

The unitary representations of the Poincaré group in any spacetime dimension

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1 Abstract

2 An extensive group-theoretical treatment of linear relativistic field equations
3 on Minkowski spacetime of arbitrary dimension $D \geq 3$ is presented. An exhaus-
4 tive treatment is performed of the two most important classes of unitary irre-
5 ducible representations of the Poincaré group, corresponding to massive and
6 massless fundamental particles. Covariant field equations are given for each
7 unitary irreducible representation of the Poincaré group with non-negative
8 mass-squared.

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1 Group-theoretical preliminaries

Elementary knowledge of the theory of Lie groups and their representations is assumed (see *e.g.* the textbooks [1, 2] or the lecture notes [3]). The basic definitions of the Lorentz and Poincaré groups together with some general facts on the theory of unitary representations are reviewed in order to fix the notation and settle down the prerequisites.

1.1 Universal covering of the Lorentz group

The group of linear homogeneous transformations $x'^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu}$ ($\mu, \nu = 0, 1, \dots, D - 1$) preserving the Minkowski metric $\eta_{\mu\nu}$ of “mostly plus” signature $(-, +, \dots, +)$,

$$\Lambda^T \eta \Lambda = \eta,$$

where Λ^T denotes the matrix transpose of Λ , is called the *Lorentz group* $O(D - 1, 1)$.

A massless particle propagates on the light-cone $x^2 = 0$. Without loss of generality, one may consider that its momentum points along the $(D - 1)$ th spatial direction. Then it turns out to be convenient to make use of the *light-cone coordinates*

$$x^{\pm} = \frac{1}{\sqrt{2}} (x^{D-1} \pm x^0) \quad \text{and} \quad x^m \quad (m = 1, \dots, D - 2),$$

where the Minkowski metric reads $\eta_{++} = 0 = \eta_{--}$, $\eta_{+-} = 1 = \eta_{-+}$ and $\eta_{mn} = \delta_{mn}$ ($m, n = 1, \dots, D - 2$).

On physical grounds, one will mainly be interested in the matrices Λ 's with determinant $+1$ and such that $\Lambda^0_0 \geq 0$. It can be shown that such matrices Λ 's also form a group that one calls the *proper orthochronous Lorentz group* denoted by $SO(D - 1, 1)^{\uparrow}$. It is connected to the identity, but not *simply connected*, that is to say, there exist loops in the group manifold $SO(D - 1, 1)^{\uparrow}$ which are not continuously contractible to a point. In order to study the representations (reps) of $SO(D - 1, 1)^{\uparrow}$, one may first determine its

62 universal covering group, *i.e.* the Lie group which is simply connected and whose Lie
 63 algebra is isomorphic to $\mathfrak{so}(D-1, 1)$, the Lie algebra of $SO(D-1, 1)^\uparrow$. For $D \geq 4$, the
 64 universal covering group, denoted $Spin(D-1, 1)$, is the double cover of $SO(D-1, 1)^\uparrow$.
 65 The spin groups $Spin(D-1, 1)$ have no intrinsically projective representations. Therefore,
 66 a single (or double) valued “representation” of $SO(D-1, 1)^\uparrow$ is meant to be a genuine
 67 representation of $Spin(D-1, 1)$.

68 **Warning:** The double cover of $SO(2, 1)^\uparrow$ is the group $SU(1, 1)$, in close analogy to the
 69 fact that the double cover of $SO(3)$ is $SU(2)$. The group $SU(2)$ is also the universal
 70 covering group of $SO(3)$, but beware that the universal cover of $SO(2, 1)^\uparrow$ is actually \mathbb{R}^3 ,
 71 covering $SO(2, 1)^\uparrow$ infinitely many times. Thus one may not speak of the spin group for
 72 the case of the proper orthochronous Lorentz group in spacetime dimension three. The
 73 analogue is that the universal cover of $SO(2) \cong U(1)$ is \mathbb{R} , that covers $U(1)$ infinitely
 74 many times, so that one may not speak of the spin group for the degenerate case of the
 75 rotation group in two spatial dimensions.

76 1.2 The Poincaré group and algebra

The transformations

$$x'^\mu = \Lambda^\mu_\nu x^\nu + a^\mu$$

where a is a spacetime translation vector, form the group of all inhomogeneous Lorentz
 transformations. If one denotes a general transformation by (Λ, a) , the multiplication law
 in the Poincaré group is given by

$$(\Lambda_2, a_2) \cdot (\Lambda_1, a_1) = (\Lambda_2 \Lambda_1, a_2 + \Lambda_2 a_1),$$

so that the *inhomogeneous Lorentz group* is the semi-direct product denoted by

$$IO(D-1, 1) = \mathbb{R}^D \rtimes O(D-1, 1).$$

77 The subgroup $ISO(D-1, 1)^\uparrow$ of inhomogeneous proper orthochronous Lorentz transfor-
 78 mations is called the *Poincaré group*. The Lie algebra $\mathfrak{iso}(D-1, 1)$ of the Poincaré group
 79 is presented by the generators $\{P_\mu, M_{\nu\rho}\}$ and by the commutation relations

$$i [M_{\mu\nu}, M_{\rho\sigma}] = \eta_{\nu\rho} M_{\mu\sigma} - \eta_{\mu\rho} M_{\nu\sigma} - \eta_{\sigma\mu} M_{\rho\nu} + \eta_{\sigma\nu} M_{\rho\mu} \quad (1)$$

$$i [P_\mu, M_{\rho\sigma}] = \eta_{\mu\rho} P_\sigma - \eta_{\mu\sigma} P_\rho, \quad (2)$$

$$i [P_\mu, P_\rho] = 0. \quad (3)$$

80 Two subalgebras must be distinguished: the Lie algebra $\mathfrak{so}(D-1, 1)$ of the Lorentz group
 81 presented by the generators $\{M_{\nu\rho}\}$ and by the commutation relations (1), and the Lie
 82 algebra \mathbb{R}^D of the Abelian translation group presented by the generators $\{P_\mu\}$ and by the
 83 commutation relations (3). The latter algebra \mathbb{R}^D is an ideal of the Poincaré algebra, as
 84 can be seen from (2). Altogether, this implies that the Lie algebra of the Poincaré group
 85 is the semi-direct sum $\mathfrak{iso}(D-1, 1) = \mathbb{R}^D \rtimes \mathfrak{so}(D-1, 1)$.

86 The Casimir elements of a Lie algebra \mathfrak{g} are homogeneous polynomials in the generators
 87 of \mathfrak{g} providing a distinguished basis of the center $\mathcal{Z}(\mathcal{U}(\mathfrak{g}))$ of the universal enveloping
 88 algebra $\mathcal{U}(\mathfrak{g})$ (see *e.g.* the part V of the lecture notes [3]). The quadratic Casimir operator
 89 of the Lorentz algebra $\mathfrak{so}(D-1, 1)$ is the square of the generators $M_{\mu\nu}$:

$$\mathcal{C}_2(\mathfrak{so}(D-1, 1)) = \frac{1}{2} M^{\mu\nu} M_{\mu\nu}. \quad (4)$$

90 The quadratic Casimir operator of the Poincaré algebra $\mathfrak{iso}(D-1, 1)$ is the square of the
91 momentum

$$\mathcal{C}_2(\mathfrak{iso}(D-1, 1)) = -P^\mu P_\mu, \quad (5)$$

92 while the quartic Casimir operator is

$$\mathcal{C}_4(\mathfrak{iso}(D-1, 1)) = -\frac{1}{2}P^2 M_{\mu\nu} M^{\mu\nu} + M_{\mu\rho} P^\rho M^{\mu\sigma} P_\sigma, \quad (6)$$

which, for $D = 4$, is the square of the Pauli-Lubanski vector W^μ ,

$$W^\mu := \frac{1}{2} \varepsilon^{\mu\nu\rho\sigma} M_{\nu\rho} P_\sigma.$$

93 1.3 ABC of unitary representations

94 The mathematical property that all non-trivial unitary reps of a non-compact simple Lie
95 group must be infinite-dimensional has some physical significance, as will be reviewed
96 later.

97 **Finite-dimensional unitary reps of non-compact simple Lie groups:** *Let $U : G \rightarrow$*
98 *$U(n)$ be a unitary representation of a Lie group G acting on a (real or complex) Hilbert*
99 *space \mathcal{H} of finite dimension $n \in \mathbb{N}$. Then U is completely reducible. Moreover, if U is*
100 *irreducible and if G is a connected simple non-compact Lie group, then U is the trivial*
101 *representation.*

102 Proof: For the property that U is completely reducible, we refer *e.g.* to the proof of
103 the proposition 5.15 in [1]. The image $U(G)$ for any unitary representation U defines a
104 subgroup of $U(n)$. Moreover, the kernel of U is a normal subgroup of the simple group
105 G . Therefore, either the representation is trivial and $\ker U = G$, or it is faithful and
106 $\ker U = \{e\}$. In the latter case, U is invertible and its image is isomorphic to its domain,
107 $U(G) \cong G$. Actually, the image $U(G)$ is a non-compact subgroup of $U(n)$ because if
108 $U(G)$ was compact, then $U^{-1}(U(G)) = G$ would be compact since U^{-1} is a continuous
109 map. But the group $U(n)$ is compact, thus it cannot contain a non-compact subgroup.
110 Therefore the representation cannot be faithful, so that it is trivial. (A different proof of
111 the second part of the theorem may be found in the section 8.1.B of [2].) \square

112 Another mathematical result which is of physical significance is the following theorem
113 on unitary irreducible representations (UIRs) of compact Lie groups.

114 **Unitary reps of compact Lie groups:** *Let U be a UIR of a compact Lie group G ,*
115 *acting on a (real or complex) Hilbert space \mathcal{H} . Then \mathcal{H} is finite-dimensional. Moreover,*
116 *every unitary representation of G is a direct sum of UIRs (the sum may be infinite).*

117 Proof: The proofs are somewhat lengthy and technical so we refer to the section 7.1 of [2]
118 for complete details. \square

119 Examples of (not so) simple groups:

- 120 • On the one hand, all (pseudo)-orthogonal groups $SO(p, q)$ are either Abelian ($p+q = 2$),
121 non-simple ($p+q = 4$) or simple ($p+q = 3$ and $p+q > 4$). Moreover, the orthogonal groups
122 ($pq = 0$) are compact, while the pseudo-orthogonal groups ($pq \neq 0$) are non-compact.
- 123 • On the other hand, the inhomogeneous Lorentz group $IO(D-1, 1)$ is non-compact
124 (both \mathbb{R}^D and $O(D-1, 1)$ are non-compact) but it is *not* semi-simple (because its normal
125 subgroup \mathbb{R}^D is Abelian).

126 2 Elementary particles as unitary irreducible representa- 127 tions of the isometry group

128 Except for the final remarks, this section is based almost *ad verbatim* on the introduction
129 of the illuminating work of Bargmann and Wigner [4], modulo some changes of notation
130 and terminology in order to follow the modern conventions.

131 The wave functions $|\psi\rangle$ describing the possible states of a quantum-mechanical system
132 form a linear vector space \mathcal{H} which, in general, is infinite-dimensional and on which a
133 positive-definite inner product $\langle\phi|\psi\rangle$ is defined for any two wave functions $|\phi\rangle$ and
134 $|\psi\rangle$ (*i.e.* they form a Hilbert space). The inner product usually involves an integration
135 over the whole configuration or momentum space and, for particles of non-vanishing spin,
136 a summation over the spin indices.

137 If the wave functions in question refer to a free particle and satisfy relativistic wave
138 equations, there exists a correspondence between the wave functions describing the same
139 state in different Lorentz frames. The transformations considered here form the group of
140 all *inhomogeneous* Lorentz transformations (including translations of the origin in space
141 and time). Let $|\psi\rangle$ and $|\psi'\rangle$ be the wave functions of the same state in two Lorentz
142 frames L and L' , respectively. Then $|\psi'\rangle = U(\Lambda, a)|\psi\rangle$, where $U(\Lambda, a)$ is a linear
143 unitary operator which depends on the transformation (Λ, a) leading from L to L' . By a
144 proper normalization, U is determined by Λ up to a factor ± 1 . Moreover, the operators U
145 form a single- or double-valued representation of the inhomogeneous Lorentz group, *i.e.*,
146 for a succession of two transformations (Λ_1, a_1) and (Λ_2, a_2) , we have

$$U(\Lambda_2, a_2)U(\Lambda_1, a_1) = \pm U(\Lambda_2\Lambda_1, a_2 + \Lambda_2 a_1). \quad (7)$$

147 Since all Lorentz frames are equivalent for the description of our system, it follows
148 that, together with $|\psi\rangle$, $U(\Lambda, a)|\psi\rangle$ is also a possible state viewed from the original
149 Lorentz frame L . Thus, the vector space \mathcal{H} contains, with every $|\psi\rangle$, all transforms
150 $U(\Lambda, a)|\psi\rangle$, where (Λ, a) is any inhomogeneous Lorentz transformation.

151 The operators U may also replace the wave equation of the system. In our discussion,
152 we use the wave functions in the “Heisenberg” representation, so that a given $|\psi\rangle$ repre-
153 sents the system for all times, and may be chosen as the “Schrödinger” wave function at
154 time $t = 0$ in a given Lorentz frame L . To find $|\psi\rangle_{t_0}$, the Schrödinger function at time
155 t_0 , one must therefore transform to a frame L' for which $t' = t - t_0$, while all other coordi-
156 nates remain unchanged. Then $|\psi\rangle_{t_0} = U(\Lambda, a)|\psi\rangle$, where (Λ, a) is the transformation
157 leading from L to L' .

158 A classification of all unitary representations of the inhomogeneous Lorentz group,
159 *i.e.* of all solution of (7), amounts, therefore, to a classification of all possible relativistic
160 wave equations. Two reps U and $\tilde{U} = VUV^{-1}$, where V is a fixed unitary operator, are
161 equivalent. If the system is described by wave functions $|\psi\rangle$, the description by

$$|\widetilde{\psi}\rangle = V|\psi\rangle \quad (8)$$

is isomorphic with respect to linear superposition, with respect to forming the inner prod-
uct of two wave functions, and also with respect to the transition from one Lorentz frame
to another. In fact, if $|\psi'\rangle = U(\Lambda, a)|\psi\rangle$, then

$$|\widetilde{\psi}'\rangle = V|\psi'\rangle = VU(\Lambda, a)V^{-1}|\widetilde{\psi}\rangle = \widetilde{U}(\Lambda, a)|\widetilde{\psi}\rangle.$$

162 Thus, one obtains classes of equivalent wave equations. Finally, it is sufficient to determine
163 the irreducible representations (irreps) since any other may be built from them.

164 Two descriptions which are equivalent according to (8) may be quite different in ap-
 165 pearance. The best known example is the description of the electromagnetic field by the
 166 field strength and the vector potential, respectively. It cannot be claimed either that
 167 equivalence in the sense of (8) implies equivalence in every physical aspect. It should be
 168 emphasized that any selection of one among the equivalent systems involves an explicit or
 169 implicit assumption as to possible interactions, *etc.* Our analysis is necessarily restricted
 170 to free particles and does not lead to any assertion about possible interactions.

171 The present discussion is not based on any hypothesis about the structure of the wave
 172 equations provided that they be covariant. In particular, it is not necessary to assume
 173 differential equations in configuration space. But it is a result of the group-theoretical
 174 analysis that every irreducible field equation is equivalent, in the sense of (8), to a system
 175 of differential equations for fields on Minkowski spacetime.

176 **Remarks:**

177 • An important theorem proved by Wigner is that any symmetry transformation that
 178 is continuously related to the identity must be represented by a linear unitary operator
 179 (see *e.g.* the appendix A of [5]). Strictly speaking, physical states are represented by
 180 *rays* in a Hilbert space. Therefore the unitary representations of the symmetry group
 181 are actually understood to be *projective* representations. In spacetime dimensions $D \geq$
 182 4, this subtlety¹ reduces to the standard distinction between single and double valued
 183 representations of the Poincaré group, as was taken for granted in the text.

184 • Notice that the previous discussion remains entirely valid if the Minkowski spacetime
 185 $\mathbb{R}^{D-1,1}$ is replaced everywhere by any other maximally symmetric spacetime (*i.e.* de Sitter
 186 spacetime dS_D , or anti de Sitter spacetime AdS_D) under the condition that the inhomogeneous
 187 Lorentz group $IO(D-1,1)$ be also replaced everywhere by the corresponding
 188 group of isometries (respectively, $O(D,1)$ or $O(D-1,2)$).

189 • In first-quantization, particles are described by fields on the spacetime and isometries
 190 are represented by unitary operators. A particle is said to be “elementary” if the rep-
 191 resentation is irreducible, and “composite” if the representation is made of a product of
 192 irreps.

193 • A modern point of view on Quantum Field Theory [5] is that a quantum field (not to
 194 be confused with the state vector discussed above) is an *operator* defined at each point
 195 of space and time, that acts in a Fock space of states, the field being represented by
 196 a superposition, for different values of the momentum, of one-particle annihilation and
 197 creation operators for particle and the associated antiparticle. The approach of [5] is
 198 to build up the quantum field by imposing Lorentz invariance at every stage. To quote
 199 Weinberg, the field equation satisfied by the *quantum* field arises almost incidentally, as a
 200 byproduct of his construction.

201 • A unitary representation of the isometry group describes the one-particle Hilbert space
 202 of states. The group-theoretical argument of Bargmann and Wigner [4] applies to the one-
 203 particle states of a free particle.² The classification of the UIRs of the Poincaré group
 204 indeed yields the Klein-Gordon equation for a massive particle, or the D’Alembert equation
 205 in the case of a massless particle [4]. This comes automatically from the group-theoretical
 206 analysis and is *not* an assumption.

207 **Summary:** On the one hand, the rules of quantum mechanics imply that quantum sym-
 208 metries correspond to unitary representations of the symmetry group carried by the Hilbert
 209 space of physical states. Furthermore, if time translations constitute a one-parameter sub-
 210 group of the symmetry group, then the Schrödinger equation for the time evolution of a

¹The case $D = 3$ is even more subtle and is treated in Appendix B.

²See *e.g.* Eq. (2.5.1) of [5] where the one-particle state vectors are denoted by $\Psi_{p,\sigma}$.

211 state vector essentially is a unitary representation of this subgroup. On the other hand,
 212 the principle of relativity dictates that the isometries of spacetime be symmetries of the
 213 physical system. All together, this implies that linear relativistic field equations may be
 214 identified with unitary reps of the isometry group.

215 3 Classification of the unitary representations

216 3.1 Induced representations

217 The method of induced reps was introduced by Wigner in his seminal paper [6] on the
 218 unitary representations of the inhomogeneous Lorentz group $IO(3,1)$ in four spacetime
 219 dimensions, which admits a straightforward generalization to any spacetime dimension D ,
 220 as reviewed now. The content of this subsection finds its origin in the section 2.5 of the
 221 comprehensive textbook [5].

From (3) one sees that all the translation generators commute with each other, so it is natural to express physical states $|\psi\rangle$ in terms of eigenvectors of the translation generators P^μ . Introducing a label σ to denote all other degrees of freedom, one thus considers states $\Psi_{q,\sigma}$ with $P_\mu \Psi_{q,\sigma} = q_\mu \Psi_{q,\sigma}$. From the infinitesimal translation $U = \mathbb{1} - iP^\mu \epsilon_\mu$ and repeated applications of it, one finds that finite translations are represented on \mathcal{H} by $U(\mathbb{1}, a) = \exp(-iP^\mu a_\mu)$, so one has

$$U(\mathbb{1}, a) \Psi_{q,\sigma} = e^{-iq \cdot a} \Psi_{q,\sigma}.$$

222 Using (2), one sees that the effect of operating on $\Psi_{p,\sigma}$ with a quantum homogeneous
 223 transformation $U(\Lambda, 0) \equiv U(\Lambda)$ is to produce an eigenvector of the translation generators
 224 with eigenvalue Λp :

$$\begin{aligned} P^\mu U(\Lambda) \Psi_{p,\sigma} &= U(\Lambda) [U^{-1}(\Lambda) P^\mu U(\Lambda)] \Psi_{p,\sigma} = U(\Lambda) ((\Lambda^{-1})_\rho{}^\mu P^\rho) \Psi_{p,\sigma} \\ &= \Lambda^\mu{}_\rho p^\rho U(\Lambda) \Psi_{p,\sigma}, \end{aligned}$$

225 since $(\Lambda^{-1})_\rho{}^\mu = \Lambda^\mu{}_\rho$. Hence $U(\Lambda) \Psi_{p,\sigma}$ must be a linear combination of the states $\Psi_{\Lambda p, \sigma'}$:

226

$$U(\Lambda) \Psi_{p,\sigma} = \sum_{\sigma'} C_{\sigma'\sigma}(\Lambda, p) \Psi_{\Lambda p, \sigma'}. \quad (9)$$

227 In general, it is possible by using suitable linear combinations of the $\Psi_{p,\sigma}$ to choose the
 228 σ labels in such a way that the matrix $C_{\sigma'\sigma}(\Lambda, p)$ is block-diagonal; in other words, so
 229 that the $\Psi_{p,\sigma}$ with σ within any one block *by themselves* furnish a representation of the
 230 Poincaré group. It is natural to identify the states of a specific particle type with the
 231 components of a representation of the Poincaré group which is irreducible, in the sense
 232 that it cannot be further decomposed in this way. It is clear from (9) that all states $\Psi_{p,\sigma}$
 233 in an irrep of the Poincaré group have momenta p^μ belonging to the orbit of a single fixed
 234 momentum, say q^μ .

