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A Study of QCD Radiation in VBF Higgs Production with VINCIA and PYTHIA

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¹ Abstract

We discuss and illustrate the properties of several parton-shower models available in 2 PYTHIA and VINCIA, in the context of Higgs production via vector boson fusion (VBF). 3 In particular, the distinctive colour topology of VBF processes allows to define observ-4 ables sensitive to the coherent radiation pattern of additional jets. We study a set of 5 such observables, using the VINCIA sector-antenna shower as our main reference, and con-6 trast it to PYTHIA's transverse-momentum-ordered DGLAP shower as well as PYTHIA's 7 dipole-improved shower. We then investigate the robustness of these predictions as suc-8 cessive levels of higher-order perturbative matrix elements are incorporated, including 9 next-to-leading-order matched and tree-level merged calculations, using POWHEG BOX 10 and SHERPA respectively to generate the hard events. 11

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Figure 1: QCD colour flow of the LO VBF Higgs production process. Due to the kinematics of the interaction, QCD radiation is directed in the forward region of the detector.

34 1 Introduction

Higgs boson production via Vector Boson Fusion (VBF) — fig. 1 — is among the most im-35 portant channels for Higgs studies at the Large Hadron Collider (LHC). With a Standard-36 Model (SM) cross section of a few pb at LHC energies, VBF accounts for order 10% of the 37 total LHC Higgs production rate [1]. The modest rate is compensated for by the signature 38 feature of VBF processes: two highly energetic jets generated by the scattered quarks, in 39 the forward and backward regions of the detector respectively, which can be tagged ex-40 perimentally and used to significantly reduce background rates. Moreover, the distinct 41 colour flow of the VBF process at leading order (LO), highlighted by the coloured thick 42 dashed lines in fig. 1, strongly suppresses any coherent bremsstrahlung into the central 43 region, leaving this region comparatively clean and well suited for precision studies of the 44 Higgs boson decay products. With over half a million Higgs bosons produced in the VBF 45 channel in total during Run II of the LHC and a projection that this will more than double 46 during Run III, studies of this process have already well and truly entered the realm of 47 precision physics. 48

On the theory side, the current state of the art for the H + 2j process in fixed-order 49 perturbation theory is inclusive next-to-next-to-next-to-leading order QCD [2], fully differ-50 ential next-to-next-to-leading order (NNLO) QCD [3–6] and next-to-leading-order (NLO) 51 electroweak (EW) calculations [7]. These calculations of course only offer their full pre-52 cision for observables that are non-zero already at the Born level, such as the total cross 53 section and differential distributions of the Higgs boson and tagging jets. For more ex-54 clusive event properties, such as bremsstrahlung and hadronisation corrections, the most 55 detailed description is offered by combinations of fixed-order and parton-shower calcula-56 tions. To this end, two recent phenomelogical studies [8,9] compared different NLO+PS 57 simulations among each other as well as to NLO and NNLO calculations. These compara-58

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⁵⁹ tive studies catered to two needs; firstly, the reliability of matched calculations was tested

 $_{60}$ $\,$ in regions where resummation effects are small. Furthermore, a more realistic estimate of

parton-shower as well as matching uncertainties was obtained by means of different shower
 and matching methods in independent implementations.

The earlier of the two studies [8] highlighted that different NLO+PS implementations 63 describe the intrinsically coherent radiation in this process quite differently, and that the 64 uncertainties arising from the choice of the shower and matching implementation can 65 persist even at the NLO-matched level. Among its central results, the study [8] confirmed 66 the observation of [10] that PYTHIA's default shower [11–13] describes the emission pattern 67 of the third jet poorly, essentially missing the coherence of the initial-final dipoles. This 68 effect was most pronounced for MADGRAPH_AMC@NLO [14] + PYTHIA, for which a global 69 recoil scheme must be used in both the time-like and space-like shower in order to match 70 the subtraction terms implemented in MADGRAPH_AMC@NLO. For POWHEG-BOX [15] +71 PYTHIA, the difference persisted when using the global recoil scheme¹. However, changing 72 to PYTHIA's alternative dipole-recoil scheme [16], which should reproduce coherence effects 73 more faithfully, improved the agreement, both with calculations starting from H + 3j as 74 well as with the angular-ordered coherent shower model in HERWIG 7 [17]. 75

The more recent study [9] highlighted a number of interesting aspects of vector boson 76 fusion that can be exploited to enhance the signal-to-background ratio in future measure-77 ments: Firstly, if the Higgs boson is boosted, the *t*-channel structure of the VBF matrix 78 elements leads to less QCD radiation when compared to the irreducible background from 79 gluon-gluon fusion. Secondly, it was found that a global jet veto provides a similarly ef-80 fective cut as a central jet veto, leading to much reduced theoretical uncertainties, and 81 in particular eliminating the need to resum non-global logarithms associated with inhib-82 ited radiation in the rapidity gap. Despite a good overall agreement between fixed-order 83 NNLO and NLO-matched parton shower predictions, the study also pointed out a few 84 subtle disagreements for highly boosted Higgs boson topologies. In these scenarios, the 85 standard fixed-order paradigm of operating with a single factorisation scale is no longer 86 appropriate, because higher-order corrections should be resummed individually for the two 87 impact factors in the structure-function approach. 88

The uncertainties arising from matching systematics in vector-boson-fusion and vector-89 boson-scattering processes (VBS) have also been studied in the past [18] with rather good 90 agreement between different showers at the level of H + 3j NLO+PS calculations [19], 91 although in that study, only the POWHEG matching scheme was considered. Very recently, 92 two extensive reviews [20,21] collected experimental results and theoretical developments 93 in VBS processes in view of the high-luminosity upgrade of the LHC as well as future 94 colliders. A summary of Monte Carlo event generators used in the modelling of VBS 95 processes in ATLAS was presented in [22]. 96

On the experimental side, recent studies of VBF Higgs production by ATLAS [23, 24] and CMS [25, 26] have used PYTHIA's default shower model matched to the NLO via the POWHEG technique, with only one of them [23] employing PYTHIA's dipole option. The associated modelling uncertainties, and ways to reduce them, therefore remain of high current relevance.

We extend the comparative study of [8] to include the new VINCIA sector-antenna shower [27] that has become available starting from PYTHIA version 8.304. Based on findings pertaining to antenna [28–31] and dipole [32–35] showers, we expect that, at least at leading colour, VINCIA's showers capture QCD coherence effects in VBF more accurately than PYTHIA's default shower. To this end, we note that the emitter-recoiler

¹We note that the global recoil scheme is the default choice only for PYTHIA's space-like DGLAP shower, while the time-like DGLAP shower uses a dipole-like recoil scheme per default.

agnostic antenna recoil employed in VINCIA is free of adverse kinematic effects [36]. We also 107 consider two new observables designed to further probe the amount of coherent radiation 108 by measuring the summed transverse energy H_T for $|\eta| < 0.5$ and for $|\eta - \eta_0| < 0.5$ 109 respectively, where η_0 is the midpoint between the two tagging jets. To investigate the 110 robustness of the predictions, we include not only POWHEG-BOX + PYTHIA [13, 15] but 111 also a new dedicated implementation of the CKKW-L merging scheme [37–39] for sector 112 showers [40], with hard events with up to four additional jets generated by SHERPA 2 [41, 113 42]. We emphasise that this is currently the only multi-jet merging approach in PYTHIA 8.3 114 which can handle VBF processes². Additionally, we highlight the systematic uncertainties 115 arising from the use of vetoed showers in the POWHEG scheme and make recommendations 116 for settings related to the use of these in PYTHIA. 117

This study is structured as follows. We begin with an overview of the setup for our simulations in section 2; starting with an overview of the fixed order, shower, matched and merged calculations and leading towards a description of the analysis we perform. We then move on to discuss the results of our analysis in section 3, with our conclusions and recommendations listed in section 4.