One has to work out the structure of the coefficients $C_{\sigma'\sigma}(\Lambda, p)$ in irreducible representations of the Poincaré group. In order to do that, note that the only functions of p^μ that are left invariant by all transformations $\Lambda^\mu{}_\nu \in SO(D-1, 1)^\uparrow$ are, of course, $p^2 = \eta_{\mu\nu} p^\mu p^\nu$ and, for $p^2 \leq 0$, also the sign of p^0 . Hence, for each value of p^2 , and (for $p^2 \leq 0$) each sign of p^0 , one can choose a standard four-momentum, say q^μ , and express any p^μ of this class as

$$p^\mu = L^\mu{}_\nu(p) q^\nu,$$

235 where L^μ_ν is some standard proper orthochronous Lorentz transformation that depends on
 236 p^μ , and also implicitly on our choice of q^μ . One can define the states $\Psi_{p,\sigma}$ of momentum
 237 p^μ by

$$\Psi_{p,\sigma} \equiv N(p) U(L(p)) \Psi_{q,\sigma}, \quad (10)$$

238 where $N(p)$ is a numerical normalization factor. Operating on (10) with an arbitrary
 239 homogeneous Lorentz transformation $U(\Lambda)$, one now finds

$$\begin{aligned} U(\Lambda)\Psi_{p,\sigma} &= N(p) U(\Lambda L(p)) \Psi_{q,\sigma} \\ &= N(p) U(L(\Lambda p)) U(L^{-1}(\Lambda p)\Lambda L(p)) \Psi_{q,\sigma}. \end{aligned} \quad (11)$$

240 The point of this last step is that the Lorentz transformation $L^{-1}(\Lambda p)\Lambda L(p)$ takes q to
 241 $L(p)q = p$, then to Λp , and finally back to q , so it belongs to the subgroup of the Lorentz
 242 group consisting of Lorentz transformations W^μ_ν that leave q^μ invariant: $W^\mu_\nu q^\nu = q^\mu$.
 243 This stability subgroup is called the *little group* corresponding to q . For any W, \bar{W} in the
 244 little group, one has

$$U(W)\Psi_{q,\sigma} = \sum_{\sigma'} D_{\sigma'\sigma}^q(W) \Psi_{q,\sigma'} \quad (12)$$

and

$$D_{\sigma'\sigma}^q(\bar{W}W) = \sum_{\sigma''} D_{\sigma'\sigma''}^q(\bar{W}) D_{\sigma''\sigma}^q(W),$$

that is to say, the coefficients $D^q(W)$ furnish a representation of the little group. In
 particular, for $W(\Lambda, p) \equiv L^{-1}(\Lambda p)\Lambda L(p)$, the equation (11) becomes

$$U(\Lambda)\Psi_{p,\sigma} = N(p) \sum_{\sigma'} D_{\sigma'\sigma}(W(\Lambda, p)) U(L(\Lambda p)) \Psi_{q,\sigma'}$$

245 or, recalling the definition (10),

$$U(\Lambda)\Psi_{p,\sigma} = \frac{N(p)}{N(\Lambda p)} \sum_{\sigma'} D_{\sigma'\sigma}(W(\Lambda, p)) \Psi_{\Lambda p, \sigma'}. \quad (13)$$

246 Apart from the question of normalization, the problem of determining the coefficients $C_{\sigma'\sigma}$
 247 in the transformation rule (9) has been reduced to the problem of determining the coeffi-
 248 cients $D_{\sigma'\sigma}$. In other words, the problem of determining all possible irreps of the Poincaré
 249 group has been reduced to the problem of finding all possible irreps of the little group,
 250 depending on the class of momentum to which q^μ belongs. This approach, of deriving
 251 representations of a semi-direct product like the inhomogeneous Lorentz group from the
 252 representations of the stability subgroup, is called the *method of induced representations*.

253 The wave function $\Psi_{p,\sigma}$ depends on the momentum, therefore its Fourier transform
 254 $\Psi_{x,\sigma}$ depends on the spacetime coordinate, so that the wave function is called a (complex)
 255 *field* on Minkowski spacetime $\mathbb{R}^{D-1,1}$ and the quantities $\Psi_{x,\sigma}$ at fixed x and for varying σ
 256 are referred to as its *physical components* at x .

257 3.2 Orbits and stability subgroups

258 The orbit of a given non-vanishing vector q^μ of Minkowski spacetime $\mathbb{R}^{D-1,1}$ under Lorentz
 259 transformations is, by definition, the hypersurface of constant momentum square p^2 . Ge-
 260 ometrically speaking, it is a quadric of curvature radius $m > 0$. More precisely, the
 261 hypersurface

- 262 • $p^2 = -m^2$ is a two-sheeted hyperboloid, each sheet of which is called a *mass-shell*.
 263 The corresponding UIR is said to be *massive*.
- 264 • $p^2 = 0$ is a cone, each half of which is called a *light-cone*. The corresponding UIR is
 265 said to be *massless* ($m = 0$).
- 266 • $p^2 = +m^2$ is a one-sheeted hyperboloid. The corresponding UIR is said to be
 267 *tachyonic*.

268 Orthochronous Lorentz transformations preserve the sign of the time component of vectors
 269 of non-positive momentum-squared, thus the orbit of a time-like (light-like) vector is the
 270 mass-shell (respectively, light-cone) to which the corresponding vector belongs.

271 **Remarks:**

- 272 • Notice that the Hilbert space carrying the irrep is indeed an eigenspace of the quadratic
 273 Casimir operator (5), the eigenvalue of which is $\mathcal{C}_2 = \pm m^2$ (the eigenvalue is real because
 274 the representation is unitary), as it should according to Schur’s lemma. Moreover, the
 275 quadratic Casimir classifies the UIRs but does not entirely characterize them.
- 276 • As quoted in Section 2, it is not necessary to assume differential equations in position
 277 space, because the group-theoretical analysis directly leads to a wave function which is
 278 a function of the momenta on the orbit, the Fourier transform of which is a function
 279 in position space obeying the Klein–Gordon equation $\square \Psi_{x,\sigma} = \pm m^2 \Psi_{x,\sigma}$. By a slight
 280 abuse of terminology, states or fields that satisfy their relativistic equations of motion are
 281 called “on-(mass-)shell” in physics literature, while those for which those equations have
 282 not been imposed) are called “off-shell”.

283 By going to a rest-frame, it is easy to show that the stabilizer of a time-like vector
 284 $q^\mu = (m, \vec{0}) \neq 0$ is the rotation subgroup $SO(D-1) \subset SO(D-1,1)^\uparrow$. For a space-
 285 like vector, one may choose a frame such that the non-vanishing momentum is along the
 286 $(D-1)$ th spatial axis: $q^\mu = (0, 0, \dots, 0, m) \neq 0$. Thus its stabilizer is the subgroup
 287 $SO(D-2,1)^\uparrow \subset SO(D-1,1)^\uparrow$. In the case of a light-like vector, the little group “*is not*
 288 *quite so obvious*” to determine, as was stressed by Wigner himself [7]. It clearly contains
 289 the rotation group $SO(D-2)$ in the space-like hyperplane \mathbb{R}^{D-2} transverse to the light-
 290 ray along the momentum. Now, we will provide an algebraic proof that the stabilizer of
 291 a light-like vector is the Euclidean group $ISO(D-2)$. According to Wigner, reviewing
 292 his $D = 4$ analysis, “*no simple argument is known (...) to show directly that the group*
 293 *of Lorentz transformations which leave a null vector invariant is isomorphic to the two-*
 294 *dimensional Euclidean group, desirable as it would be to have such an argument. Clearly,*
 295 *there is no plane in the four-space of momenta in which these transformations could be*
 296 *interpreted directly as displacements (...) because all transformations considered here are*
 297 *homogeneous*” [7]. Even though there is no simple geometric way to understand this fact,
 298 the algebraic proof reviewed here is rather straightforward.

299 Proof: By going in a light-cone frame (see Section 1.1), it is possible to express the com-
 300 ponents of a momentum p^μ obeying $p^2 = 0$ as $p_\mu = (p_-, 0, 0, \dots, 0)$. In words, one can set
 301 the component p_+ to zero, as well as all the transverse components p_m ($m = 1, \dots, D-2$).
 302 The condition that the component p_- be unaffected by a Lorentz transformation trans-
 303 lates as $0 \stackrel{!}{=} i[p_-, M_{\nu\rho}] = \eta_{-\nu} p_\rho - \eta_{-\rho} p_\nu$ due to (2). Obviously, the transformation
 304 generated by M_{+-} does modify p_- , hence it cannot be part of the little group for p .
 305 The other Lorentz generators preserve p_- , but they should also satisfy the equations
 306 $[p_m, M_{\mu\nu}] = 0 = [p_+, M_{\mu\nu}]$. It is readily seen that $i[p_m, M_{n-}] = \delta_{mn} p_- \neq 0$ (for $m = n$),
 307 therefore M_{n-} does not belong to the little group of p_μ either. We are left with the gen-
 308 erators $\{M_{mn}, M_{+n}\}$ which preserve the (vanishing) value of p_+ . It turns out to be more

309 convenient for later purpose to work with the generators $\pi_n := p_- M_{+n} = p^\mu M_{\mu n}$ instead.
 310 This redefinition does not modify the algebra since p_- commutes with all the generators
 311 of the little group. From the Poincaré algebra (1)–(3) one finds, in the light-cone frame,

$$i [M_{mn}, M_{pq}] = \delta_{np} M_{mq} - \delta_{mp} M_{nq} - \delta_{qm} M_{pn} + \delta_{qn} M_{pm}, \quad (14)$$

$$i [\pi_m, M_{np}] = \delta_{mn} \pi_p - \delta_{mp} \pi_n, \quad (15)$$

$$i [\pi_m, \pi_n] = 0. \quad (16)$$

312 As can be seen, the generators $\{M_{mn}, \pi_m\}$ span the Lie algebra of the inhomogeneous
 313 orthogonal group $ISO(D-2)$. \square

314 For later purpose, notice that the quadratic Casimir operator of the Euclidean algebra
 315 $\mathfrak{iso}(D-2)$ presented by the generators $\{M_{mn}, \pi_m\}$ and the relations (14)–(16) is the square
 316 of the “translation” generators

$$\mathcal{C}_2(\mathfrak{iso}(D-2)) = \pi^m \pi_m. \quad (17)$$

317 To end up this discussion, one should consider the case of a vanishing momentum.
 318 Of course, the orbit of a vanishing vector under linear transformations is itself while its
 319 stabilizer is the whole linear subgroup. Therefore, the subgroup of $SO(D-1, 1)^\uparrow$ leaving
 320 invariant the zero-momentum vector $p^\mu = 0$ is the whole group itself. This ends up the
 321 determination of the orbit and stabilizer of any possible vector $\in \mathbb{R}^{D-1, 1}$.

322 **Remark:** The zero-momentum ($q^\mu = 0$) representations are essentially UIRs of the little
 323 group $SO(D-1, 1)^\uparrow$ because the translation group acts trivially. The proper orthochronous
 324 Lorentz group may be identified with the isometry group of the de Sitter spacetime dS_{D-1} .
 325 In other words, the wave function of the zero-momentum representation may be interpreted
 326 as a wave function on a lower-dimensional de Sitter spacetime, and conversely. Even
 327 though their physical meaning may differ, both UIRs may be mathematically identified.

3.3 Classification

328
 329 To summarize the previous subsection, the UIRs of the Poincaré group $ISO(D-1, 1)^\uparrow$
 330 have been divided into four classes according to the four possible orbits of the momentum,
 331 summarized in the following table (where $m^2 > 0$):

Gender	Orbit	Stability subgroup	UIR
$p^2 = -m^2$	Mass-shell	$SO(D-1)$	Massive
$p^2 = 0$	Light-cone	$ISO(D-2)$	Massless
$p^2 = +m^2$	Hyperboloid	$SO(D-2, 1)^\uparrow$	Tachyonic
$p_\mu = 0$	Origin	$SO(D-1, 1)^\uparrow$	Zero-momentum

332
 333 The problem of classifying the UIRs of the Poincaré group $ISO(D-1, 1)^\uparrow$ has been reduced
 334 to the classification of the UIRs of the stability subgroup of the momentum, which are either
 335 a unimodular orthogonal group, an Euclidean group or a proper orthochronous Lorentz
 336 group.

337 Actually, the method of induced representation may also be applied to the classification
 338 of the UIRs of the Euclidean group $ISO(D-2)$, the little group of a massless particle.
 339 The important thing to understand is that the light-like momentum p^μ is fixed and that
 340 what should be examined is the action of the little group on the physical components
 341 characterized by σ . From (16) one sees that the $D-2$ “translation” generators π^i commute
 342 with each other, so it is natural to express physical states $\Psi_{p, \sigma}$ in terms of eigenvectors ξ^m
 343 of these generators π^m . Introducing a label ς to denote all remaining physical components,

344 one thus considers states $\Psi_{p,\xi,\varsigma}$ with $\pi_m \Psi_{p,\xi,\varsigma} = \xi_m \Psi_{p,\xi,\varsigma}$. The discussion presented in
 345 Subsection 3.1 may be repeated almost identically, up to the replacement of the momentum
 346 p by the eigenvector ξ , the label σ by ς , the Poincaré group $ISO(D-1,1)^\dagger$ by the
 347 Euclidean group $ISO(D-2)$ and the proper orthochronous Lorentz group $SO(D-1,1)^\dagger$
 348 by the unimodular orthogonal group $SO(D-2)$. The conclusion is therefore similar: the
 349 problem of determining all possible irreps of the massless little group $ISO(D-2)$ has
 350 been reduced to the problem of finding all possible irreps of the stability subgroup of the
 351 $(D-2)$ -vector ξ , called the *short little group* in the literature [8].

352 The massless representations induced by a non-trivial representation of the little group
 353 may therefore be divided into distinct categories, depending on the class of momentum
 354 to which ξ^m belongs. The situation is simpler here because there exist only two possible
 355 classes of orbits for a vector in the Euclidean space \mathbb{R}^{D-2} : either the origin $\xi^m = 0$, or a
 356 $(D-3)$ -sphere of radius $\mu > 0$. In the first case, the action of the elusive “translation” op-
 357 erators π^m is trivial and, effectively, the little group is identified with the short little group
 358 $SO(D-2)$. These representations are most often referred to as *helicity* representations by
 359 analogy with the $D = 4$ case. In the second case, the corresponding representations are
 360 most often referred to as *continuous spin* representations [8], even though Wigner also used
 361 the name *infinite spin* [7]. The former name originates from the fact that the transverse
 362 vector ξ^m has a continuous range of values. Nevertheless, the latter name is more adequate
 363 in some respect, as will be argued later on. Roughly speaking the point is that, on the
 364 orbit $\xi^2 = \mu^2$, the components spanned by the internal vector ξ^m take values on the sphere
 365 $S^{D-3} \subset \mathbb{R}^{D-2}$ of radius $\mu = |\xi|$. The “radius” μ of this internal sphere has actually the
 366 dimension of a mass parameter (the reason is that the sphere S^{D-3} is somehow in internal
 367 “momentum” space). Indeed, for massless representations, the parameter μ classifying the
 368 various irreps should be understood as the analogue of the mass for massive irreps, while
 369 the angular coordinates on the sphere S^{D-3} are the genuine “spin” degrees of freedom,
 370 the Fourier conjugates of which are discrete variables as is more usual for spin degrees of
 371 freedom. This point of view motivates the name “infinite spin.”³

372 To summarize, the UIRs of the Euclidean group $ISO(D-2)$ are divided into two
 373 classes according to the two possible orbits of the $(D-2)$ -vector ξ_m , summarized in the
 374 following table:

Gender	Orbit	Stability subgroup	Massless UIR
$\xi^2 = \mu^2$	Sphere	$SO(D-3)$	Infinite spin
$\xi_m = 0$	Origin	$SO(D-2)$	Helicity

375
 376 As a consequence of the method of induced representations, the physical components
 377 of a first-quantized elementary particle span a UIR of the little group. The number of
 378 local degrees of freedom (or of physical components) of the field theory is thus given by
 379 the dimension of the Hilbert space carrying the UIR of the little group. In the light of
 380 the standard results of representation theory (reviewed in Subsection 1.3) and using the
 381 method of induced representation, the UIRs of the Poincaré group may alternatively be
 382 divided into two distinct classes: the *finite-component* ones (the massive and the helicity
 383 reps) for which the (short) little group is compact, and the *infinite-component* ones (the
 384 infinite-spin, the tachyonic and the zero-momentum reps) for which the little group is
 385 non-compact.

386 **Remarks:**

³Actually, in Subsection 5.3 an explicit derivation of the continuous spin representation from a proper “infinite spin” limit of a massive representation is reviewed. All the former comments find their natural interpretation in this point of view.

387 • More precisely, the lower-dimensional cases $D = 2, 3$ are degenerate in the following
 388 sense (when $p^\mu \neq 0$). In $D = 2$, all little groups are trivial, thus all physical fields are
 389 scalars. In $D = 3$, all little groups are Abelian (massive: $SO(2)$, massless: \mathbb{R} , tachyonic:
 390 $SO(1,1)^\uparrow \cong \mathbb{R}$) hence all their UIRs have (complex) dimension one: generically, fields
 391 have one physical degrees of freedom. Notice that the helicity reps may be assigned a
 392 “conformal spin” if they are extended to irreps of the group $SO(D,2) \supset SO(D-1,1)^\uparrow$.
 393 Notice also that the “spin” of a massive representation is not discretized in $D = 3$ but
 394 can be an arbitrary real number⁴ [10] because the universal cover of $SO(2,1)^\uparrow$ covers it
 395 infinitely often.

396 • For massive and helicity representations, the number of local physical degrees of freedom
 397 may be determined from the well known formulas for the dimension of any UIR of the
 398 orthogonal groups (reviewed in Subsection 4.3 for the tensorial irreps).

399 • This group-theoretical analysis does not probe topological theories (such as Chern-
 400 Simons theory) because such theories correspond to identically vanishing representations
 401 of the little group since they have no *local* physical degrees of freedom.

402 The following corollary provides a group-theoretical explanation of the fact that com-
 403 bining the principle of relativity with the rules of quantum mechanics necessarily leads to
 404 *field* theory.

405 **Corollary:** *Every non-trivial unitary irreducible representation of the isometry group of*
 406 *any maximally-symmetric spacetime is infinite-dimensional.*

407 Proof: The Hilbert space carrying a non-trivial unitary representation of the Poincaré
 408 group is infinite-dimensional because (i) in the generic case, $q_\mu \neq 0$, the orbit is either
 409 a hyperboloid ($p^2 \neq 0$) or a cone ($p^2 = 0$) and the space of wave functions on the orbit
 410 is infinite-dimensional, (ii) the zero-momentum representations of the Poincaré group are
 411 unitary representations of the de Sitter isometry group. Thus, the proof is ended by
 412 noticing that all non-trivial unitary representations of the isometry group of (anti) de
 413 Sitter spacetimes (A) dS_D also are infinite-dimensional, because their isometry groups are
 414 *pseudo-orthogonal* Lie groups. \square

415 4 Tensorial representations and Young diagrams

416 Most of the material reviewed here may be found in textbooks such as [11]. Nevertheless,
 417 large parts of this section are either copied or adapted from the reference [12] because
 418 altogether it provides an excellent summary, both for its pedagogical and comprehensive
 419 values. The material collected in the present section goes slightly beyond what is strictly
 420 necessary for these lectures, but the reader may find it useful in specific applications.

421 4.1 Symmetric group

422 An (unlabeled) *Young diagram*, consisting of n boxes arranged in r (left justified) rows,
 423 represents a *partition* of the integer n into r *parts*:

$$n = \sum_{a=1}^r \lambda_a \quad (\lambda_1 \geq \lambda_2 \geq \dots \geq \lambda_r).$$

424 That is, λ_a is the number of boxes in the a th row. Successive row lengths are non-
 425 increasing from top to bottom. A simpler notation for the partition is the list of its parts:

⁴This peculiarity is related to the existence of anyons in three spacetime dimensions, cf. Appendix B.