¹²³ 2 Setup of the Simulation

We consider Higgs production via VBF in proton-proton collisions at the high-luminosity LHC with a centre-of-mass energy of $\sqrt{s} = 14$ TeV.

The simulation is factorised into the generation of the hard process using SHERPA 2 (for the LO merging samples) and POWHEG-BOX v2 (for the NLO matched samples) and subsequent showering with PYTHIA 8.306. A cross check is also performed using PYTHIA's internal Born-level VBF process. Details on the hard-process setups are given in section 2.1.

Since we expect the VINCIA antenna shower to account for coherence more faithfully than does PYTHIA's default "simple" p_{\perp} -ordered DGLAP shower, we take VINCIA's description as the baseline for our comparisons, contrasting it to PYTHIA's default and "dipolerecoil" options. Details on the shower setups are given in section 2.2.

Higher fixed-order corrections are taken into account at NLO+PS accuracy via the POWHEG scheme, and for VINCIA also in the CKKW-L scheme up to $\mathcal{O}(\alpha_{\rm S}^4)$. We expect that these corrections will be smaller for coherent shower models than for incoherent ones, hence these comparisons serve both to test the reliability of the baseline showers and to illustrate any ambiguities that remain after these corrections are included. Details on the matching and merging setups are given in section 2.3.

Finally, in section 2.4, we define the observables and the VBF analysis cuts that are used for the numerical studies in section 3.

Note that, since we are primarily interested in exploring the coherence properties of 143 the perturbative stages of the event simulation, most of the results will be at the so-called 144 "parton level", i.e. without accounting for non-perturbative or non-factorisable effects 145 such as hadronisation, primordial $k_{\rm T}$, or multi-parton interactions (MPI). Although this 146 is not directly comparable to physical measurements (nor is the definition universal since 147 different shower models define the cutoff differently), the factorised nature of the infrared 148 and collinear safe observables we consider imply that, while non-perturbative effects may 149 act to smear out the perturbative differences and uncertainties, they would not in general 150 be able to obviate them, thus making studies of the perturbative stages interesting in 151

 $^{^{2}}$ We do note that a technical (but due to the use of incoherent IF kinematics unphysical) fix was introduced in PYTHIA 8.242 and is planned to be re-implemented in a future version of PYTHIA 8.3.

their own right. Nevertheless, with jet $p_{\rm T}$ values going down to 25 GeV and $H_{\rm T}$ being sensitive to the overall amount of energy scattered into the central region, we include further comparisons illustrating the effect of non-perturbative corrections at the end of section 3.

156 2.1 Hard Process

For the parton-level event generation, we use a stable Higgs boson with a mass of $M_{\rm H} = 125 \, {\rm GeV}$, and we set the electroweak boson masses and widths to

 $M_{\rm Z} = 91.1876 \text{ GeV}, \quad \Gamma_{\rm Z} = 2.4952 \text{ GeV}, \quad (1)$ $M_{\rm W} = 80.385 \text{ GeV}, \quad \Gamma_{\rm W} = 2.085 \text{ GeV}.$

Electroweak parameters are derived from this set with the additional input of the electromagnetic coupling constant at the Z pole ($\alpha(M_Z)$ scheme, EW_SCHEME = 2 in SHERPA): ¹⁶¹

$$\frac{1}{\alpha(M_{\rm Z})} = 128.802\,. \tag{2}$$

We treat all flavours including the bottom quark as massless and use a diagonal CKM mixing matrix. In both SHERPA and POWHEG-BOX, we use the CT14_NNLO_as118 [43] PDF set provided by LHAPDF6 [44] with the corresponding value of α_S . For the sample generated with PYTHIA's internal VBF implementation, we use its default NNPDF23_lo_as_0130_qed PDF set [45, 46].

We consider only VBF topologies, neglecting Higgsstrahlung contributions which ap-167 pear at the same order in the strong and electroweak coupling. Identical-flavour inter-168 ference effects are neglected in events generated with POWHEG-BOX and PYTHIA, but are 169 included in events obtained with SHERPA, although their impact was found to be small [9]. 170 At NLO, the process is calculated in the structure function approximation, neglecting 171 interferences between the two quark lines. For both, internal and external events, only 172 a single scale will be assigned per event, notwithstanding that different scales could in 173 principle be assigned to the two forward-scattered quarks. Differences pertaining to the 174 scale assignment in internal and external events will be discussed in section 3.1. 175

Tree-level event samples with up to four additional jets are generated using an HPCenabled variant of SHERPA 2 [41, 42], utilising the COMIX matrix-element generator [47]. To facilitate efficient parallelised event generation and further processing, events are stored in the binary HDF5 data format [42]. The factorisation and renormalisation scales are chosen to be

$$\mu_{\rm F}^2 = \mu_{\rm R}^2 = \frac{\hat{H}_{\rm T}^2}{4} \quad \text{with} \quad \hat{H}_{\rm T} = \sum_j p_{{\rm T},j} + \sqrt{M_{\rm H}^2 + p_{{\rm T},{\rm H}}^2} \,. \tag{3}$$

and jets are defined according to the $k_{\rm T}$ clustering algorithm with R = 0.4 and a cut at 20 GeV.

PYTHIA's internal events are generated with scales governed by the two switches SigmaProcess:factorScale3VV and SigmaProcess:renormScale3VV, respectively. Their default values = 2 and = 3, respectively, correspond to the choices

$$u_{\rm F}^2 = \sqrt{m_{\rm T,V_1}^2 m_{\rm T,V_2}^2} \equiv \sqrt{(M_{\rm V_1}^2 + p_{\rm T,q_1}^2)(M_{\rm V_2}^2 + p_{\rm T,q_2}^2)}, \qquad (4)$$

$$\mu_{\rm R}^2 = \sqrt{m_{\rm T,V_1}^2 m_{\rm T,V_2}^2 m_{\rm T,H}^2} \equiv \sqrt[3]{(M_{V_1}^2 + p_{\rm T,q_1}^2)(M_{V_2}^2 + p_{\rm T,q_2}^2)m_{\rm T,H}^2},$$
(5)

with the pole masses of the exchanged vector bosons M_{V_1} , M_{V_2} , the transverse mass of the Higgs boson $m_{T,H}$, and the transverse momenta of the two final-state quarks p_{T,q_1} , p_{T,q_2} .

For NLO calculations matched to parton showers, we consider the POWHEG [48, 49] formalism. POWHEG samples are generated with POWHEG-BOX v2 [15, 50] with the factorisation and renormalisation scales chosen as

$$\mu_{\rm F}^2 = \mu_{\rm R}^2 = \frac{M_{\rm H}}{2} \sqrt{\left(\frac{M_{\rm H}}{2}\right)^2 + p_{\rm T,H}^2} \,. \tag{6}$$

Since the study in [8] did not find any significant effect from the choice of the "hdamp" parameter in POWHEG, we do not include any such damping here, corresponding to a choice of hdamp = 1.