426 $\lambda = \{\lambda_1, \lambda_2, \dots, \lambda_r\}$. For instance,

$$\{3, 3, 1\} = \begin{array}{|c|c|c|} \hline \square & \square & \square \\ \hline \square & \square & \square \\ \hline \square & & \\ \hline \end{array} .$$

427 **Examples:** There are five partitions of 4:

$$\{4\}, \{3, 1\}, \{2, 2\}, \{2, 1, 1\}, \{1, 1, 1, 1\} . \tag{18}$$

428 Partitions play a key role in the study of the symmetric group \mathfrak{S}_n . This is the group
 429 of all permutations of n objects. It has $n!$ elements and *its inequivalent irreducible rep-*
 430 *resentations may be labeled by the partitions of n .* [In the following, Greek letters λ, μ
 431 and ν will be used to specify not only partitions and Young diagrams but also irreducible
 432 representations of the symmetric group and other groups.]

433 The connection between the symmetric group and tensors was initially developed by
 434 H. Weyl. This connection can be approached in (at least) two equivalent ways.

435 **A.** Let $T_{\mu_1 \dots \mu_n}$ be a ‘generic’ n -index tensor, without any special symmetry property.
 436 [For the moment, ‘tensor’ just means a function of n indices, not necessarily with
 437 any geometrical realization. It must be meaningful, however, to *add* — and form
 438 linear combinations of — tensors of the same rank.] A *Young tableau*, or labeled
 439 Young diagram, is an assignment of the numbers $1, 2, \dots, n$ to the n boxes of a
 440 Young diagram λ . The tableau is *standard* if the numbers are increasing both
 441 along rows from left to right and down columns from top to bottom. The entries
 442 $1, \dots, n$ in the tableau indicate the n successive indices of $T_{\mu_1 \dots \mu_n}$. The tableau
 443 defines a certain symmetrization operation on these indices: *symmetrize* on the set
 444 of indices indicated by the entries in each row, then *antisymmetrize* the result on
 445 the set of indices indicated by the entries in each column. The resulting object
 446 is a tensor, \tilde{T} , with certain index symmetries. Now let each permutation of \mathfrak{S}_n
 447 act (separately) upon \tilde{T} . The $n!$ results are not linearly independent; they span
 448 a vector space $V_\lambda^{\mathfrak{S}_n}$ which supports an irreducible representation of \mathfrak{S}_n . Different
 449 tableaux corresponding to the same diagram λ yield equivalent (by not identical)
 450 representations.

451 **Example:** The partition $\{2, 2\}$ of 4 has two standard tableaux:

$$\begin{array}{|c|c|} \hline 1 & 2 \\ \hline 3 & 4 \\ \hline \end{array} \quad \text{and} \quad \begin{array}{|c|c|} \hline 1 & 3 \\ \hline 2 & 4 \\ \hline \end{array} . \tag{19}$$

452 Let us construct the symmetrized tensor $\tilde{T}_{abcd} := R_{ab|cd}$ corresponding to the second
 453 of these:

$$\begin{array}{|c|c|} \hline a & c \\ \hline b & d \\ \hline \end{array} . \tag{20}$$

First symmetrize over the first and third indices (a and c), and over the second and fourth (b and d):

$$\frac{1}{4} (T_{abcd} + T_{cbad} + T_{adcb} + T_{cdab}) .$$

454 Now antisymmetrize the result over the first and second indices (a and b), and over

455 the third and fourth (c and d),⁵ dropping the combinatorial factor $\frac{1}{16}$, we get

$$\begin{aligned}
 R_{ab|cd} &= T_{abcd} + T_{cbad} + T_{adcb} + T_{cdab} - T_{bacd} - T_{cabd} - T_{bdca} - T_{cdba} \\
 &\quad - T_{abdc} - T_{dbac} - T_{acdb} - T_{dcab} + T_{badc} + T_{dabc} + T_{bcda} + T_{dcba}.
 \end{aligned}$$

456 It is easy to check that R possesses the symmetries of the Riemann tensor. There
 457 are two independent orders of its indices (e.g. $R_{ab|cd}$ and $R_{ac|bd}$), and applying any
 458 permutation to the indices produces some linear combination of those two basic
 459 objects. On the other hand, performing on T the operations prescribed by the first
 460 tableau in (19) produces a different expression $P_{ac|bd}$, which, however, generates a
 461 two-dimensional representation of \mathfrak{S}_4 with the same abstract index structure as that
 462 generated by R . A non-standard tableau would also yield such a representation, but
 463 the tensors within it would be linear combinations of those already found. One finds

$$\begin{aligned}
 P_{ac|bd} &= T_{abcd} + T_{bacd} + T_{abdc} + T_{badc} - T_{cbad} - T_{bcad} - T_{cbda} - T_{bcda} \\
 &\quad - T_{adcb} - T_{dacb} - T_{adbc} - T_{dabc} + T_{cdab} + T_{dcab} + T_{cdba} + T_{dcba}.
 \end{aligned}$$

464 As the reader may check, no linear combinations of P can reproduce R . The
 465 objects $P_{ab|cd}$, $P_{ac|bd}$, $R_{ab|cd}$ and $R_{ac|bd}$ are linearly independent. Although R and
 466 P are characterized by the same Young *diagram*, they are associated with different
 467 standard Young *tableaux* and therefore span two *different* (although equivalent)
 468 irreducible representations of \mathfrak{S}_n . Two representations may indeed be equivalent
 469 without being identical. This happens in particular for the irreducible decomposition
 470 of the regular representation of \mathfrak{S}_n where every irreducible representation appears
 471 with a multiplicity equal to its dimension. When the dimension of an \mathfrak{S}_n irreducible
 472 representation is $d > 1$, then d copies of that irreducible representation appear in
 473 the decomposition of the regular representation of \mathfrak{S}_n and all these d representations
 474 are equivalent, although not identical.

475 **Example:** Define a *symmetrized Riemann tensor* (the *Jacobi tensor*) by $J_{ad;bc} :=$
 476 $\frac{1}{2}(R_{ab|cd} + R_{ac|bd})$. It obeys $J_{ab;cd} = J_{ba;cd} = J_{ab;dc}$. Then it is easy to show that
 477 $R_{ab|cd} = \frac{2}{3}(J_{ad;bc} - J_{bd;ac})$. Thus the tensor J has no fewer independent components
 478 and contains no less information than the tensor R , despite the extra symmetrization;
 479 R is recovered from J by an antisymmetrization. The tensors R and J are really
 480 the same tensor expressed with respect to different bases.

481 **B.** The *regular representation* of \mathfrak{S}_n is the $n!$ -dimensional representation obtained by
 482 letting \mathfrak{S}_n act by left multiplication on the formal linear combinations of elements
 483 of \mathfrak{S}_n . [That is, one labels the basis vectors of $\mathbb{R}^{n!}$ by elements of \mathfrak{S}_n , defines that
 484 action of each permutation on the basis vectors in the natural way, and extends
 485 this definition to the whole space by linearity.] Equivalently, the vector space of
 486 the regular representation is the space of real-valued functions defined on \mathfrak{S}_n . [In
 487 general the regular representation is defined with complex scalars, but for \mathfrak{S}_n it is
 488 sufficient to work with real coefficients.]

489 **Regular representation:** *The regular representation contains every irreducible*
 490 *representation with a multiplicity equal to its dimension. Each Young diagram λ*
 491 *corresponds to an irreducible representation of \mathfrak{S}_n . Its dimension and multiplicity*
 492 *are equal to the number of standard tableaux of diagram λ .*

⁵Here we adopt the convention that the second round of permutations interchanges indices with the same *names*, rather than indices in the same *positions* in the various terms. The opposite convention is tantamount to antisymmetrizing *first*, which leads to a different, but mathematically isomorphic, development of the representation theory. The issue here is analogous to the distinction between space-fixed and body-fixed axes in the study of the rotation group (active or passive transformations).

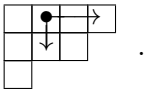
493 The symmetrization procedure described under **A.** can be transcribed to the more
 494 abstract context **B.** to construct a projection operator onto the subspace of $\mathbb{R}^{n!}$ supporting
 495 each representation. [The numerical coefficient needed to normalize the tableau operation
 496 as a projection — an operator whose square is itself — is not usually the same as that
 497 accumulated from the individual symmetrization operations. For example, to make R_{abcd}
 498 into a projection of T_{abcd} , one needs to divide by 12, not 16.]

Example: In (18), the partition $\{4\}$ corresponds to the totally symmetric four-index tensors, a one-dimensional space $V_{\{4\}}^{\mathfrak{S}_4}$. Similarly, $\{1, 1, 1, 1\}$ yields the totally antisymmetric tensors. A generic rank-four tensor, T_{abcd} , can be decomposed into the sum of its symmetric and antisymmetric parts, plus a remainder. The theory we are expounding here tells how to decompose the remainder further. The partition $\{2, 2\}$ yields two independent two-dimensional subrepresentations of the regular representation; in more concrete terms, there are two independent pieces of T_{abcd} ($\frac{1}{12} R_{ab|cd}$ and $\frac{1}{12} P_{ac|bd}$) constructed as described in connection with (19). One of these ($R_{ab|cd}$) has exactly the symmetries of the Riemann tensor; the other ($P_{ac|bd}$, coming from the first tableau of (19)) has the same abstract symmetry as the Riemann tensor, but with the indices ordered differently. Finally, each of the remaining partitions in (18), i.e., $\{3, 1\}$ and $\{2, 1, 1\}$, can be made into a standard tableau in three different ways. Therefore, each of these two representations has three separate pieces of T corresponding to it, and each piece is three-dimensional (has three independent index orders after its symmetries are taken into account). Thus the total number of independent tensors which can be formed from the irreducible parts of T_{abcd} by index permutations is

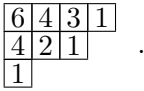
$$1^2 + 1^2 + 2^2 + 3^2 + 3^2 = 24 = 4!$$

499 which is simply the total number of permutations of the indices of T itself, as it must be.

500 To state a formula for the dimension of an irreducible representation $V_{\lambda}^{\mathfrak{S}_n}$ of \mathfrak{S}_n , we
 501 need the concept of the hook length of a given box in a Young diagram λ . The *hook*
 502 *length* of a box in a Young diagram is the number of squares directly below or directly to
 503 the right of the box, including the box once:



505 **Example:** In the following diagram, each box is labeled by its hook length:



507 One then has the following *hook length formula* for the dimension of the representation
 508 $V_{\lambda}^{\mathfrak{S}_n}$ of \mathfrak{S}_n corresponding to the Young diagram λ :

$$\dim V_{\lambda}^{\mathfrak{S}_n} = \frac{n!}{\prod(\text{hook lengths})}. \tag{21}$$

509 **Remark:** Note carefully that the “dimension” we have been discussing up to now is the
 510 number of independent *index orders* of a tensor, not the number of independent *compo-*
 511 *nents* when the tensor is realized geometrically with respect to a particular underlying
 512 vector space or manifold. The latter number depends on the dimension (say D) of that
 513 underlying space, while the former is independent of D (so long as D is sufficiently large,
 514 as we tacitly assume in generic discussions). For example, the number of components of

515 an antisymmetric two-index tensor is $\frac{D(D-1)}{2}$, but the number of its index orders is always
 516 1, except in dimension $D = 1$ where no non-zero antisymmetric tensors exist at all.

517 4.2 General linear group

518 We now turn to the representation theory of the general linear and orthogonal groups,
 519 where the (spacetime) dimension D plays a key role. The theory of partitions and of the
 520 representations of the permutation groups is the foundation on which this topic is built.

521 Let $\{v_a\}$ represent a generic element of \mathbb{R}^{D*} (or of the cotangent space at a point of a
 522 D -dimensional manifold). The action of non-singular linear operators on this space gives
 523 a D -dimensional irreducible representation $V \cong \mathbb{R}^{D*}$ of the general linear group $GL(D)$;
 524 indeed, this representation defines the group itself. The rank-two tensors, $\{T_{ab}\}$, carry a
 525 larger representation of $GL(D)$ ($V \otimes V$, of dimension D^2), where the group elements act on
 526 the two indices simultaneously. The latter representation is reducible: it decomposes into
 527 the subspace of symmetric and antisymmetric rank-two tensors $V \otimes V \cong (V \odot V) \oplus (V \wedge V)$,
 528 of respective dimensions $\frac{D(D+1)}{2}$ and $\frac{D(D-1)}{2}$. Similarly, the tensor representation of rank
 529 n , $V^{\otimes n}$, decomposes into irreducible representations of $GL(D)$ which are associated with
 530 the irreducible representations of \mathfrak{S}_n acting on the indices, which in turn are labeled by
 531 the partitions of n , hence by Young diagrams. Young diagrams with more than D rows
 532 do not contribute [if λ is a partition of n into more than D parts, then the associated
 533 index symmetrization of a D -dimensional rank- n tensor yields an expression that vanishes
 534 identically; in particular, there are no non-zero totally antisymmetric rank- n tensors if
 535 $n > D$].

536 More precisely, let λ be a Young *tableau*. The *Schur module* $V_\lambda^{GL(D)}$ is the vector space
 537 of all rank- n tensors \tilde{T} in $V^{\otimes n}$ such that:

- 538 (i) the tensor \tilde{T} is completely antisymmetric in the entries of each column of
- 539 λ ,
- 540 (ii) complete antisymmetrization of \tilde{T} in the entries of a column of λ and
- 541 another entry of λ that is on the right-hand side of the column vanishes.

542 This construction is equivalent to the construction **A**.

543 **Example:** Associated with the Young tableau (20), the tensor $R_{ab|cd}$ introduced in the
 544 subsection 4.1 obeys to the conditions (i) and (ii): $R_{ab|cd} = -R_{ba|cd} = -R_{ab|dc}$ and
 545 $R_{ab|cd} + R_{bc|ad} + R_{ca|bd} = 0$.

546 As explained in the footnote 5, if one interchanges everywhere in the previous con-
 547 structions the words “symmetric” and “antisymmetric,” then the (reducible) representa-
 548 tion spaces characterized by the same Young *diagram* [but not by the same Young *tableau*]
 549 are isomorphic and the conditions (i)-(ii) must be replaced with:

- 550 (a) the tensor is completely (or totally) symmetric in the entries of each column
- 551 of λ ,
- 552 (b) complete symmetrization of the tensor in the entries of a row of λ and
- 553 another entry of λ that sits in a lower row vanishes.

554 **Example:** Taking the standard Young tableau (20) and constructing, following the “man-
 555 ifestly symmetric convention”, the irreducible tensor associated with it, one obtains a ten-
 556 sor \mathcal{R} with the same abstract index symmetries as J [*i.e.* obeying the constraints (a)
 557 and (b)] but which is however linearly independent from J , thence linearly independent

558 from R alone. The tensor \mathcal{R} can be expressed as a linear combination of *both* R and P .
 559 Similarly, taking the first standard Young tableau in (19) and following the manifestly
 560 symmetric convention, one obtains a tensor \mathcal{P} obeying (a) and (b). This tensor is linearly
 561 independent from P alone as it is a linear combination of *both* P and R . Summarizing,
 562 associated with the Young *diagram* $\{2, 2\}$ we have the (reducible) representation space
 563 spanned by either $\{R, P\}$ in the manifestly antisymmetric convention or by $\{\mathcal{R}, \mathcal{P}\}$ in the
 564 manifestly symmetric convention.

565 **Remarks:**

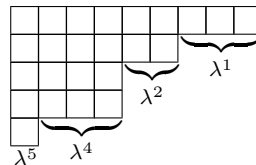
566 • An important point to note is that, by the previous construction featuring irreducible tensors
 567 with definite symmetry properties, one generates essentially *all* the finite-dimensional
 568 irreducible representations of $GL(D, \mathbb{R})$. To be more precise, $GL(D, \mathbb{R})$ tensors can be
 569 of type (p, q) , i.e., having p contravariant indices and q covariant ones. The exhaustive
 570 list of finite-dimensional irreducible representations of $GL(D, \mathbb{R})$ is provided by (p, q) -type
 571 tensors characterised by a pair of Young tableaux of rank p and q , respectively, and such
 572 that the contraction of any covariant index with a contravariant one gives zero identically.
 573 See e.g. Chapter 13 of [9] for more details.

574 • In order to make contact with an alternative road to the representation theory of $GL(D)$,
 575 one says that the irreducible representation $\Gamma_{\lambda^1 \dots \lambda^{D-1}}$ of $\mathfrak{sl}(D, \mathbb{C}) \equiv A_{D-1}$ with highest
 576 weight $\Lambda = \lambda^1 \Lambda_{(1)} + \lambda^2 \Lambda_{(2)} + \dots + \lambda^{D-1} \Lambda_{(D-1)}$ [see e.g. the Part II of the lecture notes
 577 [3] for definitions and notations] is obtained by applying the Schur functor \mathbb{S}_λ [i.e. the
 578 construction presented above] to the standard representation V , where the Young diagram
 579 is

$$\lambda = \{\lambda^1 + \dots + \lambda^{D-1}, \lambda^2 + \dots + \lambda^{D-1}, \dots, \lambda^{D-1}, 0\}.$$

580 In terms of the Young diagram for λ , the Dynkin labels λ^a ($1 \leq a \leq D - 1$) are the
 581 differences of lengths of rows: $\lambda^a = \lambda_a - \lambda_{a+1}$.

582 **Example:** If $D = 6$, then



584 is the Young diagram corresponding to the irrep $\Gamma_{3,2,0,3,1}$ of $A_5 \equiv \mathfrak{sl}(6, \mathbb{C})$.

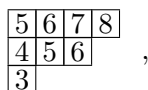
585 The dimension of the representation $V_\lambda^{GL(D)}$ of $GL(D)$ corresponding to the Young
 586 diagram λ is:

$$\dim V_\lambda^{GL(D)} = \prod \frac{D - \text{row} + \text{column}}{\text{hook length}}, \tag{22}$$

587 where the product is over the n boxes while “row” and “column” respectively give the
 588 place of the corresponding box. As was underlined before, the formula (22) is distinct
 589 from the hook length formula (21).

590 **Examples:**

591 • In the following diagram



592

593 each box is labeled by its value in the numerator of (22) for $D = 5$. Observe that, for the
 594 corresponding diagram λ , $\dim V_\lambda^{GL(5)} = 1050 \neq 70 = \dim V_\lambda^{\mathfrak{S}_5}$.

595 • The space of (anti)symmetric tensors of V of rank n are denoted by $\odot^n(V)$ (respectively,
 596 $\wedge^n(V)$). It carries an irreducible representation of $GL(D)$ labeled by a Young diagram
 597 made of one row (respectively, column) of length n . The dimensions

$$\dim \odot^n(V) = \binom{D+n-1}{n}, \quad \dim \wedge^n(V) = \binom{D}{n}, \quad (23)$$

598 are easily computed from the formula (22) and reproduce the standard results obtained
 599 from combinatorial arguments.

600 If T_1 and T_2 are tensors of ranks n_1 and n_2 , respectively, then their tensor product is
 601 a tensor of rank $n_1 + n_2$. Each factor T_j transforms under index permutation according to
 602 some representation of \mathfrak{S}_{n_j} , and under linear transformation by the corresponding repre-
 603 sentation of $GL(D)$. It follows immediately that the tensor product $T_1 \otimes T_2$ transforms as
 604 some representation of $\mathfrak{S}_{n_1} \times \mathfrak{S}_{n_2}$. This induces a representation of the full permutation
 605 group $\mathfrak{S}_{n_1+n_2}$ which is associated with a corresponding representation of $GL(D)$. It is
 606 possible to reduce these last two representations into a sum of irreducible ones. We may
 607 assume that the factor representations are irreducible, since the original tensors T_j could
 608 have been broken into irreducible parts at the outset.

609 **Littlewood–Richardson rule:** The decomposition of an “outer product” $\mu \cdot \nu$ of irre-
 610 reducible representations μ and ν of \mathfrak{S}_{n_1} and \mathfrak{S}_{n_2} , respectively, into irreducible representa-
 611 tions of $\mathfrak{S}_{n_1+n_2}$ can be determined by means of the following algorithm involving Young
 612 diagrams. The product is commutative, so it does not matter which factor is regarded as
 613 the “right-hand” one. [In practice, one should choose the simpler Young diagram for that
 614 role.]

615 (I) Label each box in the top row of the right-hand diagram, ν , by “ a ”, each box in the
 616 second row by “ b ”, etc.

617 (II) Add the labeled boxes of ν to the left-hand diagram μ , one at a time, first the as ,
 618 then the bs , ..., subject to these constraints:

- 619 (A) No two boxes in the same column are labeled with the same letter;
- 620 (B) At all stages the result is a legitimate Young diagram;
- 621 (C) At each stage, if the letters are read right-to-left along the rows, from top to
 622 bottom, one never encounters more bs than as , more cs than bs , etc.