195 2.2 Showers

The hard events defined above are showered with the three following shower models, which are all available in PYTHIA 8.306:

- VINCIA's sector antenna shower [27]. The "sector" mode is the default option for VINCIA since PYTHIA 8.304 and also enables us to make use of VINCIA's efficient CKKW-L merging [40]. We expect it to exhibit the same level of coherence as the fixed-order matrix elements, at least at leading colour (LC), since its QCD antenna functions and corresponding phase-space factorisations explicitly incorporate the soft-eikonal function for all possible (LC) colour flows. Of particular relevance to this study is its coherent treatment of "initial-final" (IF) colour flows.
- PYTHIA's default "simple shower" model [11,12], which implements p_{\perp} -ordered DGLAP evolution with dipole-style kinematics. For IF colour flows, however, the kinematic dipoles are not identical to the colour dipoles, and this can impact coherence-sensitive observables [51].

• PYTHIA's "simple shower" with the dipole-recoil option [16]. Despite its name, this not only changes the recoil scheme; in fact, it replaces the two independent DGLAP evolutions of IF dipoles by a coherent, antenna-like, dipole evolution, while keeping the DGLAP evolution of other dipoles unchanged. This option should therefore lead to radiation patterns exhibiting a similar level of coherence as VINCIA.

Ordinarily, PYTHIA would of course also add decays of the Higgs boson, and any final-214 state radiation associated with that. However, as a colour-singlet scalar with $\Gamma_{\rm H} \ll \Lambda_{\rm QCD}$ 215 and $\Gamma_{\rm H}/M_{\rm H} \sim \mathcal{O}(10^{-5})$, its decay can be treated as factorised from the production process 216 to a truly excellent approximation. For the purpose of this study, we therefore keep 217 the Higgs boson stable, to be able to focus solely on the radiation patterns of the VBF 218 production process itself, without the complication of decay products in the central region. 219 For all of the shower models, we retain PYTHIA's default PDF choice³, regardless of 220 which PDF set was used to generate the hard process. This is done to remain consistent 221 with the default shower tunings [52] and due to the better-controlled backwards-evolution 222 properties of the default set [53]. 223

Per default, the shower starting scale is chosen to be the factorisation scale of the hard process,

$$\mu_{\rm PS}^2 = \mu_{\rm F}^2 \,. \tag{7}$$

 $^{^{3}}$ NNPDF23_lo_as_0130_qed.

In VINCIA, this scale can be varied by a multiplicative "fudge" factor, controlled by Vincia:pTmaxFudge,

$$\mu_{\mathrm{PS}}^2 = k_{\mathrm{fudge}} \, \mu_{\mathrm{F}}^2 \, ,$$

while in PYTHIA, the starting scales of the initial-state and final-state showers can be varied independently,

$$\mu_{\rm PS,FSR}^2 = k_{\rm fudge,FSR} \, \mu_{\rm F}^2 \,,$$
$$\mu_{\rm PS,ISR}^2 = k_{\rm fudge,ISR} \, \mu_{\rm F}^2 \,,$$

²³⁰ controlled by TimeShower:pTmaxFudge and SpaceShower:pTmaxFudge, respectively.

In a similar vein, the strong coupling in the shower is evaluated at the shower p_{T-232} scale⁴, modified by renormalisation-scale factors k_{ren} . In PYTHIA, the strong coupling at the Z mass is set to $\alpha_S(M_Z) = 0.1365$ and independent scale factors for ISR and FSR are implemented,

$$\begin{split} &\alpha_{\rm S}^{\rm Pythia,FSR}(p_{\perp {\rm evol},FSR}^2) = \alpha_{\rm S}^{\overline{\rm MS}}(k_{\rm R,FSR}\,p_{\perp {\rm evol},FSR}^2)\,,\\ &\alpha_{\rm S}^{\rm Pythia,ISR}(p_{\perp {\rm evol},ISR}^2) = \alpha_{\rm S}^{\overline{\rm MS}}(k_{\rm R,ISR}\,p_{\perp {\rm evol},ISR}^2)\,. \end{split}$$

These can be set via TimeShower:renormMultFac and SpaceShower:renormMultFac, respectively, and are unity by default. The transverse-momentum evolution variables $p_{\perp evol,FSR}^2$ and $p_{\perp evol,ISR}^2$ are defined as in [11].

For VINCIA, on the other hand, a more refined choice can be made with separate renormalisation factors being implemented for (initial- and final-state) emissions, (initial- and final-state) gluon splittings, and (initial-state) quark conversions. These have the default settings:

$$\begin{split} k^{\rm F}_{\rm R,Emit} &= 0.66\,, \quad k^{\rm F}_{\rm R,Split} = 0.8\,, \\ k^{\rm I}_{\rm R,Emit} &= 0.66\,, \quad k^{\rm I}_{\rm R,Split} = 0.5\,, \quad k^{\rm I}_{\rm R,Conv} = 0.5\,, \end{split}$$

²⁴² which can be set via the parameters

Vincia:renormMultFacEmitF Vincia:renormMultFacSplitF

243 Vincia:renormMultFacEmitI Vincia:renormMultFacSplitI Vincia:renormMultFacConvI.

Additionally, VINCIA uses the CMW scheme [54] (while PYTHIA does not), i.e. it evaluates

²⁴⁵ the strong coupling according to

$$\alpha_{\rm S}^{\rm CMW} = \alpha_{\rm S}^{\overline{\rm MS}} \left(1 + \frac{\alpha_{\rm S}^{\overline{\rm MS}}}{2\pi} \left[C_A \left(\frac{67}{18} - \frac{\pi^2}{6} \right) - \frac{5n_f}{9} \right] \right) , \qquad (8)$$

where $\alpha_{\rm S}^{\overline{\rm MS}}(M_{\rm Z}) = 0.118$, so that

$$\alpha_{\rm S}^{\rm Vincia}(p_{\perp}^2) = \alpha_{\rm S}^{\rm CMW}(k_{\rm R} \, p_{\perp}^2) \tag{9}$$

with the VINCIA evolution variable as defined in [27].

⁴We refer to the argument of the strong coupling used in the shower as the shower renormalisation scale.

248 2.3 Matching and Merging

In the following, we will briefly review the defining features of the POWHEG NLO matching and the CKKW-L merging schemes we will use in this study. In particular, we will focus on the technicalities and practicalities to ensure a consistent use. Detailed reviews of the POWHEG schemes can for instance be found in [55] and [56]. The CKKW-L scheme is explained in detail in [39] and its extension to the VINCIA sector shower in [40].

254 2.3.1 POWHEG Matching

In the POWHEG formalism, events are generated according to the inclusive NLO cross section with the first emission generated according to a matrix-element corrected noemission probability.