623 (III) Each of the distinct diagrams constructed in this way specifies an irreducible sub-
 624 representation λ , appearing in the decomposition of the outer product. The same
 625 labeled Young diagram may arise in more than one way; the multiplicity of that
 626 representation must be counted accordingly.

627 **Remarks:**

628 • This rule enables products of *distinct* tensors to be decomposed. When the factors
 629 are the same tensor, the list is further restricted by the requirement of symmetry under
 630 interchange of the factors. This is the problem of *plethysm*, whose solution requires more
 631 complicated techniques than the Littlewood–Richardson rule.

632 • Representations with too many parts (columns of length greater than D) must be deleted
 633 from the list of subrepresentations of the $GL(D)$. [If irreducible representations of the
 634 special linear group $SL(D)$ are considered instead, every column of length D must be
 635 removed from the corresponding Young diagram.]

636 **4.3 Orthogonal group**

637 It remains to consider index contractions. Up to now we considered only covariant ten-
 638 sors, because in the intended application there is a metric tensor which serves to relate
 639 contravariant and covariant tensors. Contractions are mediated by this metric. Implicitly,
 640 therefore, one is restricting the symmetry group of the problem from the general linear
 641 group to the subgroup that leaves the metric tensor invariant, the orthogonal group $O(D)$.
 642 [If the metric has indefinite signature, the true symmetry group is a non-compact ana-
 643 logue of the orthogonal group, such as the Lorentz group. This does not affect the relevant
 644 aspects of the *finite-dimensional* representation theory.] Each irreducible $GL(D)$ represen-
 645 tation $V_\lambda^{GL(D)}$ decomposes into irreducible $O(D)$ representations $V_\nu^{O(D)}$, labeled by Young
 646 diagrams ν obtained by removing an even number of boxes from λ . The *branching rule*
 647 for this process involves a sort of inverse of the Littlewood–Richardson rule:

648 **Restriction from $GL(D)$ to $O(D)$:** *The irreps of $GL(D)$ may be reduced to direct sums*
 649 *of irreps of $O(D)$ by extracting all possible trace terms formed by contraction with products*
 650 *of the metric tensor and its inverse.*

651 The reduction is given by the branching rule for $GL(D) \downarrow O(D)$:

$$V_\lambda^{GL(D)} = V_{\lambda/\Delta}^{O(D)} \equiv V_\lambda^{O(D)} \oplus V_{\lambda/\{2\}}^{O(D)} \oplus V_{\lambda/\{4\}}^{O(D)} \oplus V_{\lambda/\{2,2\}}^{O(D)} \oplus \dots \quad (24)$$

where Δ is the formal infinite sum [13]

$$\Delta = 1 + \begin{array}{|c|c|} \hline \square & \square \\ \hline \end{array} + \begin{array}{|c|c|c|c|} \hline \square & \square & \square & \square \\ \hline \end{array} + \begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \square \\ \hline \end{array} + \dots$$

652 corresponding to the sum of all possible plethysms of the metric tensor, and where λ/μ
 653 means the sum of the Young diagrams ν such that $\nu \cdot \mu$ contains λ according to the
 654 Littlewood–Richardson rule (with the corresponding multiplicity).

655 **Examples:**

656 • The $GL(D)$ irreducible representation labeled by the Young diagram $\{2, 2\}$ decomposes
 657 with respect to $O(D)$ according to the direct sum $\{2, 2\}/\Delta = \{2, 2\} + \{2, 0\} + \{0, 0\}$ which
 658 corresponds to the decomposition of the Riemann tensor into the Weyl tensor, the traceless
 659 part of the Ricci tensor and the scalar curvature, respectively.

660 • The $GL(D)$ irreducible representation labeled by the Young diagram $\{n\}$ decomposes
 661 with respect to $O(D)$ according to the direct sum $\{n\}/\Delta = \{n\} + \{n-2\} + \{n-4\} + \dots$,
 662 corresponding to the decomposition of a completely symmetric tensor or rank n into its
 663 traceless part, the traceless part of its trace, *etc.* This provides an alternative proof of the
 664 obvious fact that the number of independent components of a traceless symmetric tensor
 665 of rank n is equal to the number of independent components of a symmetric tensor of rank
 666 n minus the number of independent components of a symmetric tensor of rank $n-2$ (its
 667 trace): $\dim V_{\{n\}}^{O(D-2)} = \dim V_{\{n\}}^{GL(D)} - \dim V_{\{n-2\}}^{GL(D)}$. Using the formula (23) allows to show
 668 that

$$\dim V_{\{n\}}^{O(D)} = \frac{(D+2n-2)(D+n-3)!}{n!(D-2)!}. \quad (25)$$

669 The very useful formula (25) contains as a particular case the well-known fact that all the
 670 traceless symmetric tensorial representations of $O(2)$ are two-dimensional (indeed, any
 671 UIR of an Abelian group is of complex dimension one). Moreover, the traceless symmetric
 672 tensorial representations of rank n of the rotation group $O(3)$ are the well-known integer
 673 spin representations of dimension equal to $2n+1$.

674 The following theorem is very important (see e.g. the first reference of [11]):

675 **Vanishing irreps for (pseudo-)orthogonal groups:** *Whenever the sum of the lengths*
 676 *of the first two columns of a Young diagram λ is greater than $D = p+q$, then the irreducible*
 677 *representation of $O(p, q)$ labeled by λ is identically zero.*

678 Young diagrams such that the sum of the lengths of the first two columns does not
 679 exceed D are said to be *allowed*.

680 **Finite-dimensional irreps of (pseudo-)orthogonal groups:** *Each non-zero finite-*
 681 *dimensional irreducible representation of $O(p, q)$ is isomorphic to a completely traceless*
 682 *tensorial representation, the symmetry properties of which are labeled by an allowed Young*
 683 *diagram λ .*

684 The dimension of the tensorial irrep is determined by the following rule due to King [14]:

685 (α) The numbers $D - 1, D - 3, D - 5, \dots, D - 2r + 1$ are placed in the end boxes
 686 of the 1st, 2nd, 3rd, \dots , r th rows of the diagram λ . A labeled Young diagram of
 687 n numbers is then constructed by inserting in the remaining boxes of the diagram,
 688 numbers which increase by one in passing from one box to its left-hand neighbor.

689 (β) This labeled Young diagram is extended to the limit of the triangular Young diagram
 690 τ of r rows. This produces a Young diagram $\tilde{\lambda}$ the a th row of which has length equal
 691 the maximum between the two integers $\tau_a = r - a + 1$ and λ_a .

692 (γ) The series of numbers in any row of the Young diagram $\tilde{\lambda}$ is then extended by
 693 inserting in the remaining boxes of the diagram, numbers which decrease by one
 694 in passing from one box to its right-hand neighbor. The resulting numbers will be
 695 called the “King length.”

696 (δ) The row lengths $\lambda_1, \lambda_2, \dots, \lambda_r$ are then added to all of the numbers of the Young
 697 diagram $\tilde{\lambda}$ which lie on lines of unit slope passing through the first box of the 1st,
 698 2nd, \dots , r th rows, respectively, of the Young diagram λ .

699 The dimension is equal to the product of the integers in the resulting labeled Young
 700 diagram $\tilde{\lambda}$ divided by the product of

- 701 - the hook length of each box of λ , and of
- 702 - the King length of each box of $\tilde{\lambda}$ outside λ .

703 **Examples:**

704 • In the following diagram, allowed for $D = 5$,



706 each box is labeled by its King length, while in the diagram



708 each box is labeled by the number obtained at the very end of King’s rule. Observe that,
 709 for the corresponding diagram λ , it was not necessary to perform the steps (β)-(γ) and
 710 that, $\dim V_\lambda^{O(5)} = 231 < 1050 = \dim V_\lambda^{GL(5)}$.

711 • In the following Young diagram $\lambda = \{2, 2, 1\}$, allowed for $D = 5$,

$$\begin{array}{|c|c|} \hline 5 & 4 \\ \hline 3 & 2 \\ \hline 0 & \\ \hline \end{array},$$

712

713 each box is labeled by the number obtained after step (α) . The step (β) is now necessary
 714 and gives the Young diagram $\tilde{\lambda} = \{3, 2, 1\}$. At the end of steps (γ) and (δ) , respectively,
 715 the result is

$$\xrightarrow{(\gamma)} \begin{array}{|c|c|c|} \hline 5 & 4 & 3 \\ \hline 3 & 2 & \\ \hline 0 & & \\ \hline \end{array} \xrightarrow{(\delta)} \begin{array}{|c|c|c|} \hline 7 & 6 & 4 \\ \hline 5 & 3 & \\ \hline 1 & & \\ \hline \end{array},$$

716

717 so that $\dim V_{\lambda}^{O(5)} = \frac{7 \cdot 6 \cdot 5 \cdot 4 \cdot 3}{(4 \cdot 3 \cdot 2) \cdot (3)} = 35 < 75 = \dim V_{\lambda}^{GL(5)}$.

718 • The space of traceless symmetric tensors of V of rank n carries an irreducible represen-
 719 tation of $O(D)$ labeled by a Young diagram made of one row of length n for which the
 720 dimension (25) is easily reproduced from the King rule, since the rules (β) - (γ) may be
 721 omitted

722 • Computing the number of components of the Weyl tensor and of a symmetric, traceless,
 723 rank-two tensor in $D = 4$ dimensions, enables one to give the decomposition $\{2, 2\}/\Delta =$
 724 $\{2, 2\} + \{2, 0\} + \{0, 0\}$ of the Riemann tensor into the Weyl tensor, the traceless part
 725 of the Ricci tensor and the scalar curvature, respectively, in terms of the corresponding
 726 dimensions. This gives the well-known result $20 = 10 + 9 + 1$.

727 **Unitary irreps of orthogonal groups:** *Each non-zero inequivalent UIR of $O(D)$ cor-*
 728 *responds to an allowed Young diagram λ , and conversely.*

729 Proof: The orthogonal group is compact, thence any UIR is finite-dimensional (see Subsec-
 730 tion 1.3). Furthermore, any finite-dimensional irrep of the orthogonal group is labeled by
 731 an allowed Young diagram. Moreover, an important result is that any finite-dimensional
 732 representation may be endowed with a sesquilinear form which makes it unitary. \square

733 The quadratic Casimir operator of the orthogonal algebra $\mathfrak{so}(D)$ presented by its gen-
 734 erators and its commutation relations

$$i [M_{\mu\nu}, M_{\rho\sigma}] = \delta_{\nu\rho} M_{\mu\sigma} - \delta_{\mu\rho} M_{\nu\sigma} - \delta_{\sigma\mu} M_{\rho\nu} + \delta_{\sigma\nu} M_{\rho\mu} \quad (26)$$

735 is the sum of square of the generators (similarly to the definition (4) for $\mathfrak{so}(D - 1, 1)$ since
 736 these two *complex* algebras are isomorphic). Its eigenvalue on a finite-dimensional irrep
 737 labeled by an allowed Young diagram $\lambda = \{\lambda_1, \lambda_2, \dots, \lambda_r\}$ is given in the subsection 9.4.C
 738 of [2]:

$$\left[\mathcal{C}_2(\mathfrak{so}(D)) - \sum_{a=1}^r \lambda_a (\lambda_a + D - 2a) \right] V_{\lambda}^{O(D)} = 0. \quad (27)$$

739 **Examples:**

740 • The UIRs of the Abelian group $O(2) \cong U(1)$ are labeled by one integer only, which is
 741 the eigenvalue of the single generator on the irrep, say $h \in \mathbb{Z}$. The only allowed Young
 742 diagrams are made of a single row of length equal to the non-negative integer $s = |h|$. The
 743 traceless symmetric tensorial representations of $O(2)$ are two-dimensional, the sum of the
 744 two irreps labeled by $h = \pm s$. The formula (27) with $D = 2$, $r = 1$ and $\lambda_1 = s$ gives the
 745 obvious eigenvalue s^2 , since the quadratic Casimir operator of the rotation group $O(2)$ is
 746 equal to the square of the single generator.

747 • The quadratic Casimir operator of the rotation group $O(3)$ is the square of the angular
 748 momentum. The irrep of $O(3)$ with spin $s \in \mathbb{N}$ is labeled by the allowed Young diagram
 749 made of a single row of length equal to the integer s . The formula (27) with $D = 3$, $r = 1$
 750 and $\lambda_1 = s$ gives the celebrated eigenvalue $s(s + 1)$.

751 • The irrep of $O(D)$ carried by the space of traceless symmetric tensors of rank n is labeled
 752 by the allowed Young diagram $\{n\}$ made of a single row of length equal to an integer n .
 753 The formula (27) with $r = 1$ and $\lambda_1 = n$ gives the eigenvalue $n(n + D - 2)$ for the quadratic
 754 Casimir operator.

755 The following branching rule is extremely useful in the process of dimensional reduc-
 756 tion.

Restriction from $GL(D)$ to $GL(D - 1)$: *The restriction to the subgroup $GL(D - 1) \subset GL(D)$ of a finite-dimensional irrep of $GL(D)$ determined by the Young diagram λ contains each irrep of $GL(D - 1)$ labeled by Young diagrams μ such that*

$$\lambda_1 \geq \mu_1 \geq \lambda_2 \geq \mu_2 \geq \dots \geq \mu_{r-1} \geq \lambda_r \geq \mu_r \geq 0,$$

757 *with multiplicity one. The same theorem holds for the restriction $O(D) \downarrow O(D - 1)$ where*
 758 *λ is an allowed Young diagram.*

759 These rules are discussed in the section 8.8.A of [2]. They may be summarized in the
 760 following branching rule for $GL(D) \downarrow GL(D - 1)$,

$$V_\lambda^{GL(D)} = V_{\lambda/\Sigma}^{GL(D-1)} \equiv V_\lambda^{GL(D-1)} \oplus V_{\lambda/\{1\}}^{GL(D-1)} \oplus V_{\lambda/\{2\}}^{GL(D-1)} \oplus V_{\lambda/\{3\}}^{GL(D-1)} \oplus \dots \quad (28)$$

761 where Σ is the formal infinite sum of all Young diagrams made of a single row.

Example: The branching rule applied to symmetric irrep labeled by a Young diagram $\{n\}$ made of one row of length n gives as a result:

$$\{n\}/\Sigma = \{n\} + \{n - 1\} + \{n - 2\} + \dots + \{1\} + \{0\}.$$

762 This implies the obvious fact that a completely symmetric tensor of rank n whose indices
 763 run over D values may be decomposed as a sum of completely symmetric tensors of rank
 764 $n, n - 1, \dots, 1, 0$ whose indices run over $D - 1$ values. A non-trivial instance of the
 765 branching rule for $O(D) \downarrow O(D - 1)$ is that the same result is true for *traceless* symmetric
 766 tensors as well.

767 4.4 Auxiliary variables

768 Let λ be a Young diagram with s columns and r rows.

769 The Schur module $V_\lambda^{GL(D)}$ in the “manifestly antisymmetric convention” can be built
 770 *via* a convenient construction in terms of polynomials in $s \times D$ graded variables satisfy-
 771 ing appropriate conditions. More precisely, the vector space $V_\lambda^{GL(D)}$ is isomorphic to a
 772 subspace of the associative algebra

$$\mathcal{A} = (\otimes^s \wedge \mathbb{R}^{D*}) \otimes C^\infty(\mathbb{R}^D) = \otimes_{C^\infty(\mathbb{R}^D)}^s \Omega(\mathbb{R}^D) \quad (29)$$

773 of s tensor products of antisymmetric forms. The elements of \mathcal{A} are called *multiforms* [15].

774 The D generators of the I th factor \mathbb{R}^{D*} in $(\otimes^s \wedge \mathbb{R}^{D*})$ are written $d_I x^\mu$ ($\mu = 0, 1, \dots, D$).
 775 By definition, the multiform algebra \mathcal{A} is presented by the graded commutation relations

$$d_I x^\mu d_J x^\nu = (-)^{\delta_{IJ}} d_J x^\nu d_I x^\mu, \quad (30)$$

776 where the wedge products are not written explicitly. The condition (i) of Subsection 4.2
 777 is automatically verified for any element $\Phi \in \mathcal{A}$ due to the fact that the variables are
 778 anticommuting in a fixed column ($I = J$). The $GL(D)$ -irreducibility condition (ii) of
 779 Subsection 4.2 is implemented by the conditions

$$\left(d_I x \cdot \frac{\partial^L}{\partial (d_J x)} - \delta_{IJ} \ell_I \right) \Phi = 0, \quad (I \leq J) \quad (31)$$

780 where the dot stands for the contraction of the indices, ℓ_I for the length of the I th column
 781 in the Young diagram λ and ∂^L stands for “left” derivative. By the Weyl construction,
 782 an element $\Phi \in \mathcal{A}$ satisfying (31) belongs to the Schur module $V_\lambda^{GL(D)}$. Following the
 783 discussion of Subsection 4.3, if λ denotes an allowed Young diagram, such an element
 784 $\Phi \in V_\lambda^{GL(D)}$ is irreducible under the (pseudo)-orthogonal group $O(p, q)$ ($p + q = D$) if it
 785 is traceless, that is

$$\left(\frac{\partial^L}{\partial(d_I x)} \cdot \frac{\partial^L}{\partial(d_J x)} \right) \Phi = 0, \quad (\forall I, J) \quad (32)$$

786 where the dot stands now for the contraction of indices via the use of the metric preserved
 787 by $O(p, q)$. An element $\Phi \in \mathcal{A}$ such that (31)-(32) are fulfilled belongs to the Schur module
 788 $V_\lambda^{O(p,q)}$ labeled by the Young diagram λ .

789 The Schur module $V_\lambda^{GL(D)}$ admits another convenient realization in terms of polyno-
 790 mials in $r \times D$ commuting variables. In other words, the vector space $V_\lambda^{GL(D)}$ is isomorphic
 791 to a subspace of the polynomial algebra in the variables u_a^μ ($a = 1, 2, \dots, r$) where the
 792 index a corresponds to each row. The condition (a) of Subsection 4.2 is automatically
 793 verified for any such polynomial due to the fact that the variables are commuting in a
 794 fixed row. The $GL(D)$ -irreducibility condition (b) of Subsection 4.2 is implemented by
 795 the conditions

$$\left(u_a \cdot \frac{\partial}{\partial u_b} - \delta_{ab} \lambda_a \right) \Phi = 0, \quad (a \leq b) \quad (33)$$

796 where the dot still stands for the contraction of the indices. The degree of homogeneity of
 797 the polynomial Φ in the variables u_a^μ (for fixed a) is λ_a . The corresponding coefficients are
 798 tensors irreducible under the general linear group. By the Weyl construction, a polynomial
 799 $\Phi(u_a)$ satisfying (33) belongs to the Schur module $V_\lambda^{GL(D)}$. Again, such an element
 800 $\Phi \in V_\lambda^{GL(D)}$ is irreducible under the (pseudo)-orthogonal group $O(p, q)$ ($p + q = D$) iff it
 801 is traceless, that is

$$\left(\frac{\partial}{\partial u_a} \cdot \frac{\partial}{\partial u_b} \right) \Phi = 0, \quad (\forall a, b) \quad (34)$$

802 where the dot stands for the contraction of indices via the use of the metric preserved by
 803 $O(p, q)$. A polynomial $\Phi(u_a)$ such that (33)-(34) are fulfilled belongs to the Schur module
 804 $V_\lambda^{O(p,q)}$ labeled by an allowed Young diagram λ .

805 **Example:** Consider an irreducible representation of the orthogonal group $O(D)$ labeled
 806 by the Young diagram $\{n\}$ made of a single row of length equal to an integer n . The
 807 polynomial $\Phi(u) \in V_{\{n\}}^{O(D)}$ obeys to the irreducibility conditions

$$\left(u \cdot \frac{\partial}{\partial u} - n \right) \Phi = 0, \quad \left(\frac{\partial}{\partial u} \cdot \frac{\partial}{\partial u} \right) \Phi = 0. \quad (35)$$

They mean that the polynomial is homogeneous (of degree equal to n) and harmonic, so
 that its components correspond to a symmetric traceless tensor of rank n :

$$\Phi(u) = \frac{1}{n!} \Phi_{\mu_1 \dots \mu_n} u^{\mu_1} \dots u^{\mu_n}, \quad \delta^{\mu_1 \mu_2} \Phi_{\mu_1 \mu_2 \mu_3 \dots \mu_n} = 0.$$

808 Of course the integral of the square of such a polynomial over \mathbb{R}^D is, in general, infinite.
 809 But the restriction of an harmonic polynomial on the unit sphere $\vec{u}^2 = 1$ is square
 810 integrable on S^{D-1} . This restriction is called a *spherical harmonic* of degree n . Therefore
 811 the space of spherical harmonics of degree n provides an equivalent realization of the Schur
 812 module $V_{\{n\}}^{O(D)}$. For $D = 3$, the space $V_{\{n\}}^{O(3)}$ is spanned by the usual spherical harmonics
 813 $Y_n^m(\theta, \phi)$ on the two-sphere with $|m| \leq n$.