Since the shower kernels in the POWHEG no-emission probability are replaced by the 258 ratio of the real-radiation matrix element to the Born-level one, it is independent of the 259 shower it will later be matched to. It is, however, important to stress that generally, 260 the POWHEG ordering variable will not coincide with the ordering variable of the shower. 261 Starting a shower with a different ordering variable at the POWHEG scale of the first emis-262 sion might thus lead to over- or undercounting emissions. A simple method to circumvent 263 this was presented in [57]. There, the shower is started at the phase space maximum 264 (a so-called "power shower" [58]) and emissions harder than the POWHEG one are vetoed 265 until the shower reaches a scale below the scale of the first emission. For general ordering 266 variables, there is, however, no guarantee that once the shower falls below the scale of the 267 POWHEG emission it will not generate a harder emission later on in the evolution. This is 268 especially important if the shower is not ordered in a measure of hardness but e.g. in emis-269 sion angles, such as the HERWIG \tilde{q} shower [59]. In these cases, it is advisable to recluster 270 the POWHEG emission and start a truncated and vetoed shower off the Born state [48], 271 see also [60-62] for the use of truncated showers in merging schemes. This scheme also 272 avoids the issue that in vetoed showers, all emissions in the shower off a Born+1-jet state 273 are compared against the POWHEG emission as if they were the first emission themselves. 274 But from the point of view of kinematics and colour they will still be the second, third, 275 etc. 276

However, since all showers we consider here are ordered in a notion of transverse momentum, it shall suffice for our purposes to use the simpler "vetoed power shower" scheme. To this end, we have amended the existing POWHEG user hook for PYTHIA's showers by a dedicated one for POWHEG+VINCIA, which has been included in the standard release of PYTHIA starting from version 8.306; see appendix A for detailed instructions.

For both PYTHIA and VINCIA, we use a vetoed shower with the POWHEG $p_{\rm T}$ and 282 d_{ij} definitions, corresponding to the mode POWHEG:pTdef = 1. We define the POWHEG 283 scale with respect to the radiating leg and use PYTHIA's definition of emitter and recoiler, 284 corresponding to the modes POWHEG:pTemt = 0 and POWHEG:emitted = 0. Per default. 285 we choose to define the scale of the POWHEG emission by the minimum $p_{\rm T}$ among all 286 final-state particles, i.e. use POWHEG:pThard = 2, according to the suggestion in [63]. As 287 an estimate of the uncertainty of this choice, we vary the $p_{T,hard}$ scale to be the LHEF 288 scale and the $p_{\rm T}$ of the POWHEG emission, corresponding to the modes POWHEG:pThard = 289 0 and POWHEG:pThard = 1, respectively. 290

The purpose of these settings is to ensure maximally consistent scale definitions while not reverting to the (more involved) "truncated and vetoed shower" scheme mentioned above. While we deem the choices made here appropriate for the case at hand they remain ambiguous, effectively introducing systematic matching uncertainties into the (precision) calculation. As a means of estimating these uncertainties, we will discuss the influence of the $p_{\rm T,hard}$ scale setting on physical observables below in section 3.

297 2.3.2 CKKW-L Merging

Multi-leg merging schemes aim at correcting parton shower predictions away from the soft 298 and collinear regions. In the CKKW-L merging scheme [39], multiple inclusive tree-level 299 event samples are combined to a single inclusive one by introducing a (somewhat arbitrary) 300 "merging scale" $t_{\rm MS}$ which separates the matrix-element ($t > t_{\rm MS}$) from the parton-shower 301 $(t < t_{\rm MS})$ region. In this way, over-counting of emissions is avoided while accurate parton-302 shower resummation in logarithmically enhanced regions and leading-order accuracy in the 303 regions of hard, well-separated jets is ensured if the merging scale is chosen appropriately. 304 The missing Sudakov suppression in higher-multiplicity configurations is calculated 305 post-facto by the use of truncated trial showers between the nodes of the most probable 306 "shower history". In this context, the shower history represents the sequence of interme-307 diate states the parton shower at hand would (most probably) have generated to arrive 308 at the given *n*-jet state. Usually, this sequence is constructed by first finding all possible 309 shower histories and subsequently choosing the one that maximises the branching proba-310 bility, i.e., the product of branching kernels and the Born matrix element. As we employ 311 this scheme with VINCIA's sector shower, a few comments are in order. The objective of 312 the sector shower is to replace the probabilistic shower history by a deterministic history, 313 governed by the singularity structure of the matrix element. This means that at each 314 point in phase space only the most singular branching contributes. In the shower, this is 315 ensured by vetoing any branchings that do not abide by this; in the merging, this results 316 in a faster and less resource-intensive algorithm, as it is no longer required to generate a 317 large number of possible histories. Details and subtleties of VINCIA's sectorised CKKW-L 318 implementation can be found in [40]. 319

The CKKW-L merging scheme is in principle implemented for all showers in PYTHIA 320 8.3. However, the intricate event topology of VBF processes currently prohibits the use 321 of PYTHIA's default merging implementation⁵. We hence limit ourselves to study the 322 effect of merging with VINCIA, and have adapted VINCIA's CKKW-L implementation [40] 323 so that VBF processes are consistently treated. Specifically, the flag Vincia:MergeVBF = 324 on should be used, which restricts the merging to only consider shower histories that retain 325 exactly two initial-final quark lines. As a consequence, there must not be any "incomplete 326 histories" (histories that do not cluster back to a VBF Born configuration); this should be 327 guaranteed as long as the input event samples are of the VBF type only and no QED or 328 EW emissions are generated. A complete list of relevant settings for the use of VINCIA's 329 CKKW-L merging is collected in appendix B. 330

331 2.4 Analysis

We use the anti- $k_{\rm T}$ algorithm [64] with R = 0.4, as implemented in the FASTJET [65] package, to cluster jets in the range,

$$p_{\rm T} > 25 \,\,{\rm GeV}\,, \quad |\eta| < 4.5\,.$$

In addition, we employ typical VBF cuts to ensure that the two "tagging jets" are sufficiently hard, have a large separation in pseudorapidity, and are located in opposite hemispheres:

 $m_{j_1,j_2} \ge 600 \text{ GeV}, \quad |\Delta \eta_{j_1,j_2}| \ge 4.5, \quad \eta_{j_1} \cdot \eta_{j_2} \le 0.$

³³⁷ We consider the following observables:

⁵We note that a technical fix for this was available in PYTHIA 8.245 and will become available again in PYTHIA 8.3 in the future.

Pseudorapidity Distributions: at the Born level, the two tagging jets already have nontrivial pseudorapidity distributions. These are sensitive to showering chiefly via recoil effects and via the enhancement of radiation towards the beam directions. The third (and subsequent) jets are of course directly sensitive to the generated emission spectra. To minimise contamination from final-state radiation off the tag ging jets, we also consider the pseudorapidity of the radiated jet(s) relative to the midpoint of the two tagging jets,

$$\eta_{j_i}^* = \eta_{j_i} - \eta_0 , \qquad (10)$$

³⁴⁵ with the midpoint defined by:

$$\eta_0 = \frac{1}{2}(\eta_{j_1} + \eta_{j_2}) \ . \tag{11}$$

• Transverse Momentum Distributions: we expect coherence effects for the radiated jets (i > 2) to be particularly pronounced for radiation that is relatively soft in comparison to the characteristic scale of the hard process. Conversely, the transverse momenta of the two tagging jets should mainly be affected indirectly, via momentum-conservation (recoil) effects.

• Scalar Transverse Momentum Sum: as a more inclusive measure of the summed jet activity in the central rapidity region, we consider the scalar transverse momentum sum of all reconstructed jets (defined as above, i.e., with $p_{\rm T} > 25$ GeV),

$$H_{\rm T} = \sum_{j} |p_{{\rm T},j}|, \qquad (12)$$

in two particular regions:

- in the central rapidity region, $\eta \in \left[-\frac{1}{2}, +\frac{1}{2}\right]$

- around the midpoint of the tagging jets, $\eta^* \in \left[-\frac{1}{2}, +\frac{1}{2}\right]$, cf eq. (10).

We point out that, due to the way it is constructed, the second of these regions is not sensitive to the tagging jets, as it is not possible for them to fall within this region. Unlike the previous two observables, $H_{\rm T}$ is sensitive to the overall radiation effect in the given region, not just that of a certain jet multiplicity. As such, we expect $H_{\rm T}$ to give a measure of the all-orders radiation effects.