814 **Remarks:**

- The infinitesimal generators of the pseudo-orthogonal group $O(p, q)$ are represented by the operators

$$M_{\mu\nu} = i \sum_{a=1}^r u_a^\rho \left(g_{\rho\mu} \frac{\partial}{\partial u_a^\nu} - g_{\rho\nu} \frac{\partial}{\partial u_a^\mu} \right).$$

815 Reordering the factors and making use of (33)-(34) allows to reproduce the formula (27)
816 for the eigenvalues of the quadratic Casimir operator.

- Instead of polynomial functions in the commuting variables, one may equivalently consider *distributions* obeying to the same conditions. The space of solutions would carry an equivalent irrep, as follows from the highest-weight construction of the representation. However, it does not make sense any more of talking about the “coefficients” of the homogeneous distribution so that the link with the equivalent tensorial representation is more intricate.

823 The example of the spherical harmonics suggests that it might be convenient to realize
824 any unitary module of the orthogonal group $O(D)$ as a space of functions on the unit hy-
825 persphere S^{D-1} satisfying some linear differential equations. Better, the symmetry under
826 the orthogonal group would be made manifest by working with homogeneous harmonic
827 functions on the *ambient space* \mathbb{R}^D , evaluated on any hypersphere $S^{D-1} \subset \mathbb{R}^D$.

828 **Spherical harmonics:** *To any UIR of the isometry group $O(D)$ of a hypersphere S^{D-1} ,*
829 *one may associate manifestly covariant differential equations for functions on S^{D-1} em-*
830 *bedded in \mathbb{R}^D whose space of solutions carry the corresponding UIR.*

831 Proof: Any UIR of the isometry group $O(D)$ corresponds to a Schur module $V_\lambda^{O(D)}$ which
832 may be realized as the space of polynomials $\Phi(\vec{u}_a)$ such that (33)-(34) are obeyed. Let
833 us introduce the notation: $\vec{x} := \vec{u}_1$ and $\vec{t}_{a-1} := \vec{u}_a$ for $a = 2, \dots, r$. One interprets
834 the polynomial $\Phi(\vec{x}, \vec{t}_a)$ (where the index a runs from 1 to $r-1$) as a tensor field on the
835 Euclidean space \mathbb{R}^D parametrized by the Cartesian coordinates \vec{x} , with some auxiliary
836 variables \vec{t}_a implementing the tensor components. The conditions (33)-(34) for a and b
837 strictly greater than 1 imply that

$$\left(t_a \cdot \frac{\partial}{\partial t_b} - \delta_{ab} \lambda_a \right) \Phi = 0, \quad (a \leq b) \quad \left(\frac{\partial}{\partial t_a} \cdot \frac{\partial}{\partial t_b} \right) \Phi = 0, \quad (36)$$

where $\underline{\lambda} = \{\lambda_2, \dots, \lambda_r\}$ is the Young diagram obtained from λ by removing its first row.
Thus the components of the “tensor field” $\Phi(\vec{x}, \vec{t}_a)$ carry an irreducible representation
of $O(D)$ labeled by $\underline{\lambda}$. The conditions (33) for $a = b = 1$ imply that

$$\left(x \cdot \frac{\partial}{\partial x} - \lambda_1 \right) \Phi = 0,$$

838 so the polynomial $\Phi(\vec{x}, \vec{t}_a)$ is homogeneous of degree λ_1 in the radial coordinate $|\vec{x}|$.
839 The condition (34) for $a = b = 1$ is interpreted as the Laplace equation

$$\left(\frac{\partial}{\partial x} \cdot \frac{\partial}{\partial x} \right) \Phi = 0 \quad (37)$$

840 on the ambient space \mathbb{R}^D , it implies that the tensor field Φ is harmonic in ambient space.
841 The condition (33) for $b > a = 1$ states that the radial components vanish,

$$\left(x \cdot \frac{\partial}{\partial t_a} \right) \Phi = 0, \quad (38)$$

842 so the tensor components are longitudinal to the hyperspheres S^{D-1} . Therefore the evalu-
 843 ation of the non-vanishing components of $\Phi(\vec{x}, \vec{t}_a)$ on the unit hypersphere $|\vec{x}| = 1$ is an
 844 *intrinsic* tensor field living on the hypersphere S^{D-1} and whose tensor components carry
 845 an irrep of the stability subgroup $O(D-1)$ labeled by $\underline{\lambda}$. These tensor fields generalize the
 846 spherical harmonics to the generic case $r \geq 1$. Finally, the condition (34) for $b > a = 1$
 847 states that the tensor field is divergenceless in ambient space,

$$\left(\frac{\partial}{\partial x} \cdot \frac{\partial}{\partial t_a} \right) \Phi = 0. \quad (39)$$

848 The differential equations (37) and (39) are written in ambient space but they may be
 849 reformulated in intrinsic terms on the hypersphere, at the price of losing the manifest
 850 covariance under the full isometry group $O(D)$. \square

851 4.5 Euclidean group

852 The method of induced representations was introduced in Subsection 3.1 for the Poincaré
 853 group $ISO(D-1, 1)^\uparrow$ and applied to the Euclidean group $ISO(D-2)$ in Subsection
 854 3.3. Focusing on the faithful (*i.e.* with a non-trivial action of the translation generators)
 855 irreps of the *inhomogeneous* orthogonal group, all of them are induced from an UIR of the
 856 stability subgroup. Using the results of the previous section 4.3, one may summarize the
 857 final result into the following classification.

858 **Unitary irreps of the inhomogeneous orthogonal groups:** *Each inequivalent UIR*
 859 *of the group $IO(D)$ with a non-trivial action of its Abelian normal subgroup is associated*
 860 *with a positive real number μ and an allowed Young diagram of the subgroup $O(D-1)$,*
 861 *and conversely.*

862 The orbits of the linear action of the orthogonal group $O(D)$ on the Euclidean space
 863 \mathbb{R}^D are the hyperspheres S^{D-1} of radius R . The isometry group of any such hypersphere
 864 S^{D-1} is precisely $O(D)$. Considering a region of fixed size on these hyperspheres, in
 865 the limit $R \rightarrow \infty$ the sphere becomes a hyperplane \mathbb{R}^{D-1} . Therefore the homogeneous
 866 and inhomogeneous orthogonal groups are related by some infinite radius limit: $O(D) \rightarrow$
 867 $IO(D-1)$. Such a process is frequently referred to as an *Inönü-Wigner contraction* in the
 868 physics literature [16]. This is better seen at the level of the Lie algebra. Specializing the
 869 D th directions, the commutation relations (26) take the form

$$i [M_{mn}, M_{pq}] = \delta_{np} M_{mq} - \delta_{mp} M_{nq} - \delta_{qm} M_{pn} + \delta_{qn} M_{pm}, \quad (40)$$

$$i [M_{mD}, M_{pq}] = \delta_{mn} M_{pD} - \delta_{mp} M_{nD}, \quad (41)$$

$$i [M_{mD}, M_{pD}] = M_{pm}. \quad (42)$$

870 where the latin letters take $D-1$ values. Defining $M_{mD} = R P_m$ and taking the limit
 871 $R \rightarrow \infty$ (with P_m fixed) in the relations (40)-(42) lead to

$$i [M_{mn}, M_{pq}] = \delta_{np} M_{mq} - \delta_{mp} M_{nq} - \delta_{qm} M_{pn} + \delta_{qn} M_{pm}, \quad (43)$$

$$i [P_m, M_{pq}] = \delta_{mn} P_p - \delta_{mp} P_n, \quad (44)$$

$$i [P_m, P_p] = 0. \quad (45)$$

872 As can be seen, the generators $\{M_{mn}, P_m\}$ span the Lie algebra of the inhomogeneous
 873 orthogonal group $IO(D-1)$. The former argument proves the contraction $\mathfrak{so}(D) \rightarrow$
 874 $\mathfrak{iso}(D-1)$.

875 The limit of a sequence of irreps of the homogeneous orthogonal group $O(D)$, in which
 876 one performs an Inönü-Wigner contraction, is automatically a representation of the inho-
 877 mogeneous orthogonal group $IO(D-1)$ (if the limit is not singular). An interesting issue

878 is the inverse problem: which irreps of $IO(D-1)$ may be obtained as the limit of such a
 879 sequence of irreps of $O(D)$? The problem is non-trivial because, generically, the limit of
 880 a sequence of irreps is a *reducible* representation.

881 **Contraction of UIRs of the homogeneous orthogonal groups:** *Each inequivalent*
 882 *UIR of the group $IO(D-1)$ with a non-trivial action of its Abelian normal subgroup may*
 883 *be obtained as the contraction of a sequence of UIRs of the group $O(D)$.*

884 More precisely, the Inönü-Wigner contraction $R \rightarrow \infty$ of a sequence of UIRs of $O(D)$,
 885 labeled by allowed Young diagrams $\nu = \{s, \lambda_1, \dots, \lambda_r\}$ such that the limit of the quotient
 886 s/R is a fixed positive real number μ , is the UIR of $IO(D-1)$ labeled by the parameter
 887 μ and the Young diagram $\lambda = \{\lambda_1, \dots, \lambda_r\}$.

888 Proof: The use of the spherical harmonics construction discussed at the end of Subsection
 889 4.4 is very convenient here. The main idea is to solve the homogeneity condition in a
 890 neighborhood of $x^D \neq 0$ as follows:

$$\Phi(x^m, x^D, t_a) = z^s \phi\left(\frac{x^m}{z}, t_a\right), \quad (46)$$

where $\vec{x} = (x^m, x^D)$ and $\phi(y^m, t_a) := \Phi(y^m, \frac{s}{\mu}, t_a)$. In other words, one may perform a
 convenient change of coordinates from the *homogenous coordinates* (x^m, x^D) to the set
 (y^m, z) where

$$y^m = \frac{x^m}{z}$$

are the *inhomogenous coordinates* (on the projective space $\mathbb{P}\mathbb{R}^{D-1}$ minus the point at
 infinity $z = 0$) and

$$z = \frac{\mu x^D}{s}$$

891 is a scale variable. The magic is that the equations for the generalized spherical harmonics
 892 have a well-behaved limit $x^D \rightarrow \infty$ in terms of $\phi(y^m, t_a)$ when x^D/s is fixed to be equal
 893 to the ratio z/μ , where z and μ are finite [17]. To see that, one should use the relations

$$\begin{aligned} \frac{\partial}{\partial x^m} &= \frac{1}{z} \frac{\partial}{\partial y^m}, \\ \frac{\partial}{\partial x^D} &= \frac{\mu}{s} \left(\frac{\partial}{\partial z} - \frac{1}{z} y^m \frac{\partial}{\partial y^m} \right). \end{aligned} \quad (47)$$

894 Moreover, the equations in this limit may be identified with equations for the proper UIR
 895 of the inhomogeneous orthogonal group $IO(D-1)$ realized homogeneously in terms of the
 896 inhomogenous coordinates. \square

Example: The simplest instance is when $\lambda = \{0\}$ because one considers the sequence of
 harmonic functions $\Phi(x^m, x^D)$ of homogeneity degree s . The Laplace operator acting on
 $\Phi(x^m, x^D)$ reads in terms of $\phi(y^m)$ as follows

$$\Delta_{\mathbb{R}^D} \Phi = z^{s-2} \left[\frac{\partial}{\partial y} \cdot \frac{\partial}{\partial y} + \frac{\mu^2}{s^2} \left(s(s-1) - (2s-1) \left(y \cdot \frac{\partial}{\partial y} \right) + \left(y \cdot \frac{\partial}{\partial y} \right)^2 \right) \right] \phi,$$

due to the homogeneity condition (46) and the relations (47). The Laplace equation
 $\Delta_{\mathbb{R}^D} \Phi = 0$ is thus equivalent to the equation

$$\left[\frac{\partial}{\partial y} \cdot \frac{\partial}{\partial y} + \frac{\mu^2}{s^2} \left(s(s-1) - (2s-1) \left(y \cdot \frac{\partial}{\partial y} \right) + \left(y \cdot \frac{\partial}{\partial y} \right)^2 \right) \right] \phi = 0,$$

897 whose limit for $s \rightarrow \infty$ is the Helmholtz equation $[\Delta_{\mathbb{R}^{D-1}} + \mu^2] \phi = 0$, where $\Delta_{\mathbb{R}^{D-1}} =$
 898 $\frac{\partial}{\partial y} \cdot \frac{\partial}{\partial y}$. The space of solutions of the Helmholtz equation carries an UIR of $IO(D-1)$
 899 induced from a trivial representation of the stability subgroup $O(D-2)$.

900 5 Relativistic field equations

901 The *Bargmann-Wigner programme* amounts to associating, with any given UIR of the
 902 Poincaré group, a manifestly covariant differential equation whose (positive-energy) so-
 903 lutions transform according to the corresponding UIR. Physically, it might be natural
 904 to restrict this programme to the two most important classes of UIRs: the massive and
 905 massless representations. Mathematically, this restriction is convenient because the group-
 906 theoretical analysis is simpler since any of these UIRs is induced from an UIR of a uni-
 907 modular orthogonal group $SO(n)$ (with $D-3 \leq n \leq D-1$), as can be checked easily on
 908 the tables of Subsection 3.3.

909 In 1948, this restricted programme was completed by Bargmann and Wigner in four
 910 dimensions when, for each such UIR of $ISO(3,1)^\dagger$, a relativistic field equation was written
 911 whose positive-energy solutions transform according to the corresponding UIR [4]. But
 912 this case ($D=4$) will not be reviewed here in details because it may cast shadow on the
 913 generic case. Indeed, it is rather peculiar in many respects:

- 914 • The quadratic and quartic Casimir operators essentially classify the UIRs, but this is
 915 no more true in higher dimensions where more Casimir operators are necessary and
 916 the classification quickly becomes technically cumbersome in this way. Moreover,
 917 one should stress that the eigenvalues of the Casimir operators do not character-
 918 ize uniquely an irreducible representation (for instance, the quadratic and quartic
 919 Casimir operators vanish for all helicity representations).
- 920 • The (complex) Lorentz algebra $\mathfrak{so}(3,1)$ is isomorphic to the direct sum of two (com-
 921 plex) rotation algebras $\mathfrak{so}(3) \cong \mathfrak{sp}(2)$. These isomorphisms allow the use of the
 922 convenient “dotted-undotted” formalism for the finite-dimensional (non-unitary) ir-
 923 reps of the spin group $Spin(3,1)$.
- 924 • The symmetric tensor-spinor fields are sufficient to cover all inequivalent cases.
- 925 • The helicity short little group $SO(2)$ is Abelian, therefore its irreps are one-dimensional,
 926 for fixed helicity. Notice that the helicity is discretized because the representa-
 927 tion of the “little group” $SO(2)$ is a restriction of the representation of the group
 928 $Spin(3) \cong SU(2)$ which has no intrinsically projective representations.
- 929 • The infinite-spin short little group $SO(1)$ is trivial, thus there are only two inequiv-
 930 alent infinite-spin representations (single- or double-valued) [6].
- 931 • *etc.*

932 Moreover, there exists an extensive literature on the subject of UIRs of $ISO(3,1)^\dagger$ and
 933 we refer to the numerous pedagogical reviews available for more details on the four-
 934 dimensional case (see *e.g.* the inspiring presentations of [5] and [19]).

935 It is standard to require time reversal and parity symmetry of the field theory. More
 936 precisely, the field equations we will consider are covariant under the two previous transfor-
 937 mations. As a consequence of the time reversal symmetry, the representation is *irreducible*
 938 under the group $ISO(D-1,1)$ but *reducible* under the Poincaré group $ISO(D-1,1)^\dagger$:

939 the Hilbert space of solutions contain both positive and negative energy solutions. Fur-
 940 thermore, the parity symmetry implies that the representation is *irreducible* under the
 941 inhomogeneous Lorentz group $IO(D-1, 1)$ but *reducible* under the group $ISO(D-1, 1)$
 942 (for instance, both chiralities are present in the massless case for D even). To conclude, the
 943 Bargmann-Wigner programme is actually understood as associating, with any given UIR
 944 of the inhomogeneous Lorentz group, a manifestly covariant differential equation whose
 945 solutions transform according to the corresponding UIR.

946 5.1 General procedure

947 The lesson on induced representations that we learned from Wigner implies the following
 948 strategy:

- 949 1. Pick a unitary representation of the (short) little group.
- 950 2. Introduce a wave function on $\mathbb{R}^{D-1,1}$ taking values in some (possibly non-unitary)
 951 representation of the Lorentz group $O(D-1, 1)$ the restriction of which to the (short)
 952 little group contains the representation of step 1.
- 953 3. Write a system of linear covariant equations, differential in position space x^μ thus
 954 algebraic in momentum space p_ν , for the wave function of step 2. These equations
 955 may not be independent.
- 956 4. Fix the momentum and check in convenient coordinates that the field equations of
 957 step 3 put to zero all “unphysical” components of the wave function. More precisely,
 958 verify that its non-vanishing components carry the unitary representation of step 1.

959 Proof: The fact that the set of linear differential equations is taken to be manifestly
 960 covariant ensures that the Hilbert space of their solutions carries a (infinite-dimensional)
 961 representation of $IO(D-1, 1)$. The fourth step determines the representation of the little
 962 group by which it is induced. \square

963 In the physics literature, the fourth step is referred to as “looking at the physical
 964 degrees of freedom.” If the (possibly reducible) representation is proven to be unitary,
 965 then this property is summarized in a “no-ghost theorem.”

966 The Klein-Gordon equation $(p^2 \pm m^2)\Psi = 0$ is always, either present in the system of
 967 covariant equations or a consequence thereof. Consequently, the Klein-Gordon equation
 968 will be assumed implicitly from now on in the step 3. Therefore, the step 4 will be
 969 immediately performed in a proper Lorentz frame. (We refer the reader to the Subsection
 970 3.2 for more details.)

971 The two completions [20] and [21] of the Bargmann-Wigner programme for finite-
 972 component representations in Minkowski spacetime of dimension $D > 3$ are reviewed,
 973 respectively, in the appendix A and in the subsections 5.2-5.3 for single-valued UIRs of
 974 the Poincaré group.⁶

975 The tachyonic case⁷ is more briefly discussed in Subsection 5.4. The zero-momentum
 976 representations are not considered here since they essentially are the unitary irreducible
 977 representations of the de Sitter spacetime dS_{D-1} . The latter have been reviewed in [23].

978 The Bargmann-Wigner programme for fractional-spin fields in three spacetime di-
 979 mensions has been completed in [25]. More generally, the exhaustive completion of the

⁶Spinorial irreps may be adressed analogously by supplementing the system of differential equations with Dirac-like equations and gamma-trace constraints (see *e.g.* [17, 22] for more details).

⁷The discussion presented in the section 5.4 was not published before, it directly derives from private conversations between X.B. and J. Mourad.

980 Bargmann-Wigner programme (for all representations) in Minkowski spacetime of dimen-
981 sion $D = 3$ is briefly summarised in Appendix B.

982 5.2 Massive representations

983 The Bargmann-Wigner programme is easy to complete for massive UIRs because the
984 massive stability subgroup is the orthogonal group $O(D - 1) \subset O(D - 1, 1)$. By going
985 to a rest-frame, the time-like momentum vector takes the form $p^\mu = (m, \vec{0}) \neq 0$. The
986 physical components of the field are thus carrying a tensorial irrep of the group $O(D -$
987 $1)$ of orthogonal transformations in the spatial hyperplane \mathbb{R}^{D-1} orthogonal to p^μ . In
988 other words, the linear field equations should remove all components including time-like
989 directions. These unphysical components are responsible for the fact that the Fock space
990 is not endowed with a positive-definite norm.

991 **Step 1.** From the sections 1.3 and 4, one knows that any unitary representation of
992 the orthogonal group $O(D - 1)$ is a sum of UIRs which are finite-dimensional and thus,
993 equivalent to a tensorial representation. Let us consider the UIR of $O(D - 1)$ labeled by
994 the allowed Young diagram $\lambda = \{\lambda_1, \lambda_2, \dots, \lambda_r\}$ (*i.e.* the sum of the lengths of its first
995 two columns does not exceed $D - 1$).

996 **Step 2.** The simplest way to perform the Bargmann-Wigner programme in the
997 massive case is to choose a covariant wave function whose components carry the (finite-
998 dimensional and non-unitary) tensorial irrep of the Lorentz group $O(D - 1, 1)$ labeled by
999 the Young diagram λ . As explained in the subsection 4.4, a convenient way of realiz-
1000 ing this is in terms of a wave function $\Phi(p, u_a)$ polynomial in the auxiliary commuting
1001 variables u_a^μ satisfying the irreducibility conditions (33)-(34).