The analysis is performed using the RIVET analysis framework [66, 67] and based on the one used in [8].

364 **3** Results

In this section, we present the main results of our study based on the setup described in the last section. In fig. 2, the exclusive jet cross sections for up to 7 jets are shown at LO+PS and NLO+PS (via the POWHEG scheme) accuracy at the Born level. While there are very large differences between the three shower predictions at the leading order, there is good agreement between the NLO+PS predictions at least for the 2- and 3-jet cross sections.



Figure 2: Exclusive jet cross sections at LO+PS (*left*) and POWHEG NLO+PS (*right*) accuracy. The bands are obtained by a variation of the default shower starting scale by a factor of two or the variation of the hard scale, respectively.

371 3.1 Leading Order

It is instructive to start by studying the properties of the baseline leading-order + shower
 calculations, without including higher fixed-order corrections.

We use the leading-order event samples generated with SHERPA and by default let 374 the factorisation scale $\mu_{\rm F}^2$ define the shower starting scale. As a way to estimate the 375 uncertainty associated with this choice, we vary the shower starting scale $\mu_{\rm PS}^2$ by a factor 376 $k_{\rm fudge} \in \left[\frac{1}{2}, 2\right], \ \mu_{\rm PS}^2 = k_{\rm fudge} \mu_{\rm F}^2$. Strictly speaking, shower starting scales not equal to 377 the factorisation scale lead to additional PDF ratios in the no-branching probabilities 378 generated by the shower, but for factor-2 variations these are consistent with unity (since 379 the PDF evolution is logarithmic) and we therefore neglect them. Compared to the shower 380 starting scale, variations of the shower renormalisation scale only have a marginal effect 381 and are therefore not shown here. As we are primarily concerned with the shower radiation 382 patterns, we do not vary the scales in the fixed-order calculation. The effect of those 383 variations have been studied extensively in the literature before, cf. e.g. [8, 18]. 384

In fig. 3, the transverse momentum distributions of the two tagging jets and as well 385 as the transverse momentum and pseudorapidity distributions of the third-hardest jet are 386 shown. While the tagging jet $p_{\rm T}$ spectra agree well between VINCIA and PYTHIA with 387 dipole recoil, differences are visible for the third-jet observables, with similar shapes but 388 a slightly larger rate produced by the PYTHIA dipole-recoil shower. The distributions 389 obtained with the PYTHIA default shower, on the other hand, neither agree in shape nor 390 in the rate with the other two. In fact, almost no suppression of radiation in the central-391 rapidity region is visible and the shower radiation appears at a much higher transverse 392 momentum scale. The high emission rate in the default shower also implies that the 393 tagging jets receive much larger corrections with this shower than with the other models, 394 as evident from the tagging jet $p_{\rm T}$ distributions. 395

Figure 4 shows the $H_{\rm T}$ distributions in the previously defined central and midpoint regions. As for the third-jet pseudorapidity and transverse-momentum distributions, there is only a minor disagreement between PYTHIA dipole-recoil shower and VINCIA, while PYTHIA's DGLAP shower generates significantly more radiation in both regions.

For all observables considered here, we also note that the variation of the shower starting scale has a much more pronounced effect on the PYTHIA default shower than on VINCIA or on PYTHIA when the dipole-recoil option is enabled. Moreover, the starting-scale



Figure 3: Transverse momentum of the first tagging jet (*top left*), second tagging jet (*top right*), third jet (*bottom left*), and pseudorapidity of the third jet (*bottom right*) at LO+PS accuracy. The bands are obtained by a variation of the default shower starting scale by a factor of two.

variation affects the $p_{\rm T}$ distribution of the third jet more than it does the pseudorapidity distribution. This indicates that, while a tailored shower starting scale for the default shower might be able to mimic the phase space-suppression of the dipole/antenna showers to some extent, this would not by itself be sufficient to represent the dipole-antenna emission pattern of the third jet.



Figure 4: Scalar transverse momentum sum in the central rapidity region (*left*) and around the rapidity midpoint of the tagging jets (*right*) at LO+PS accuracy. The bands are obtained by a variation of the default shower starting scale by a factor of two.

We close this subsection by comparing showers off our externally generated Born-level 408 VBF events (i.e., ones generated by SHERPA and passed to PYTHIA for showering) to show-409 ers off internally generated ones (i.e., ones generated by PYTHIA's HiggsSM:ff2Hff(t:WW) 410 and HiggsSM:ff2Hff(t:ZZ) processes). This is intended as a cross check for any effects 411 caused by differences in how PYTHIA treats external vs internal events. For instance, for 412 external events, the external generator is responsible not only for computing the hard 413 cross section but also for setting the shower starting scale, via the HDF5 scales dataset 414 (equivalent to the Les Houches SCALUP parameter [68,69]). For our VBF events, the choice 415 made in SHERPA is identical to the factorisation scale eq. (3), 416

SHERPA VBF events:
$$\mu_{\rm PS}^2 \equiv \mu_{\rm F}^2 = \frac{\hat{H}_{\rm T}^2}{4} = \frac{1}{4} \left(\sum_j p_{{\rm T},j} + \sqrt{M_{\rm H}^2 + p_{{\rm T},{\rm H}}^2} \right)^2$$

For internally generated VBF events, PYTHIA's choice of the factorisation scale, and thereby also the shower starting scale, is designed to reflect the off-shellness of the two virtual-boson *t*-channel propagators, cf. eq. (5),

PYTHIA VBF events:
$$\mu_{\rm PS}^2 \equiv \mu_{\rm F}^2 = \sqrt{m_{{\rm T},{\rm V}_1}^2 m_{{\rm T},{\rm V}_2}^2} \equiv \sqrt{(M_{{\rm V}_1}^2 + p_{{\rm T},{\rm q}_1}^2)(M_{{\rm V}_2}^2 + p_{{\rm T},{\rm q}_2}^2)}$$

This choice ensures that the factorisation scale and shower starting scale will always be at least of order M_V^2 even when the outgoing quarks have low $p_T \ll M_V$, while for very large p_T values, it asymptotes to the geometric mean of the quark p_T values. While the minimum of the SHERPA choice is of the same order, $\mathcal{O}(M_H) \sim \mathcal{O}(M_V)$, the largetransverse-momentum limit is considerably larger. The expectation is therefore that, in the absence of matching or merging corrections, SHERPA-generated Born events will lead to higher amounts of hard shower radiation than PYTHIA-generated ones.

In fig. 5, the ratio of the two PYTHIA showers to VINCIA is shown for the $p_{\rm T}$ and $H_{\rm T}$ spectra using (left) PYTHIA LO and (right) SHERPA LO events. We immediately note that, in the low- p_{\perp} limit, the excess of soft radiation generated by PYTHIA's default shower (red) persists in both samples. In the high- p_{\perp} regions, the agreement between the simple shower and the two dipole/antenna options (blue and yellow) tends to be best for PYTHIA's internal hard process. This likely originates from the lower value for the default shower



Figure 5: Ratio of PYTHIA to VINCIA at LO+PS accuracy, comparing internal (*left*) and external (*right*) events. The bands are obtained by a variation of the factorisation scale (internal events) and shower starting scale (external events) by a factor of two.

433 starting scale in PYTHIA, which, as discussed above, imitates the propagator structure 434 of the Born process as closely as possible and hence *should* to some extent set a natural 435 boundary for strongly ordered propagators in the shower. For the dipole/antenna showers, 436 the sensitivity to the starting scale is far milder, as the relevant kinematic information is 437 encoded in the dipole invariant masses independently of the choice of starting scale.

438 3.2 Next-to-Leading Order Matched

In fig. 6, the POWHEG-matched transverse momentum distributions of the four hardest 439 jets are collected. In comparison to the LO+PS case discussed in the last section, it 440 is directly evident that the Born-jet $p_{\rm T}$ distributions are in good agreement between all 441 three shower models, including the default PYTHIA one, for which the tagging jet $p_{\rm T}$ 442 distributions undershoot the VINCIA curve only by an approximately constant factor of 443 order of five per cent. After POWHEG matching, almost perfect agreement is found for the 444 tagging jet transverse momentum distributions obtained with VINCIA and PYTHIA with 445 dipole recoil, as can be seen in fig. 8. The NLO corrections are, however, slightly smaller 446 for the former. The scale choice of the POWHEG emission has only mild effects on all three 447 showers for these tagging-jet observables. 448



Figure 6: Transverse momentum of the first tagging jet (top left), second tagging jet (top right), third jet (top left), and fourth jet (top right) at NLO+PS accuracy in the POWHEG scheme. The bands are obtained by a variation of the hard scale in the vetoed showers as explained in the text.

Good agreement is also found between all three shower models for the $p_{\rm T}$ of the third 449 jet, as shown in the bottom left panel of fig. 6. It must be noted that, again in the 450 case of the PYTHIA default shower, this agreement is subject to appropriately vetoing 451 harder emissions than the POWHEG one, which requires the definition of the POWHEG 452 scale according to the minimal $p_{\rm T}$ in the event, corresponding to the POWHEG:pThard = 453 2 setting, cf. section 2.3.1. Other choices again lead to too hard third jets and heavily 454 increased radiation in the central rapidity region, as can be inferred from the (relative) 455 rapidity distributions of the third jet in the top row of fig. 7, where the importance of 456 a judicious POWHEG scale choice is especially visible. As for the tagging jet spectra, the 457 agreement in both the third-jet transverse momentum and rapidity predictions between 458 VINCIA and the dipole-improved PYTHIA shower is almost perfect, as shown in fig. 9. 459 While the correction (which in this case is essentially a LO matrix-element correction) 460 is positive for VINCIA, it is negative for the dipole-improved PYTHIA shower. Moreover, 461 in the case of VINCIA, this correction affects mostly the high- $p_{\rm T}$ and the central-rapidity 462 region, whereas for PYTHIA's dipole-improved shower, the correction is negligible at zero 463 rapidity but bigger (and almost) constant at larger rapidities as well as for the transverse 464 momentum. 465



Figure 7: Pseudorapidity (*left column*) and relative rapidity to the tagging jets (*right column*) of the third jet (*top row*) and fourth jet (*bottom row*) at NLO+PS accuracy in the POWHEG scheme. The bands are obtained by a variation of the hard scale in the vetoed showers as explained in the text.



Figure 8: Detailed comparison of the PYTHIA dipole and VINCIA LO+PS and POWHEG NLO+PS predictions for the transverse momentum of the first tagging jet (*left*) and the second tagging jet (*right*).



Figure 9: Detailed comparison of the PYTHIA dipole and VINCIA LO+PS and POWHEG NLO+PS predictions for the transverse momentum (*left*) and rapidity of the third jet (*right*).



Figure 10: Scalar transverse momentum sum for $|\eta| < 0.5$ (*left*) and around the rapidity midpoint of the tagging jets (*right*) at NLO+PS accuracy in the POWHEG scheme. The bands are obtained by a variation of the hard scale in the vetoed showers as explained in the text.

The bottom right pane in fig. 6 and the bottom row in fig. 7 compare the $p_{\rm T}$ and (relative) rapidity predictions of the three shower models. While again rather good agreement in these distributions is found for the VINCIA shower and the dipole-improved PYTHIA shower, PYTHIA's default shower produces a harder spectrum, located more in the central rapidity region. Here, it is worthwhile noting that for two-jet POWHEG matching, the emission of the fourth jet is uncorrected in either of the shower models, so that the effects visible in these distributions are solely produced by the showers.

Lastly, fig. 10 shows the scalar transverse momentum for $|\eta| < 0.5$ (left) and around the 473 tagging jet midpoint (right) in the POWHEG NLO+PS scheme. In both distributions, the 474 three shower models produce similar results for $H_{\rm T} > 40 \text{ GeV}$, while in the complementary 475 region again only VINCIA and the dipole-improved PYTHIA shower agree. In this soft region, 476 the default PYTHIA shower again predicts more radiation than the other two. As before, 477 a variation of the POWHEG scale choice leads to significant effects in the predictions of 478 PYTHIA's default shower, but has only mild effects on the dipole-improved shower and 479 VINCIA. 480

481 3.3 Comparison of Matching and Merging

In figs. 11 to 13, we compare the VINCIA NLO-matched predictions presented in the last section to an $\mathcal{O}(\alpha_{\rm S})$ tree-level merged calculation using the CKKW-L scheme implemented for VINCIA. For the latter, we include the exclusive zero-jet and inclusive Sudakov-weighted 1-jet predictions in the plots (dashed lines).

The uncertainty bands of the merged predictions (labelled VINCIA MESS $O(\alpha_{\rm S})$) are obtained by a variation of the shower renormalisation scale as per section 2.2. As VINCIA's merging implementation reweights event samples by a ratio of the strong coupling as used in the shower to the one used in the fixed-order calculation, this variation effectively amounts to an intertwined scale variation of the hard process as well. The uncertainty bands of the NLO-matched calculation are obtained by the variation of the $p_{\perp,\text{hard}}$ scale as in the previous section.



Figure 11: Comparison between LO+PS, POWHEG NLO+PS, and CKKW-L-merged predictions for the transverse momentum of the first (*left*) and second (*right*) tagging jet.



Figure 12: Comparison between LO+PS, POWHEG NLO+PS, and CKKW-L-merged predictions for the transverse momentum (left) and pseudorapidity (right) of the third jet.



Figure 13: Comparison between LO+PS, POWHEG NLO+PS, and CKKW-L-merged predictions for the scalar transverse momentum sum for $|\eta| < 0.5$ (*left*) and around the pseudorapidity midpoint of the tagging jets (*right*).



Figure 14: Tree-level merged predictions with up to four additional jets for the pseudorapidity (*left*) and transverse momentum (*right*) of the Higgs and tagging jets system.