1002 **Step 3.** The massive Klein-Gordon equation

$$(p^2 + m^2)\Phi = 0 \quad (48)$$

1003 has to be supplemented with the transversality conditions

$$\left(p \cdot \frac{\partial}{\partial u_a}\right)\Phi = 0, \quad (49)$$

1004 of the wave function.

1005 **Step 4.** Looking at a fixed-momentum mode in its corresponding rest-frame $p^\mu =$
1006 $(m, \vec{0})$ leads to the fact that the components of the wave function along the timelike
1007 momentum are set to zero by (49): $\Phi = \Phi(p, \vec{u}_a)$. In words, Φ does not depend on the
1008 time components $u_a^0, \forall a$. In this case, the conditions (33)-(34) read as irreducibility
1009 conditions under the orthogonal group $O(D - 1)$. \square

1010 **Example:** Massive symmetric representations with “spin” equal to s correspond to Young
1011 diagrams $\lambda = \{s\}$ made of one row of length equal to the integer s . In four spacetime
1012 dimensions, this representation is precisely what is usually called a “massive spin- s field.”⁸
1013 The covariant wave function $\Phi(p, u)$ obeys to the irreducibility conditions (33)-(34) of the
1014 components

$$\left(u \cdot \frac{\partial}{\partial u} - s\right)\Phi = 0, \quad \left(\frac{\partial}{\partial u} \cdot \frac{\partial}{\partial u}\right)\Phi = 0. \quad (50)$$

The wave function Φ is homogeneous of degree s and harmonic in the auxiliary variable u .
If the wave function $\Phi(p, u)$ is polynomial in the auxiliary variable u , then its components

⁸To our knowledge, the Bargmann-Wigner programme for the massive integer-spin representations in four-dimensional Minkowski spacetime was addressed along the lines reviewed here for the first time by Fierz in [26].

correspond to a symmetric tensor of rank s

$$\Phi(p, u) = \frac{1}{s!} \Phi_{\mu_1 \dots \mu_s}(p) u^{\mu_1} \dots u^{\mu_s},$$

1015 which is traceless

$$\eta^{\mu_1 \mu_2} \Phi_{\mu_1 \mu_2 \mu_3 \dots \mu_s}(p) = 0. \quad (51)$$

1016 The covariant field equations are the massive Klein-Gordon equation together with the
1017 transversality condition

$$\left(p \cdot \frac{\partial}{\partial u} \right) \Phi = 0, \quad (52)$$

1018 which reads in components as

$$p^{\mu_1} \Phi_{\mu_1 \mu_2 \dots \mu_s}(p) = 0. \quad (53)$$

1019 The non-vanishing components of a solution of (53) must be along the spatial directions,
1020 *i.e.* only $\Phi_{i_1 \dots i_s}(p)$ may be $\neq 0$. This symmetric tensor field is traceless with respect to
1021 the spatial metric: $\delta^{i_1 i_2} \Phi_{i_1 i_2 i_3 \dots i_s}(p) = 0$, thus the physical components carry a symmetric
1022 irrep of the orthogonal group $O(D-1)$, the dimension of which can be computed by making
1023 use of the formula (25). The polynomial wave function $\Phi(p, u)$ evaluated on the internal
1024 unit hypersphere $u^i u_i = 1$ corresponds to a decomposition of the physical components in
1025 terms of the spherical harmonics on the internal hypersphere S^{D-2} , which is an equivalent,
1026 though rather unusual, way of representing the physical components (usually, the use of
1027 spherical harmonics is reserved to the “orbital” part of the wave function).

1028 The quartic Casimir operator of the Poincaré algebra is easily evaluated in components
1029 in the rest frame

$$\begin{aligned} & -\frac{1}{2} P^2 M_{\mu\nu} M^{\mu\nu} + M_{\mu\rho} P^\rho M^{\mu\sigma} P_\sigma \\ & = \frac{1}{2} m^2 (M_{ij} M^{ij} + 2M_{i0} M^{i0}) - m^2 M_{i0} M^{i0} = m^2 \frac{1}{2} M_{ij} M^{ij}, \end{aligned}$$

1030 giving as a final result for a massive representation associated with a Young diagram λ

$$\begin{aligned} \mathcal{C}_4(\mathfrak{iso}(D-1, 1)) &= \mathcal{C}_2(\mathfrak{iso}(D-1, 1)) \mathcal{C}_2(\mathfrak{so}(D-1)), \\ &= m^2 \sum_{a=1}^r \lambda_a (\lambda_a + D - 2a - 1), \end{aligned} \quad (54)$$

1031 where the eigenvalues of the quadratic Casimir operator of the rotation algebra are given
1032 by the formula (27).

1033 **Example:** In any dimension D , the eigenvalue of the quartic Casimir operator for a
1034 massive symmetric representation of rank s is equal to $m^2 s(s+D-3)$. In four spacetime
1035 dimensions, the square of the Pauli-Lubanski vector acting on a massive field of spin- s is
1036 indeed equal to $m^2 s(s+1)$.

1037 Each massive representation in $D \geq 4$ dimensions may actually be obtained as the
1038 first Kaluza–Klein mode in a dimensional reduction from $D+1$ down to D dimensions.
1039 There is no loss of generality because the massive little group $SO(D-1)$ in D dimension is
1040 identified with the $(D+1)$ -dimensional helicity (short) little group. Such a Kaluza–Klein
1041 mechanism leads to a Stückelberg formulation of the massive field.

The massless limit $m \rightarrow 0$ of a massive irrep with λ fixed is, in general, reducible
because the irrep of the massive little group $SO(D-1)$ is restricted to the helicity (short)
little group $SO(D-2) \subset SO(D-1)$. This argument combined with the known branching

rule for $O(D - 1) \downarrow O(D - 2)$ (reviewed in Subsection 4.3) allows to prove that the massless limit of a massive irrep of the homogeneous Lorentz group labeled by a fixed Young diagram λ contains each helicity irrep labeled by Young diagrams μ such that

$$\lambda_1 \geq \mu_1 \geq \lambda_2 \geq \mu_2 \geq \dots \geq \mu_{r-1} \geq \lambda_r \geq \mu_r \geq 0,$$

1042 with multiplicity one. The zero modes of a dimensional reduction from $D + 1$ down to D
 1043 dimensions are determined by the same rule.

1044 **Example:** The zero modes of the dimensional reduction of a massive symmetric repre-
 1045 sentations with “spin” equal to s are all helicity symmetric representations with integer
 1046 “spins” not greater than the integer s , each with multiplicity one. For the dimensional
 1047 reduction of a gravitational theory (*i.e.* a spin-two particle), one recovers the usual result
 1048 that the massless spectrum is made of one “graviton” (spin-2), one “photon” (spin-1) and
 1049 one “dilaton” (spin-0).

1050 5.3 Massless representations

1051 The quartic Casimir operator of the Poincaré algebra is evaluated easily in components in
 1052 the light-cone coordinates (see Subsection 3.2 for notations),

$$-\frac{1}{2} P^2 M_{\mu\nu} M^{\mu\nu} + M_{\mu\rho} P^\rho M^{\mu\sigma} P_\sigma = 0 + M_{m+} P^+ M^{m-} P_- = \pi_m \pi^m,$$

1053 giving as a final result for a massless representation

$$C_4(\mathfrak{iso}(D - 1, 1)) = C_2(\mathfrak{iso}(D - 2)) = \mu^2 \tag{55}$$

1054 where the quadratic Casimir operator of the massless little group is written in (17).

1055 5.3.1 Helicity representations

1056 Helicity representations correspond to the case $\mu = 0$, so that $\pi^m = 0$ and in practice the
 1057 representation is induced from a representation of the orthogonal group $O(D - 2)$.

1058 **Step 1.** Again, any unitary representation of the orthogonal group $O(D - 2)$ is a
 1059 sum of finite-dimensional UIRs. Let us consider the UIR of the helicity short little group
 1060 $O(D - 2)$ labeled by the allowed Young diagram $\lambda = \{\lambda_1, \lambda_2, \dots, \lambda_r\}$ (that is, the sum of
 1061 the lengths of its first two columns does not exceed $D - 2$):

$$\lambda = \begin{array}{l} \boxed{} \lambda_1 \\ \boxed{} \lambda_2 \\ \boxed{} \lambda_3 \\ \dots \vdots \\ \boxed{\phantom{\lambda_{r-1}}} \lambda_{r-1} \\ \boxed{} \lambda_r \end{array} . \tag{56}$$

1062 The step 2 is more subtle to perform than for massive representations because the
 1063 field equations must set to zero all components along the light-cone of the covariant wave
 1064 function, because they are unphysical. In other words, the covariant wave equations should
 1065 remove *two* directions, and not only one like in the massive case. This fact implies that
 1066 the transversality is not a sufficient condition any more, it must be supplemented either
 1067 by other equations or by gauge symmetries asserting that one may quotient the solution
 1068 space by pure gauge fields. In these lecture notes, one focuses on two *gauge-invariant*
 1069 formulations which may be respectively referred to as “Bargmann-Wigner formulation”
 1070 in terms of the field strength and “gauge-fixed formulation” in terms of the potential.

1071 *Bargmann-Wigner equations*

1072 The so-called “Bargmann-Wigner equations” were actually first written by Dirac [27]
 1073 in four-dimensional Minkowski spacetime in spinorial form. Their name originates from
 1074 their decisive use in the completion of the Bargmann-Wigner programme [4]. The gener-
 1075 alization of the Bargmann-Wigner equations to any dimension was presented in [21] for
 1076 tensorial irreps (reviewed here) and in [22] for spinorial irreps. The latter programme had
 1077 previously been completed in [24] with different equations.

1078 **Step 2.** Let $\bar{\lambda} = \{\lambda_1, \lambda_1, \lambda_2, \dots, \lambda_r\}$ be the Young diagram depicted as

$$\bar{\lambda} = \begin{array}{c} \boxed{} \\ \boxed{} \\ \boxed{} \\ \boxed{} \\ \vdots \\ \boxed{} \\ \boxed{} \end{array} \begin{array}{l} \lambda_1 \\ \lambda_1 \\ \lambda_2 \\ \lambda_3 \\ \vdots \\ \lambda_{r-1} \\ \lambda_r \end{array} . \quad (57)$$

1079 It is obtained from the Young diagram λ represented in (56) by adding a row of equal
 1080 length on top of the first row of λ . The Young diagram $\bar{\lambda}$ has at least two rows of equal
 1081 lengths and the sum of the lengths of its first two columns does not exceed D . The
 1082 covariant wave function is chosen to take values in the Schur module $V_{\bar{\lambda}}^{O(D-1,1)}$ realized
 1083 in the manifestly antisymmetric convention. Following Subsection 4.4, the wave function
 1084 $\mathcal{K}(p, d_I x)$ is taken to be a polynomial in the graded variables $d_I x^\mu$ ($I = 1, 2, \dots, \lambda_1$)
 1085 obeying the commutation relations (30). Moreover, the irreducibility conditions of the
 1086 components under the Lorentz group $O(D-1, 1)$ are

$$\left(d_I x^\mu \frac{\partial^L}{\partial(d_I x^\mu)} - \delta_{IJ} \bar{\ell}_I \right) \mathcal{K} = 0, \quad (I \leq J) \quad (58)$$

1087 where $\bar{\ell}_I$ stands for the length of the I th column in the Young diagram $\bar{\lambda}$, and

$$\left(\eta^{\mu\nu} \frac{\partial^L}{\partial(d_I x^\mu)} \frac{\partial^L}{\partial(d_I x^\nu)} \right) \mathcal{K} = 0. \quad (59)$$

1088 **Step 3.** The covariant field equations may be summarized in the assertion that the
 1089 wave function is a “harmonic” multiform in the sense that, $\forall I$, it is “closed”

$$\left(p_\mu d_I x^\mu \right) \mathcal{K} = 0, \quad (60)$$

1090 and “coclosed” (*i.e.* transverse)

$$\left(p^\mu \frac{\partial^L}{\partial(d_I x^\mu)} \right) \mathcal{K} = 0. \quad (61)$$

1091 The operators $p \cdot d_I x$ act as “exterior differentials” (or “curls”), they are nilpotent and obey
 1092 graded commutation relations. As one can easily see, the field equations (60) and (61),
 1093 considered together, imply the massless Klein-Gordon equation. Actually, the equations
 1094 (60) may even be imposed off-shell, whereas the equations (61) only hold on-shell [21].

1095 **Step 4.** In the light-cone frame (see Section 1.1), the components of the momentum
 1096 may be taken to be $p_\mu = (p_-, 0, 0, \dots, 0)$ with $p_- \neq 0$. On the one hand, the transversality
 1097 condition (61) implies that the wave function does not depend on the variables $d_I x^+$. On
 1098 the other hand, the closure condition (60) reads $(p_- d_I x^-) \mathcal{K} = 0$, the general solution of
 1099 which is $\mathcal{K} = (\prod_I p_- d_I x^-) \phi$, where ϕ depends neither on $d_I x^-$ nor on $d_I x^+$ (due to the

1100 transversality condition). In other words, the directions along the light-cone have been
 1101 removed, since $\phi = \phi(p, d_I x^m)$ ($m = 1, 2, \dots, D - 2$). Focusing on this field, one may
 1102 show that the irreducibility conditions (58) become, in terms of the function ϕ ,

$$\left(d_I x^m \frac{\partial^L}{\partial (d_J x^m)} - \delta_{IJ} \ell_I \right) \phi = 0, \quad (I \leq J) \quad (62)$$

1103 where $\ell_I = \bar{\ell}_I - 1$, and the trace conditions (59) implies

$$\left(\delta^{mn} \frac{\partial^L}{\partial (d_I x^m)} \frac{\partial^L}{\partial (d_J x^n)} \right) \phi = 0. \quad (63)$$

1104 Since ℓ_I is the length of the I th column of the Young diagram λ , the system of equations
 1105 (62)-(63) states that the components of the function ϕ carry a tensorial irrep of the or-
 1106 thogonal group $O(D - 2)$. Therefore, the same is true for the physical components of the
 1107 wave function \mathcal{K} . \square

1108 This may be reformulated covariantly by saying that the closure (60) of the wave
 1109 function implies that

$$\mathcal{K} = \left(\prod_{I=1}^{\lambda_1} p_\mu d_I x^\mu \right) \phi. \quad (64)$$

1110 In components, this means that the tensor \mathcal{K} is equal to λ_1 curls of the tensor ϕ . This
 1111 motivates the name “field strength” for the wave function $\mathcal{K}(p, d_I x)$, the components of
 1112 which are irreducible under the Lorentz group (when evaluated on zero-mass shell) and
 1113 labeled by $\bar{\lambda}$, and the name “potential” or “gauge field” for the wave function $\phi(p, d_I x)$,
 1114 the components of which may be taken to be irreducible under the general linear group,
 1115 with symmetries labeled by the Young diagram λ .

1116 Examples:

- The helicity vectorial representation corresponds to a Young diagram $\lambda = \{1\}$ made of a single box. In four spacetime dimensions, this representation is precisely what is usually called a “vector gauge field”. The Young diagram $\bar{\lambda} = \{1, 1\}$ is a single column made of two boxes. The wave function in momentum space is given by

$$\mathcal{K} = \frac{1}{2} \mathcal{K}_{\mu\nu}(p) dx^\mu dx^\nu$$

which carries an irrep of $GL(D, \mathbb{R})$: the antisymmetric rank-two representation. As one can see, the wave function actually is a differential two-form, the components of which transforming as an antisymmetric tensor of rank two. The field equations (60) and (61), respectively, read in components

$$p_\mu \mathcal{K}_{\nu\rho} + p_\nu \mathcal{K}_{\rho\mu} + p_\rho \mathcal{K}_{\mu\nu} = 0 \quad (\text{Bianchi identities})$$

and

$$p^\mu \mathcal{K}_{\mu\nu} = 0 \quad (\text{transversality conditions}).$$

1117 The differential two-form \mathcal{K} is indeed harmonic (closed and coclosed). In physical terms,
 1118 one says that the field strength $\mathcal{K}_{\mu\nu}$ obeys to the Maxwell equations. As usual, the Bianchi
 1119 identities imply that the field strength derives from a potential: $\mathcal{K}_{\mu\nu} = p_\mu \phi_\nu - p_\nu \phi_\mu$. In
 1120 the light-cone coordinates, the transversality implies that the components $\mathcal{K}_{+\nu}$ vanish,
 1121 thus the only non-vanishing components are $\mathcal{K}_{-n} = p_- \phi_n$. Therefore the only physical
 1122 components correspond to a $(D - 2)$ -vector in the hyperplane transverse to the light-cone.

- Helicity symmetric representations with “helicity” (or “spin”) equal to s correspond to Young diagrams $\lambda = \{s\}$ made of one row of length equal to the integer s . In four

spacetime dimensions, this representation is precisely what is usually called a “massless spin- s field”. The Young diagram $\bar{\lambda} = \{s, s\}$ is a rectangle made of two row of length equal to the integer s . The wave function is thus a polynomial in the auxiliary variables

$$\mathcal{K} = \frac{1}{2^s} \mathcal{K}_{\mu_1\nu_1|\dots|\mu_s\nu_s} d_1x^{\mu_1} d_1x^{\nu_1} \dots d_sx^{\mu_s} d_sx^{\nu_s}$$

1123 satisfying the irreducibility equations (58)-(59) with $\ell_I = 2, \forall I \in \{1, \dots, s\}$. The tensor
1124 \mathcal{K} is, by construction, antisymmetric in each of the s sets of two indices

$$\mathcal{K}_{\mu_1\nu_1|\dots|\mu_s\nu_s} = -\mathcal{K}_{\nu_1\mu_1|\dots|\mu_s\nu_s} = \dots = -\mathcal{K}_{\mu_1\nu_1|\dots|\nu_s\mu_s}. \quad (65)$$

1125 Moreover, the complete antisymmetrization over any set of three indices gives zero and
1126 all its traces are zero on-shell, so that the on-shell tensor \mathcal{K} indeed belongs to the space
1127 irreducible under the Lorentz group $O(D-1, 1)$ characterized by a two-row rectangular
1128 Young diagram of length s . In four-dimensional Minkowski spacetime, the irrep of the
1129 Lorentz group $O(3, 1)$ carried by the on-shell tensor \mathcal{K} is usually denoted as $(s, 0) \oplus (0, s)$.
1130 More precisely, the symmetry properties of the tensor $\mathcal{K}_{\mu_1\nu_1|\dots|\mu_s\nu_s}$ are labeled by the
1131 Young tableau

$$\begin{array}{|c|c|c|c|} \hline \mu_1 & \mu_2 & \dots & \mu_s \\ \hline \nu_1 & \nu_2 & \dots & \nu_s \\ \hline \end{array} .$$

1132 The equation (64) means that the components of the tensor $\mathcal{K}_{\mu_1\nu_1|\dots|\mu_s\nu_s}$ are essentially
1133 the projection of $p_{\mu_1} \dots p_{\mu_s} \phi_{\nu_1 \dots \nu_s}$ on the tensor field irreducible under $GL(D, \mathbb{R})$ with
1134 symmetries labeled by the above Young tableau. The physical components $\phi_{n_1 \dots n_s}$ of the
1135 symmetric tensor gauge potential $\phi_{\nu_1 \dots \nu_s}$ are along the $D-2$ directions transverse to the
1136 light-cone. The number of physical degrees of freedom of a helicity symmetric field of rank
1137 s can be computed by making use of the formula (25).

1138 • The helicity symmetric representation with “spin” equal to 2 corresponds to the gravi-
1139 ton. The field strength has the symmetry properties of the Riemann tensor. Its on-shell
1140 tracelessness indicates that it corresponds to the (linearized) Weyl tensor. The equations
1141 (60) are the Bianchi identities for the linearized Riemann tensor in flat spacetime, whereas
1142 the equations (61) hold as a consequence of the sourceless Einstein equations linearized
1143 around flat spacetime.