Taking into account their respective accuracies, we observe good agreement between 493 the matched and the merged predictions for the transverse momentum and pseudorapidity 494 spectra. We expect the small differences that are visible to trace back mainly to the lack 495 of unitarity in the CKKW-L scheme. This explanation is supported by the fact that the 496 merged calculation overshoots the matched ones and that e.g. for the p_{T,i_3} distribution, 497 the inclusive Sudakov-reweighted 1-jet contribution already agrees in shape and magnitude 498 with the matched distributions, while the exclusive zero-jet contributions only adds to the 499 rate, i.e overall normalisation. In addition, we wish to note again that the mismatch of the 500 POWHEG and VINCIA ordering variables is only treated approximately via the use of vetoed 501 showers, while the correct shower history is taken into account in the merged calculation. 502 Furthermore, we have used two different renormalisation and factorisation scales in the 503 two calculations. Because the renormalisation scale variation in VINCIA's merging affects 504 the renormalisation scale of the hard process, as alluded to above, the renormalisation 505 scale mismatch is covered to some degree by the scale variations in the merging. 506

The situation is different for the $H_{\rm T}$ distributions, cf. fig. 13. In the merged calculation, 507 more soft radiation is predicted in the central pseudorapidity region than in the matched 508 one. The distribution is solely governed by the one-jet sample there, while the zero-jet 509 sample contributes significantly above 60 GeV only. In the midpoint region, however, the 510 merged calculation predicts the same shape as the matched one, but with an overall bigger 511 rate. Barely any contribution stems from the exclusive zero-jet sample in this observable. 512 This confirms the properties of the two $H_{\rm T}$ observables mentioned in section 2.4. When 513 the observable is defined over the central rapidity region, it is sensitive to the radiation of 514 the third jet in the soft region, i.e. for $H_{\rm T} \lesssim 60 \text{ GeV}$, but becomes sensitive to the tagging 515 jets in the complementary hard region, i.e. above around 60 GeV. In contrast, defining 516 the observable over the region around the pseudorapidity midpoint of the two tagging jets 517 cleans it from almost all contributions stemming from the Born configuration (only a tiny 518 contribution from soft radiation off the Born survives). Due to this property, the latter of 519 the two definitions is particularly suited in the study of the radiation pattern regarding 520 its coherence. 521

The comparison of NLO matching and $\mathcal{O}(\alpha_{\rm S})$ tree-level merging provides a strong cross check of both methods.



Figure 15: Tree-level merged predictions with up to four additional jets for the scalar transverse momentum sum in the central (left) and midpoint (right) pseudorapidity region.

524 3.4 Merged with up to Four Jets

In addition to the one-jet merged calculation of the last section, we here present a tree-525 level merged calculation with up to four additional jets (i.e., 6 jets in total when counting 526 the tagging jets) using VINCIA'S CKKW-L implementation. We consider the effect of ad-527 ditional hard jets on the spectra of the pseudorapidity and transverse momentum of the 528 Higgs plus tagging jets system as well as the herein before mentioned scalar transverse 529 momentum sum in the two pseudorapidity regions. The uncertainty bands of the merged 530 calculation shown in the figures are obtained by a variation of the renormalisation scale 531 prefactors $k_{\rm R}$, c.f. section 2.2, in VINCIA's shower and merging, again effectively represent-532 ing a variation of the renormalisation scale in the hard process as well, cf. section 3.3. As 533 visible from fig. 15, the inclusion of additional hard jets does not change the pseudorapid-534 ity spectrum, but increases the rate of the transverse momentum spectrum in the high- $p_{\rm T}$ 535 region. This correction is exactly what is expected from a multi-jet merged calculation. 536 The dashed lines in fig. 15 represent the different multi-jet contributions to the merged 537 prediction. Again as expected, the Born sample dominates in the low- $p_{\rm T}$ region and the 538 one-jet sample in the region around 40 GeV, whereas higher multiplicities take over in 539 the harder regions above $\sim 70 \text{ GeV}$. It is worth highlighting, however, that, at least in 540 the region 70 GeV $\leq p_{\rm T} \leq 150$ GeV, the two-jet sample dominates with only sub-leading 541 corrections from the three- and four-jet samples. 542

Figure 14 shows the $H_{\rm T}$ distributions in the central and midpoint pseudorapidity re-543 gions defined in section 2.4. As for the one-jet merged prediction presented in section 3.3, 544 the high- $H_{\rm T}$ region is dominated by the Born sample, while for small $H_{\rm T}$, the samples with 545 additional jets define the shape. Although all samples with additional jets contribute to 546 the central $H_{\rm T}$ over the full shown spectrum, the three-jet sample (denoted 1 *j* in fig. 14) is 547 the dominant extra-jet sample everywhere. Above approximately 60 GeV, the Born sam-548 ple becomes the predominant one, highlighting again that this region is sensitive mainly 549 to the tagging jets. Corrections from the multi-jet merging are negligible there. 550

As before, the situation is different in the midpoint region between the two tagging jets (right-hand pane in fig. 14). There, the Born sample has almost no impact (< 5%) on the $H_{\rm T}$ distribution and the one-jet sample (denoted 1*j* in fig. 14) dominates in the



Figure 16: Detailed comparison of PYTHIA DGLAP and VINCIA LO+PS predictions at parton-level, hadron-level, and hadron-level plus MPI for the transverse momentum of the first tagging jet (*left*) and the second tagging jet (*right*).

region $\lesssim 70 \text{ GeV}$, while the two-jet sample (denoted 2j in fig. 14) does in the region 70 GeV $\lesssim H_{\rm T} \lesssim 100$ GeV. This emphasises the finding of the last section that the midpoint $H_{\rm T}$ is clean of contributions from the tagging jets and therefore more relevant in the study of coherence effects in QCD radiation.

558 3.5 Hadronisation and Multi-Parton Interactions

Although we focused on the parton level throughout this study, we wish to close by estimating the size of non-perturbative corrections arising from hadronisation, fragmentation, and multi-parton interactions. To this end, we employ PYTHIA's string fragmentation and interleaved MPI model [11] using the default PYTHIA [52] and VINCIA [27] tunes.

Figures 16 to 18 compare PYTHIA's simple shower and VINCIA predictions on the 563 parton level, hadron level, and hadron level with MPIs at LO+PS accuracy. As expected 564 from the cuts employed in our analysis, cf. section 2.4, the inclusion of non-perturbative 565 effects in either of the two simulations has only a negligible effect on most observables 566 studied here, although the discrepancy between the two showers is slightly mitigated. A 567 notable exception are the VINCIA predictions for the $H_{\rm T}$ in the two pseudorapidity regions 568 defined in section 2.4, for which the inclusion of MPIs leads to a substantial excess in 569 radiation in the soft region. This means, that in those regions the coherent suppression 570 of radiation by VINCIA is overwhelmed by the soft radiation off secondary (non-VBF-like) 571 interactions, at least with our set of cuts. It should be noted here that firstly, this excess 572



Figure 17: Detailed comparison of PYTHIA DGLAP and VINCIA LO+PS predictions at parton level, hadron level, and hadron-level plus MPI for the transverse momentum (*left*) and pseudorapidity of the third jet (*right*).

is not visible in the distributions obtained with PYTHIA's simple shower, and secondly,
the discrepancy between the simple shower and VINCIA overpowers the MPI effect greatly.
As such, the inclusion of hadron-level and MPI effects emphasise that VINCIA's antenna
shower reproduces QCD coherence effects more faithfully than PYTHIA's simple shower.

577 4 Conclusion

We have here studied the effect of QCD radiation in VBF Higgs production, focusing in 578 particular on how the coherent emission patterns exhibited by this process are modelled 579 by various parton-shower approaches that are available in the PYTHIA event generator, 580 and how significant the corrections to that modelling are, from higher fixed-order matrix 581 elements. From a QCD point of view, the main hallmark of VBF is that gluon emission in 582 the central region originates from intrinsically coherent interference between initial- and 583 final-state radiation. In DGLAP-style showers, which are anchored in the collinear limits 584 and treat ISR and FSR separately, this interplay can only be captured at the azimuthally 585 integrated level via angular ordering, while it is a quite natural element in dipole- and 586 antenna-based formalisms, in which initial-final colour flows enter on an equal footing with 587 final-final and initial-initial flows. Hence we would expect the latter (dipole/antenna-style) 588 approaches to offer more robust and reliable modelling of the radiation patterns in VBF 589 than the former (DGLAP-based) approaches. 590

Figure 18: Detailed comparison of PYTHIA DGLAP and VINCIA LO+PS predictions at parton level predictions for the central $H_{\rm T}$ (*left*) and midpoint $H_{\rm T}$ (*right*).