1144 Remark:

1145 One can find some early indications for the existence of the tensor $\mathcal{K}^{\mu_1\nu_1|\dots|\mu_s\nu_s}$ in the
1146 paper [28] where Weinberg constructs free quantum field operators that have a nonzero
1147 expectation value between the vacuum and one-particle states for massless particles of
1148 helicity $\pm s$ in four spacetime dimensions. In Weinberg’s approach, one cannot find the
1149 classical (or “first-quantized”) field strength tensor $\mathcal{K}^{\mu_1\nu_1|\dots|\mu_s\nu_s}$ that we have built above,
1150 but instead a quantum *operator* (in so-called “second-quantization”) that we denote here
1151 $\hat{\mathcal{K}}_{\pm}^{\mu_1\nu_1|\dots|\mu_s\nu_s}$ and that transforms like a tensor under Lorentz transformations. This
1152 operator is built out of the product $[p^{\mu_1} e_{\pm}^{\nu_1}(\vec{p}) - p^{\nu_1} e_{\pm}^{\mu_1}(\vec{p})] \dots [p^{\mu_s} e_{\pm}^{\nu_s}(\vec{p}) - p^{\nu_s} e_{\pm}^{\mu_s}(\vec{p})]$
1153 featuring the two polarisation “vectors” $e_{\pm}^{\mu}(\vec{p})$. On the one hand, solving the Bianchi
1154 identities for the field strength $\mathcal{K}^{\mu_1\nu_1|\dots|\mu_s\nu_s}$ allows to write the latter as an expression
1155 involving s derivatives of a completely symmetric gauge potential $\phi_{\mu_1 \dots \mu_s}$. This potential
1156 satisfies [21] the second-order Fronsdal field equations [29] and is the building block for
1157 the construction of an interacting quantum field theory with long-range interactions. On
1158 the other hand, the canonical quantization of the free field theory with field strength
1159 tensor \mathcal{K} gives rise to Weinberg’s quantum field operator $\hat{\mathcal{K}}_{\pm}$. The same remarks apply to

1160 the relation between the generalised field strength (64) and its second-quantized version
 1161 in [30].

1162 *Gauge-fixed equations*

1163 The following equations are somewhat unusual, but they proved to be crucial in the
 1164 completion of the Bargmann-Wigner programme for the infinite spin representations [17].

1165 **Step 2.** Let $\widehat{\lambda} = \{\lambda_1 - 1, \lambda_2 - 1, \dots, \lambda_r - 1\}$ be the Young diagram depicted as

$$\widehat{\lambda} = \begin{array}{c} \boxed{} \lambda_1 - 1 \\ \boxed{} \lambda_2 - 1 \\ \boxed{} \lambda_3 - 1 \\ \dots \vdots \\ \boxed{} \lambda_{r-1} - 1 \\ \boxed{} \lambda_r - 1 \end{array}, \quad (66)$$

1166 obtained from the Young diagram λ represented in (56) by removing the first column
 1167 of λ . Therefore the sum of the length of the first two columns of the Young diagram
 1168 $\widehat{\lambda}$ does not exceed $D - 2$. The covariant wave function is chosen to take values in the
 1169 Schur module $V_{\widehat{\lambda}}^{O(D-1,1)}$ realized in the manifestly symmetric convention. Actually, as
 1170 anticipated in Subsection 4.4, it turns out to be crucial to regard the wave function $\Phi(p, u_a)$
 1171 as a *distribution* in the commuting auxiliary variables u_a^μ , obeying to

$$\left[\left(u_a \cdot \frac{\partial}{\partial u_b} \right) - \widehat{\lambda}_a \delta_{ab} \right] \Phi = 0, \quad (a \leq b). \quad (67)$$

$$\left(\frac{\partial}{\partial u_a} \cdot \frac{\partial}{\partial u_b} \right) \Phi = 0, \quad (68)$$

1172 **Step 3.** Proper field equations are the transversality condition (49) combined with
 1173 the equation

$$(p \cdot u_a) \Phi = 0. \quad (69)$$

1174 The equations (69) and (49) are the respective analogues of the closure and coclosure
 1175 conditions (60)-(61). A drastic difference is that the operators $p \cdot u_a$ are not nilpotent
 1176 (thus there is no underlying cohomology). Actually, the equation (69) has no solution if
 1177 Φ is assumed to be a polynomial in all the variables.

1178 **Step 4.** Equation (69) can be solved as

$$\Phi = \delta(u_a \cdot p) \phi, \quad (70)$$

1179 where the distribution $\phi(p, u_a)$ may actually be assumed to be a function depending poly-
 1180 nomially on the auxiliary variables u_a for the present purpose. The Dirac delta is a
 1181 distribution of homogeneity degree equal to minus one, hence the irreducibility conditions
 1182 (67)-(68) imply that

$$\left[\left(u_a \cdot \frac{\partial}{\partial u_b} \right) - \lambda_a \delta_{ab} \right] \phi = 0 \quad (a \leq b), \quad (71)$$

$$\left(\frac{\partial}{\partial u_a} \cdot \frac{\partial}{\partial u_b} \right) \phi = 0. \quad (72)$$

1183 The function ϕ is defined from (70) modulo the equivalence relation

$$\phi \sim \phi + \sum_{a=1}^r (u_a \cdot p) \epsilon_a \quad (73)$$

1184 where ϵ_a are arbitrary functions. This means that (70) is equivalent to the alternative
 1185 road towards the Bargmann-Wigner programme: the gauge symmetry principle with the
 1186 irreducible components of $(u_a \cdot p) \epsilon_a$ being pure gauge fields. As mentioned before, this
 1187 path will not be addressed here (see *e.g.* [21] and refs therein for more discussions on the
 1188 gauge-invariance issue). Therefore, one may say that the equation (69) is the “remnant”
 1189 of the gauge symmetries (73). In the light-cone coordinates, the gauge symmetries (73)
 1190 imply that one may choose a representative ϕ which does not depend on the variables u_a^-
 1191 (the gauge is “fixed”). The transversality condition (49) implies that ϕ is also transverse,
 1192 implying no dependence on u_a^+ (“gauge shoots twice”). Thus ϕ depends only on the trans-
 1193 verse auxiliary variables u_a^m , so one concludes by observing that the physical components
 1194 of ϕ carry a tensorial irrep of $O(D-2)$ labeled by λ . \square

1195 5.3.2 Infinite spin representations

1196 Infinite spin representations correspond to the case $\mu \neq 0$ and, in practice, the repre-
 1197 sentation of the massless little group $IO(D-2)$ is induced from a representation of the
 1198 orthogonal group $O(D-3)$. The parameter μ is a real parameter with the dimension of a
 1199 mass. Wigner proposed a set of manifestly covariant equations to describe fields carrying
 1200 these UIR in four spacetime dimensions [31]. They have been generalized to arbitrary
 1201 infinite-spin representations in any dimension [17].⁹

1202 **Step 1.** Again, any unitary representation of the orthogonal group $O(D-3)$ is a
 1203 sum of finite-dimensional UIRs. Let us consider the UIR of the helicity short little group
 1204 $O(D-3)$ labeled by the allowed Young diagram $\lambda = \{\lambda_1, \lambda_2, \dots, \lambda_r\}$ (that is, the sum of
 1205 the lengths of its first two columns does not exceed $D-3$).

1206 **Step 2.** In order to have manifest covariance, it is necessary to lift the eigenvalues
 1207 ξ^m of the generators π^m in the massless little group to a D -vector ξ^μ . In practice, the
 1208 covariant wave function is taken to be a distribution $\Phi(p, \xi, u_a)$ satisfying the conditions
 1209 (33)-(34). The tensorial components associated with the commuting variables u_a belong to
 1210 the Schur module of the Lorentz group $O(D-1, 1)$ labeled by an allowed Young diagram
 1211 λ .

1212 **Step 3.** Relativistic equations describing a first-quantized particle with infinite spin
 1213 are

$$(p \cdot \xi) \Phi = 0, \quad (74)$$

$$\left(p \cdot \frac{\partial}{\partial \xi} - i \right) \Phi = 0, \quad (75)$$

$$(\xi^2 - \mu^2) \Phi = 0, \quad (76)$$

1214 together with the transversality conditions

$$(p \cdot u_a) \Phi = 0, \quad (77)$$

$$\left(p \cdot \frac{\partial}{\partial u_a} \right) \Phi = 0, \quad (78)$$

$$\left(\xi \cdot \frac{\partial}{\partial u_a} \right) \Phi = 0. \quad (79)$$

1215 This system of equations is far from being independent. For instance, compatibility con-
 1216 dition of the systems (74)-(75) or (77)-(78) is the massless Klein-Gordon equation.

⁹More recent developments (as well as a list of open challenges) have been reviewed in [18].

1217 **Step 4.** The equation (75) reflects the fact that the couples (p, ξ) and $(p, \xi + \alpha p)$ are
 1218 physically equivalent for arbitrary $\alpha \in \mathbb{R}$. Indeed, one gets

$$\Phi(p, \xi + \alpha p) = e^{i\alpha} \Phi(p, \xi) \quad (80)$$

1219 from Equation (75). The equation (76) states that the internal vector ξ is a space-like
 1220 vector while the mass-shell condition states that the momentum is light-like. From the
 1221 equation (74), one obtains that the internal vector is transverse to the momentum. All
 1222 together, one finds that ξ may be taken to live on the hypersphere S^{D-3} of radius μ
 1223 embedded in the transverse hyperplane \mathbb{R}^{D-2} . In brief, the “continuous spin” degrees
 1224 of freedom essentially correspond to $D - 3$ angular variables, whose Fourier conjugates
 1225 are discrete variables analogous to the usual spin degrees of freedom. Finally, proceeding
 1226 analogously to the “gauge-fixed” field equations of the helicity representations, one may
 1227 show [17] that the conditions (77)-(79) concretely remove three unphysical directions in
 1228 the components, so that the final result is a tensorial irrep of the short little group $O(D-3)$
 1229 fixing both the momentum p and the internal vector ξ .

1230 From the group theoretical point of view, the UIR of the homogeneous and inho-
 1231 mogeneous orthogonal groups are related by an Inönü-Wigner contraction $O(D-1) \rightarrow$
 1232 $IO(D-2)$ (see Subsection 4.5). It follows that one can obtain the continuous spin rep-
 1233 resentations from the massive ones in a suitable massless limit $m \rightarrow 0$ since their little
 1234 group UIRs are related by a contraction. The quartic Casimir operator of the Poincaré
 1235 group for the massive representation is related to its Young diagram ν labeling the UIR
 1236 of the little group $O(D-1)$ via the formula (54):

$$\mathcal{C}_4(\mathfrak{iso}(D-1, 1)) = m^2 \sum_{a=1}^r \nu_a (\nu_a + D - 2a - 1), \quad (81)$$

1237 In order to keep \mathcal{C}_4 non-vanishing, the massless limit must be such that the product of the
 1238 “spin” $\nu_1 = s$ and the mass m remains finite. More precisely, one needs $sm \rightarrow \mu$ in order
 1239 to reproduce (55), so that the spin goes to infinity while the row lengths ν_a for $a \neq 1$ are
 1240 kept equal to λ_{a-1} [17, 32]. The Fourier transform (in the internal space spanned by ξ)
 1241 of the field equations (74)-(79) may be obtained in this way from the field equations of a
 1242 massive representation in “gauge-fixed” form (see [17] for more details). This limit is very
 1243 similar to the contraction of Subsection 4.5.

1244 5.4 Tachyonic representations

1245 The tachyonic representations have some similarities with the massive representations.
 1246 The simpler one is the analogue of the Klein-Gordon equation, up to a change of sign
 1247 for the mass term. The other similarity is that the linear equations should remove the
 1248 components along the momentum. Of course, the major difference is that the momentum
 1249 is space-like. The quartic Casimir operator of the Poincaré algebra is also evaluated easily
 1250 in components, giving as a final result for a tachyonic representation,

$$\mathcal{C}_4(\mathfrak{iso}(D-1, 1)) = \mathcal{C}_2(\mathfrak{iso}(D-1, 1)) \mathcal{C}_2(\mathfrak{so}(D-2, 1)), \quad (82)$$

1251 where the eigenvalues of the quadratic Casimir operator of the rotation algebra are given
 1252 by the formula (27).

1253 **Step 1.** The first step is more involved for the tachyonic case since it requires the
 1254 exhaustive knowledge of the UIR theory for the groups $SO(D-2, 1)^\dagger$. Fortunately,
 1255 complete results are available [33, 34]. The steps 2-3 further require the completion of

1256 the Bargmann-Wigner programme for the isometry group $SO(D-2,1)^\uparrow$ of the de Sitter
1257 spacetime dS_{D-2} . This has been done in [23].¹⁰

1258 Let us assume that this programme has been performed through an ambient space
1259 formulation, analogous to the one of the spherical harmonics, as discussed in the subsection
1260 4.4. More explicitly, let us consider that the physical components of the wave function
1261 have been realized via a function on the hyperboloid dS_{D-2} of radius $\mu > 0$ embedded
1262 in $\mathbb{R}^{D-2,1}$ with some set of auxiliary commuting vectors of $\mathbb{R}^{D-2,1}$ (for the spin degrees
1263 of freedom) and the corresponding $O(D-2,1)$ -covariant field equations of the UIR are
1264 known explicitly. The step 1 is therefore assumed to be performed.

1265 **Step 2.** In order to have manifest Lorentz invariance, all auxiliary variables are lifted
1266 to D -vectors: the coordinates of the internal de Sitter spacetime are denoted by ξ^μ and
1267 the auxiliary variables by u_A^μ . The wave function is taken to be $\Phi(p, \xi, u_A)$, where the
1268 internal vector ξ plays a role similar to the one in the infinite-spin representations. An
1269 important distinction is that in the ambient space formulation, one would evaluate the
1270 wave function on the hypersurface $\xi^2 = \mu^2$ instead of imposing this relation on the wave
1271 function, as in (76). The $O(D-2,1)$ -covariant field equations for the UIR of the little
1272 group $O(D-2,1)$ must be $O(D-1,1)$ -covariantized accordingly. Concretely, this implies
1273 that the components of the covariant wave function carry an (infinite-dimensional) irrep
1274 of the Lorentz group.

1275 **Step 3.** These covariantized field equations and the tachyonic Klein-Gordon equation
1276 $(p^2 - m^2)\psi = 0$ must be supplemented by two equations: say the orthogonality condition
1277 (74), similarly to the infinite spin representation, and the transversality condition (49),
1278 similarly to the massive representation. The orthogonality condition (74) may be replaced
1279 by another transversality equation for the vector ξ .

1280 **Step 4.** Now, the equation (74) implies that the internal vector belongs to the hy-
1281 perplane $\mathbb{R}^{D-2,1}$ orthogonal to the momentum p . Its intersection with the hypersurface
1282 $\xi^2 = \mu^2$ restricts ξ to the internal de Sitter space $dS_{D-2} \subset \mathbb{R}^{D-2,1}$. Moreover, the condi-
1283 tion (49) sets to zero all components of the wave function along the momentum. Therefore,
1284 the remaining components are physical and carry an UIR of the little group $O(D-2,1)$
1285 by construction (see step 2). \square

1286 **Example:** The simplest non-trivial example corresponds to a tachyonic representation
1287 of the inhomogeneous Lorentz group $IO(D-1,1)$ induced by a representation of the
1288 little group $O(D-2,1)$ corresponding to “massive scalar field” on the “internal de Sitter
1289 spacetime” dS_{D-2} with $D \geq 4$. This UIR belongs to the *principal continuous series* of
1290 UIR of the group $O(D-2,1)$ and it may be realized as the space of harmonic functions
1291 on $\mathbb{R}^{D-2,1}$ of (complex) homogeneity degree s equal to $\frac{3-D}{2} + i\sigma$ (with σ a positive real
1292 parameter) evaluated on the unit one-sheeted hyperboloid $dS_{D-2} \subset \mathbb{R}^{D-2,1}$. They can be
1293 regarded as a generalization of the spherical harmonics in the Lorentzian case, where the
1294 degree is a complex number. The eigenvalue of the quadratic Casimir operator (4) of the
1295 little group $O(D-2,1)$ on this representation is equal to

$$C_2(\mathfrak{so}(D-2,1)) = \left(\frac{D-3}{2}\right)^2 + \sigma^2. \quad (83)$$

1296 The d’Alembertian on the unit hyperboloid evaluated on such functions is precisely equal
1297 to the former eigenvalue (as is true for the Laplacian on the unit sphere evaluated on spher-
1298 ical harmonics) so the corresponding fields on the internal spacetime dS_{D-2} are indeed
1299 “massive”. Inserting the above result in (82), one sees that the quartic Casimir operator is

¹⁰The Bargmann-Wigner programme in AdS_D , with field equations that generalise the ones presented in Section 5.3, were obtained in [35]. Similar equations were obtained later in the dS_D signature [23].

1300 negative for the corresponding tachyonic representation. In four-dimensional Minkowski
 1301 spacetime, this implies that the Pauli-Lubanski vector is time-like. The Lorentz-covariant
 1302 wave function is taken to be $\Phi(p, \xi)$ evaluated on $\xi^2 = 1$ and the corresponding relativistic
 1303 equations for the induced tachyonic representation may be chosen as

$$(p^2 - m^2) \Phi = 0, \tag{84}$$

$$\left(p \cdot \frac{\partial}{\partial \xi} \right) \Phi = 0, \tag{85}$$

$$\left(\frac{\partial}{\partial \xi} \cdot \frac{\partial}{\partial \xi} \right) \Phi = 0, \tag{86}$$

$$\left(\xi \cdot \frac{\partial}{\partial \xi} - s \right) \Phi = 0, \tag{87}$$

1304 where one should remember that $s = \frac{3-D}{2} + i\sigma$. Notice the formal analogy with the
 1305 system of equations (48), (52)) and (50) for a massive symmetric tensor field.

1306 **Remark:** There might be sometimes confusion in the folklore surrounding the tachyons.
 1307 We would like to insist on the fact that the tachyonic representations are indeed *unitary* (by
 1308 definition). Still, their physical interpretation is problematic because they are *not causal*
 1309 in the sense that one may show that the support of their propagator requires superluminal
 1310 propagation. Roughly speaking, the acausality is obvious because the momentum is space-
 1311 like, $p^2 = +m^2$. The confusing point is that one may try to circumvent this problem in
 1312 the following way: solving $p^2 - m^2 = 0$ by $p^\mu = (im, \vec{0})$ enforces causality, but the price to
 1313 pay is the loss of unitarity. Indeed, the energy is pure imaginary, hence a naive plane-wave
 1314 $e^{\pm i p_0 x^0}$ is actually a non-integrable exponential $e^{\pm m x^0}$. These remarks are summarized
 1315 in the following table:

$E = p_0$	$ \vec{p} $	Unitarity	Causality
0	m	OK	KO
$\pm im$	0	KO	OK

1317 Nevertheless, the tachyonic representations should not be discarded too quickly on
 1318 such physical grounds. Actually, if tachyonic representations appear in the spectrum of
 1319 a theory, then it merely signals a local instability of the field theory in the sense that
 1320 the perturbation theory is performed around an unstable vacuum, and the tachyon might
 1321 roll to a stable vacuum (if any). For instance, the Higgs particle is described by nothing
 1322 but a tachyonic scalar field (induced by the trivial representation of the little group).
 1323 By analogy, one may wonder if some infinite-component tachyonic field (induced by a
 1324 non-trivial representation of the little group) could not play a role in some huge Brout–
 1325 Englert–Higgs mechanism providing mass to an infinite tower of gauge fields in various
 1326 massless irreps.

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1334 A Siegel-Zwiebach equations

1335 The Bargmann-Wigner programme for finite-component representations in Minkowski
 1336 spacetime of any dimension $D > 3$ was completed for massless helicity representations by
 1337 Siegel and Zwiebach in [24] and generalised to massive representations in Siegel’s lecture
 1338 notes [20]. Only the massless representations will be reviewed here since the case of massive
 1339 representations follows by dimensional reduction, as mentioned in the subsection 5.2.

1340 *Siegel-Zwiebach equations*

1341 The main idea behind these equations is the covariantisation of the condition that the
 1342 “translation” generators π_n of the massless little group $IO(D-2)$ must act trivially on
 1343 physical states of the helicity representations (cf. Subsections 3.2-3.3). Let us rewind the
 1344 procedure initiated in Subsection 5.3.1:

1345 **Steps 1 and 2.** These first steps are identical to the case of Bargmann-Wigner
 1346 equations, *i.e.* the wave function is a field strength $\mathcal{K}(p, d_I x)$ taking values in an irrep of
 1347 the Lorentz group $O(D-1, 1)$ labeled by the Young diagram $\bar{\lambda}$.

1348 **Step 3.** The generators of the Lorentz algebra $\mathfrak{so}(D-1, 1)$ can be decomposed as the
 1349 sum $M_{\mu\nu} = L_{\mu\nu} + S_{\mu\nu}$ of the “orbital” part (transforming the positions or momenta) and
 1350 the “spin” part (transforming the irrep labeled by the Young diagram $\bar{\lambda}$),

$$1351 \quad L_{\mu\nu} = -i \left(p_\mu \frac{\partial}{\partial p^\nu} - p_\nu \frac{\partial}{\partial p^\mu} \right), \quad S_{\mu\nu} = -i \left(d_I x_\mu \frac{\partial}{\partial (d_I x^\nu)} - d_I x_\nu \frac{\partial}{\partial (d_I x^\mu)} \right). \quad (88)$$

1351 The Siegel-Zwiebach equations for $s \neq 0$ take the simple form

$$(p^\mu S_{\mu\nu} - i s p_\nu) \mathcal{K} = 0. \quad (89)$$

1352 They imply the massless Klein-Gordon equation $p^2 \mathcal{K} = 0$ (since $s \neq 0$). In fact, one can
 1353 check that the quadratic and quartic Casimir operators both vanish as a consequence of
 1354 (89).¹¹ Notice that a similar “spin-enslaving” relation, leading to (89), was recently given
 1355 in [36].

1356 **Step 4.** In the light-cone frame (see Section 1.1) where the components of the mo-
 1357 mentum are $p_\mu = (p_-, 0, 0, \dots, 0)$ with $p_- \neq 0$, the system (89) of equations splits as

$$1358 \quad \pi_n \mathcal{K} = 0, \quad (S_{+-} - i s) \mathcal{K} = 0, \quad (90)$$

1359 where $\pi_n := p_- S_{+n} = p^\mu S_{\mu n}$ (with $n = 1, 2, \dots, D-2$) are generators corresponding the
 1360 “translation” subgroup $\mathbb{R}^{D-2} \subset IO(D-2)$ of the massless little group.¹² On the one hand,
 1361 the fact that these generators π_n act trivially ensures that the massless representation is
 1362 a helicity representation, *i.e.* only the generators S_{mn} of the rotations in the transverse
 1363 plane act non-trivially. Moreover, the condition $\pi_n \mathcal{K} = 0$ implies that the field strength
 1364 \mathcal{K} in the light-cone frame has a maximal (respectively, minimal) number of factors $d_I x^-$
 1365 (respectively, $d_I x^+$).¹³ Therefore, the physical components of the field strength read
 1366 $\mathcal{K} = (\prod_I p_- d_I x^-) \phi$, where ϕ depends neither on $d_I x^-$ nor on $d_I x^+$. On the other hand,
 1367 the eigenvalue $S_{+-} = i s$ of the Lorentz generator

$$S_{+-} = -i \left(d_I x^+ \frac{\partial}{\partial (d_I x^+)} - d_I x^- \frac{\partial}{\partial (d_I x^-)} \right) \quad (91)$$

¹¹In order to check that the quartic Casimir operator acts trivially, it useful to notice that $M_{\mu\nu}$ can be replaced everywhere by $S_{\mu\nu}$ inside the definition (6). In $D = 4$, this property is obvious in terms of the Pauli-Lubanski vector.

¹²See Subsection 3.2. Note that $M_{+n} = S_{+n}$ and $M_{mn} = S_{mn}$ in this light-cone frame, since the corresponding orbital parts of the generators of the little group act trivially on the momentum.

¹³See [20] for an elegant derivation of these facts from (90).

1368 implies that the Young diagram $\bar{\lambda}$ must have s columns. This is because the operator
 1369 S_{+-} is a number operator (up to a coefficient i) for the total number of covariant indices
 1370 $-$ minus the number of covariant indices $+$, and in every column of the field strength
 1371 there is no index $+$ and one index $-$. The conclusion that is reached is the same as in
 1372 Subsection 5.3.1.

1373 *Equivalence with Bargmann-Wigner equations*

1374 In fact, the Siegel-Zwiebach equations are equivalent to the Bargmann-Wigner equa-
 1375 tions reviewed in Subsection 5.3.1. For instance, the closure and coclosure conditions (60)
 1376 and (61) imply (89). This follows from the identity

$$\begin{aligned} p^\mu S_{\mu\nu} &= -i p^\mu \left(d_I x_\mu \frac{\partial}{\partial(dx_I^\nu)} - d_I x_\nu \frac{\partial}{\partial(dx_I^\mu)} \right) \\ &= -i \left(p_\mu d_I x^\mu \right) \frac{\partial}{\partial(dx_I^\nu)} - d_I x_\nu \left(p^\mu \frac{\partial}{\partial(dx_I^\mu)} \right). \end{aligned} \quad (92)$$

1377 In the last term, one recognises between the parentheses the divergence operator acting
 1378 on the column I , which gives zero because of the coclosure condition (61). As for the first
 1379 term on the right-hand side of the above equation, one can rewrite it as

$$-i \left(p_\mu d_I x^\mu \right) \frac{\partial}{\partial(dx_I^\nu)} = -i \frac{\partial}{\partial(dx_I^\nu)} \left(p_\mu d_I x^\mu \right) - i p_\mu \left[dx_I^\mu, \frac{\partial}{\partial(dx_I^\nu)} \right]. \quad (93)$$

1380 The first term on the right-hand side gives zero on the field strength because of the
 1381 closure relation (60), while the last term gives $+i s p_\nu$ because of the commutation relations
 1382 $\left[dx_I^\mu, \frac{\partial}{\partial(dx_I^\nu)} \right] = -s \delta_\nu^\mu$.

1383 The covariant proof that the Siegel-Zwiebach equations imply Bargmann-Wigner
 1384 equations is more cumbersome and will not be presented here. Anyway, this equivalence
 1385 is guaranteed from the light-cone frame analysis.

1386 B Bargmann-Wigner programme in three dimensions

1387 In this appendix we review results obtained in the literature concerning the Wigner and
 1388 Bargmann-Wigner programmes in Minkowski spacetime of dimension $D = 2 + 1$. The
 1389 former programme was achieved in [10] along the lines of the seminal paper [6] by Wigner.

1390 There are four classes of UIRs of the Poincaré group $ISO(2, 1)^\uparrow$:

- 1391 1) Zero-momentum representations, labeled by the eigenvalue $c \in \mathbb{R}$ of the quadratic
 1392 Casimir operator $\mathcal{C}_2[\mathfrak{so}(2, 1)]$ of the Lorentz algebra;¹⁴
- 1393 2) Massive representations, labeled by mass $m > 0$ and spin $s \in \mathbb{R}$;
- 1394 3) Massless representations:
 - 1395 1. helicity representations, either single-valued (bosonic) or double-valued (fermionic);
 - 1396 2. infinite-spin representations, labeled by a dimensionful parameter $\mu > 0$;
- 1397 4) Tachyonic representations, labeled by a dimensionful parameter $m > 0$ and by a
 1398 dimensionless parameter $s \in \mathbb{R}$ (the analogue of spin).

1399 In what follows, we briefly summarize exhaustive results on the completion of the Bargmann-
 1400 Wigner programme in $D = 2 + 1$ dimensions for the four classes of UIRs listed above.

¹⁴Strictly speaking, the principal and complementary series are labeled by two real parameters, not only by the value of the Casimir operator.

1401 B.1 Zero momentum representations

1402 Effectively, the zero momentum representation of the Poincaré group $ISO(2, 1)^\uparrow$ are UIRs
 1403 of the Lorentz subgroup $SO(2, 1)^\uparrow$. The latter were classified in [33]. We also refer
 1404 the reader to [37] for a physicist-friendly classification of the irreps of the Lorentz group
 1405 $SO(2, 1)^\uparrow$.

1406 We will not repeat these well-known results here. For the purpose of the Bargmann-
 1407 Wigner programme, it is enough to know that the UIRs of $SO(2, 1)^\uparrow$ are labeled by the
 1408 real eigenvalue of the quadratic Casimir operator $C_2[\mathfrak{so}(2, 1)]$ of the Lorentz algebra (and
 1409 another real parameter for the principal and complementary series, cf. Footnote 14). Since
 1410 the momentum is vanishing, the states span a constant field ψ on Minkowski spacetime
 1411 taking values in these UIRs of the Lorentz group $SO(2, 1)^\uparrow$. A relativistic equation is then
 1412 $(C_2[\mathfrak{so}(2, 1)] - c)\psi = 0$, which asserts that the states ψ are eigenvectors of the Casimir
 1413 operator with eigenvalue $c \in \mathbb{R}$.

1414 B.2 Massive representations

1415 Consider a massive representation labeled by mass $m > 0$ and spin $s \in \mathbb{R}$.

1416 B.2.1 (Half-)integer spins

1417 For integer spin $s \in \mathbb{N}$, the Klein-Gordon equation (48) together with the tracelessness
 1418 condition (51) and the transversality condition (53) for a totally symmetric tensor $\varphi_{\mu_1 \dots \mu_s}$
 1419 provide relativistic field equations whose positive-energy solutions represent the corre-
 1420 sponding UIR. Equivalently, for non-vanishing integer spin $s \in \mathbb{N}_0$, they can be summa-
 1421 rized by the following set of equations:

$$1422 \eta^{\mu_1 \mu_2} \varphi_{\mu_1 \dots \mu_s} = 0, \quad m \varphi_{\mu_1 \dots \mu_s} \pm \epsilon_{\mu_1 \nu \rho} p^\nu \varphi^\rho_{\mu_2 \dots \mu_s} = 0. \quad (94)$$

1423 where we take $\epsilon_{012} = -1$. Notice that the transversality condition (53) directly follows
 1424 from the second equation in (94). Moreover, note that there is no need to explicitly
 1425 symmetrize the last equation in its free indices when the tracelessness and transversality
 1426 conditions hold true. In turn, the Klein-Gordon equation follows from repeated application
 1427 of the second equation in (94). The two possible signs in the last equation stand for the two
 1428 possible values $\pm s$ of the “helicity” of the massive particle. This system of equations can
 1429 be generalized to the AdS_3 background and be supersymmetrized, see [38] and references
 therein.

1430 B.2.2 Fractional spins

1431 In the case of the massive UIRs where the real number s is neither integer nor half-
 1432 integer (“fractional spin”, see e.g. [39] for a review), one should stress that although the
 1433 number of physical components is one (the UIRs of the massive little group $SO(2)$ are
 1434 one-dimensional since this group is Abelian) nevertheless their corresponding covariant
 1435 description necessarily involve relativistic field equations with an infinite number of com-
 1436 ponents (since there are no finite-dimensional irreps of the Lorentz group $SO(2, 1)^\uparrow$ with
 1437 such values of the spin).

1438 The positive-energy solutions to the system of the four equations (48), (50), (51),
 1439 (52) formally describe a massive UIR of mass m and spin $s \in \mathbb{R}$ (as can be checked by
 1440 computing the value of the quartic Casimir operator). Note that the field $\Phi(p, u)$ is not
 1441 polynomial in the auxilliary vector u^μ when $s \notin \mathbb{N}$. Finding a suitable functional space is
 1442 a subtle issue that we will not attempt to address. In fact, the construction of manifestly
 1443 $IO(2, 1)$ -covariant field equations proved to be a rather difficult task.

1444 Several approaches have been followed in the literature. We refer to reader to [39] and
 1445 the introduction of the paper [40] for reviews. In the following, we will review the results
 1446 obtained in [41] for the linear relativistic equations whose positive-energy solutions span
 1447 the massives UIRs where the spin s is neither integer nor half-integer.

1448 The Cortes-Plyushchay equations proposed in [41] read¹⁵

$$V_\mu \psi = 0, \quad V_\mu := s P_\mu - i \epsilon_{\mu\nu\lambda} P^\nu \widetilde{M}^\lambda + m \widetilde{M}_\mu, \quad (95)$$

1449 where the three operators $\widetilde{M}_\mu := \frac{1}{2} \epsilon_{\mu\nu\rho} M^{\nu\rho}$ generate the $\mathfrak{so}(2,1)$ Lorentz algebra in
 1450 $D = 2 + 1$ dimensions ($i [\widetilde{M}_\mu, \widetilde{M}_\nu] = \epsilon_{\mu\nu\rho} \widetilde{M}^\rho$), so that the quadratic Casimir (4) is equal
 1451 to $\mathcal{C}_2[\mathfrak{so}(2,1)] = -\widetilde{M}_\mu \widetilde{M}^\mu$. In the above equations (95), the real number s is assumed to
 1452 be nonzero. Contracting the above equations with \widetilde{M}^μ , P^μ and $\epsilon^{\mu\nu\lambda} P_\mu \widetilde{M}_\lambda$ produces the
 1453 following three equations

$$((s-1)W + m \widetilde{M}^2) \psi = 0, \quad (sP^2 + mW) \psi = 0, \quad (P^2 \widetilde{M}^2 + W(m-W)) \psi = 0, \quad (96)$$

1454 where the scalar $W := P^\mu \widetilde{M}_\mu$ is, in three spacetime dimension, the analogue of the Pauli-
 1455 Lubanski vector. Since by assumption both s and m are non-zero, these three equations
 1456 are equivalent to

$$(m^2 \widetilde{M}^2 - s(s-1)P^2) \psi = 0, \quad (sP^2 + mW) \psi = 0, \quad P^2(P^2 + m^2) \psi = 0. \quad (97)$$

1457 If one discards the trivial representation of the Poincaré group where $P_\mu = 0 = \widetilde{M}_\mu$, one
 1458 gets the following three equations:

$$(\widetilde{M}^2 + s(s-1)) \psi = 0, \quad (W - sm) \psi = 0, \quad (P^2 + m^2) \psi = 0, \quad (98)$$

1459 the last two being the Pauli-Lubanski condition and the Klein-Gordon equation, whereas
 1460 the first sets the quadratic Casimir of the Lorentz group to $\mathcal{C}_2[\mathfrak{so}(2,1)] = s(s-1)$, which
 1461 indicates that the field ψ takes value in an irrep of the Lorentz group labeled by s . The
 1462 positive-energy solutions of the above field equations (98) transform in the UIR of mass
 1463 m and spin s . More directly, in the Lorentz frame where $p^\mu = (m, 0, 0)$, the Cortes-
 1464 Plyushchay equations (95) yield

$$(\widetilde{M}_0 - s) \psi = 0, \quad (\widetilde{M}_1 - i \widetilde{M}_2) \psi = 0. \quad (99)$$

1465 If one takes $L_\pm := \widetilde{M}_1 \pm i \widetilde{M}_2$ as raising/lowering operators of the Lorentz algebra $\mathfrak{so}(2,1)$,
 1466 then these equations assert that the state of momentum $p^\mu = (m, 0, 0)$ is a lowest-weight
 1467 state of $\mathfrak{so}(2,1)$. This implies that the positive-energy solutions are fields taking values in
 1468 a representation of the Lorentz algebra bounded from below. For $s \notin \frac{1}{2}\mathbb{N}$, one concludes
 1469 that the field ψ takes values in an infinite-dimensional UIR of the Lorentz algebra $\mathfrak{so}(2,1)$
 1470 belonging to the discrete series.

1471 The cases with $s = -j < 0$, where $j \in \frac{1}{2}\mathbb{N}$ is a non-vanishing (half)integer, correspond
 1472 to the non-unitary spin- j irreducible representations of the Lorentz algebra $\mathfrak{so}(2,1)$ with
 1473 quadratic Casimir $\mathcal{C}_2[\mathfrak{so}(2,1)] = j(j+1)$, in which case the Cortes-Plyushchay equation
 1474 propagates the massive fields with (half)integer spins discussed around (94).

1475 Manifest covariance groups the three components of the equations as the components
 1476 of a vector, but let us mention that only two of the three equations (95) are enough to
 1477 produce the third one. These equations are integrable in the sense that the commutator
 1478 $[V_\mu, V_\nu] \psi$ vanishes on a field ψ solution of (95). We refer to [40] for an extended discussion
 1479 of these equations.

¹⁵One can show that the operator V_μ can be obtained by the dimensional reduction of the Siegel-Zwiebach massless operator in (89).

1480 B.3 Massless representations

1481 The massless little group in $D = 2 + 1$ spacetime dimensions is $ISO(1) \cong \mathbb{R}$ that is
 1482 abelian, hence massless UIRs are one-dimensional and labeled by a single real parameter
 1483 $\mu \in \mathbb{R}$. Therefore, all massless UIRs of the Poincaré group $ISO(2, 1)^\dagger$ have a single
 1484 physical component. Nevertheless, we will stick to the distinction “helicity” vs “infinite-
 1485 spin” representations.

1486 B.3.1 Helicity representations

The helicity representations correspond to the particular case $\mu = 0$. Two case arises
 whether the representation of the Lorentz group $SO(2, 1)^\dagger$ is either single or double val-
 ued: the “helicity” is effectively zero or one-half, which corresponds to the fact that a
 massless field in three spacetime dimensions can always be dualized to a massless scalar
 or a Dirac spinor, as will be reviewed now. The manifestly covariant field equations are
 similar to those for the massless helicity cases in $D > 3$ studied above, except that only
 symmetric (spinor-)tensor gauge fields $\varphi_{\mu_1 \dots \mu_s} = \varphi_{(\mu_1 \dots \mu_s)}$ are allowed (the spinor index
 is not written). Equivalently, only field strengths $\mathcal{K}_{\mu_1 \nu_1 | \dots | \mu_s \nu_s}$ labeled by rectangular
 two-row Young diagrams are allowed. Moreover, higher (gamma-)traces of those field
 strengths must be set to zero. Indeed, if in three dimensions one were to set to zero the
 single (gamma-)trace of the field strength \mathcal{K} , one would obtain that the field strength
 itself should vanish on-shell, resulting in the absence of propagating degrees of freedom.
 More precisely, upon Hodge-dualizing the s pairs of antisymmetric indices of the spin- s
 field strength one obtains a totally symmetric (spinor-)tensor

$$\tilde{\mathcal{K}}_{\mu_1 \dots \mu_s} := \frac{1}{2^s} \epsilon_{\mu_1 \nu_1 \rho_1} \cdots \epsilon_{\mu_s \nu_s \rho_s} \mathcal{K}^{\nu_1 \rho_1 | \dots | \nu_s \rho_s},$$

1487 where the latter (spinor-)tensor is completely symmetric in its spacetime indices.

1488 The closure and coclosure conditions on the field strength \mathcal{K} are equivalent to coclosure
 1489 and closure condition on its dual:

$$\partial^{\mu_1} \tilde{\mathcal{K}}_{\mu_1 \mu_2 \dots \mu_s} = 0, \quad \partial_\mu \tilde{\mathcal{K}}_{\nu \rho_1 \dots \rho_{s-1}} - \partial_\nu \tilde{\mathcal{K}}_{\mu \rho_1 \dots \rho_{s-1}} = 0. \quad (100)$$

1490 The field strength $\tilde{\mathcal{K}}$ begin closed, it is exact:

$$\tilde{\mathcal{K}}_{\mu_1 \dots \mu_s} = p_{\mu_1} \cdots p_{\mu_s} \phi, \quad (101)$$

1491 where ϕ is a (spinor) scalar.

1492 The higher-trace equations on the field strength \mathcal{K} for a propagating, massless helicity
 1493 representation in three dimensions, are then for bosons

$$\eta^{\mu_1 \mu_2} \tilde{\mathcal{K}}_{\mu_1 \mu_2 \mu_3 \dots \mu_s} = 0, \quad s > 1, \quad (102)$$

1494 with the usual massless Klein-Gordon and Maxwell equations for $s = 0$ and 1, respectively,
 1495 and for fermions

$$\gamma^\mu \tilde{\mathcal{K}}_{\mu \nu_2 \dots \nu_s} = 0 \quad (103)$$

1496 for the spin $s + \frac{1}{2} > \frac{1}{2}$ cases; the spin- $\frac{1}{2}$ case being of course given by $\gamma^\mu \partial_\mu \phi = 0$, where
 1497 again, the spinor indices are not written and the three γ^μ matrices are three Dirac (in fact
 1498 Pauli) matrices in $D = 2 + 1$ dimensions.

1499 The conclusion is that all these descriptions of bosonic (respectively, fermionic) mass-
 1500 less fields are dual to each others, for all (half-)integer values of the “spin” s , in accordance
 1501 with fact that the positive-energy solutions of the above Bargmann-Wigner equations

(102) (respectively, (103), for fermions) carry a single (respectively, double) valued helicity representations of the Poincaré group $ISO(2,1)^\uparrow$. Concretely, these fields are dual a scalar (or spinor) field. More explicitly, the on-shell duality relation between the gauge fields $\varphi_{\mu_1 \dots \mu_s}$, the field strengths $\mathcal{K}_{\mu_1 \nu_1 | \dots | \mu_s \nu_s}$ and the massless scalar (or spinor) field ϕ is (101).

1507 B.3.2 Infinite spin representations

1508 The positive-energy solutions of the Wigner equations (74)-(76), reviewed in Subsec-
1509 tion 5.3.2, transform in the massless infinite-spin representation of the Poincaré group
1510 $ISO(2,1)^\uparrow$, labeled by $\mu > 0$. The paper [42] provided an extensive discussion of massless
1511 infinite-spin particles in $D = 2 + 1$ dimensions.

1512 B.4 Tachyonic representations

1513 Finally, in order to be exhaustive, we end this section by mentioning that the relativistic
1514 equations (84)-(87) provide an exhaustive solution of the Bargmann-Wigner programme
1515 in the tachyonic case. Indeed, the little group $SO(1,1)$ of a spacelike momenta in $D = 2+1$
1516 dimensions is Abelian, thus its UIRS of are labeled by a single parameter $s \in \mathbb{R}$.

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