To this end, we have compared the VINCIA antenna shower to PYTHIA's default ("sim-591 ple") shower, including both its (default) DGLAP and its dipole-improved option ("dipole 592 recoil"). We have shown that at leading order, large discrepancies pertaining to the radi-593 ation of additional jets in the central rapidity regions exist between the default PYTHIA 594 predictions and the ones obtained with the dipole option and VINCIA, while the latter two 595 appear more consistent. This effect even concerns observables related to the tagging jets, 596 i.e. those jets which are described by the matrix element and not the shower. We have 597 confirmed that these findings apply to both external (LHA) and internal events. 598

After matching the showers to the NLO, these discrepancies mostly vanish for observ-599 ables sensitive to the tagging jets or third jet only, while larger effects remain visible in 600 observables sensitive to higher jet multiplicities. These findings are largely consistent with 601 the ones from an earlier study [8], although it is worth highlighting that the disagreement 602 found for the default PYTHIA shower is fairly less pronounced here after matching it to the 603 NLO via the POWHEG scheme. We consider this to be an effect of a more careful treatment 604 of the ordering-variable mismatch between POWHEG and PYTHIA. Based on this, we rec-605 ommend varying the POWHEG:pThard mode contained in the PowhegHooks classes to gain 606 an estimate of systematic matching uncertainties. To reduce the uncertainties pertaining 607 to the use of vetoed showers with POWHEG samples, a truncated and vetoed shower should 608 be used with both PYTHIA and VINCIA. As alluded to above, such a scheme is not (yet) 609 available for either of the showers considered in the present study. 610

In addition to NLO matching, we have studied the effect of including higher-multiplicity tree-level matrix elements in the shower via the CKKW-L merging scheme in VINCIA. We

have confirmed that the NLO-matched and one-jet merged calculations lead to comparable 613 predictions for observables sensitive to the third jet. For a set of inclusive observables, 614 we presented predictions from a tree-level merged calculation at $\mathcal{O}(\alpha_s^4)$. This yields cor-615 rections of the order of 20% in the hard tail above around 60 GeV of the transverse 616 momentum spectrum of the Higgs-plus-tagging-jet system. Considering the mild correc-617 tions in the ranges studied here, it is evident that the sample with four additional jets (i.e. 618 the 2+4-jet sample) will contribute significantly only in the very hard tails $H_{\rm T} \gg 100 {\rm ~GeV}$ 619 and $p_{\perp,\mathrm{H}ii} \gg 150 \text{ GeV}$. 620

Although not the main focus of this study, we have gained a first estimate of nonperturbative corrections on the observables studied here. While we generally found only minor changes from the inclusion of hadron-level corrections, the inclusion of MPIs had a relatively more significant effect on VINCIA's predictions than on the ones obtained with PYTHIA's default shower. This affected the rate of radiation in soft as well as central pseudorapidity regions, i.e. precisely the regions in which VINCIA predicts a strong coherent suppression, so that the MPI contamination becomes relatively more important.

With this study we also proposed two new observables, the scalar transverse momentum 628 sum in the central pseudorapidity region and around the pseudorapidity midpoint between 629 the two tagging jets. We have shown that both of these observables are sensitive to multi-630 jet radiation, but highlighted that the former becomes dominated by the tagging jets in the 631 hard region $H_{\rm T} \gtrsim 60$ GeV. As an alternative, we demonstrated that the $H_{\rm T}$ sum around 632 the midpoint between the tagging jets is free of this contamination, with the Born sample 633 only giving a negligible contribution. Due to the strong suppression of radiation in this 634 region, both observables do however receive corrections from the modelling of multi-parton 635 interactions, which would be relevant to study further. 636

While it has been considered a coherent shower before, this has been the first time that the radiation pattern of the VINCIA antenna shower was studied with a dedicated focus on its coherence. At the same time, we have here showcased NLO matching and tree-level merging methods with VINCIA, which are both publicly available as of the PYTHIA 8.306 release.

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652 A POWHEG+VINCIA Setup

As mentioned in section 2.3.1, a dedicated vetoed-shower UserHook for POWHEG+VINCIA was developed as part of this work and is included in the standard PYTHIA distribution from version 8.306 onwards. At the time of submission of this manuscript, it is included in the file PowhegHooksVincia.h, in the directory include/Pythia8Plugins/, which also contains the standard PowhegHooks.h file. (Note that these two files may be merged into one in a future release; if so, simply omit the corresponding step below.)

Assuming you have a main program that is set up to run POWHEG+PYTHIA (such as the example program main31.cc included with PYTHIA), the following changes (highlighted in red) will modify it to run POWHEG+VINCIA:

662 663 664 665	 Include the PowhegHooksVincia.h header file: #include "Pythia8Plugins/PowhegHooksVincia.h" (you can leave any existing #include "Pythia8Plugins/PowhegHooks.h" state- ment; the two will not interfere with each other).
666 667 668 669	 Replace the POWHEG+PYTHIA user hook pointer by a POWHEG+VINCIA one: shared_ptr<powheghooks> powhegHooks; powhegHooks = make_shared<powheghooksvincia>(); pythia.setUserHooksPtr((UserHooksPtr)powhegHooks);</powheghooksvincia></powheghooks>
670	In addition, the following settings should be used:
671 672 673	 Switch on VINCIA's showers and allow them to fill all of phase space: PartonShowers:model = 2 # Use Vincia's shower model. Vincia:pTmaxMatch = 2 # Power showers (to be vetoed by hook).
674 675	• Enable shower vetoes via the PowhegHooksVincia (same as for PowhegHooks): POWHEG: veto = 1 # Turn shower vetoes on
676 677 678 679 680 681	 Turn QED/EW showers and interleaved resonance decays off: Vincia:ewMode = 0 # Switch off QED/EW showers. Vincia:interleaveResDec= off # No interleaved resonance decays. While enabling QED showers (Vincia:ewMode = 1 2) should not pose any problems in the matching, it is not validated (yet). We recommend against using the EW shower (Vincia:ewMode = 3) with the POWHEG matching.
682 683 684 685 686 687	 Since POWHEG-BOX event samples come unpolarised, VINCIA's helicity shower should be turned off (the helicity shower needs a polarised Born state): Vincia:helicityShower = off # Use helicity-averaged antennae. We note that VINCIA offers the possibility to polarise Born configurations using matrix elements provided via interfaces to external generators. We have not studied this in the present work.
688 689 690 691	 In the POWHEG-specific settings, the number of outgoing particles in the Born process is defined as usual, e.g. =2 for the 2 → 2 example in main31.cc, or =3 for the 2 → 3 VBF-type processes studied in this work: POWHEG:nFinal = 3 # Number of outgoing particles in the Born process.
692 693 694 695	• We highly recommend varying the POWHEG:pThard mode, for both PYTHIA and VINCIA, to estimate matching systematics. This is how the shaded bands in most of the plots shown in this paper were obtained. POWHEG:pThard = 2 # Vary (=0,=1,=2) to estimate matching systematics.

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696 697	• We also recommend checking all accepted emissions rather than only the first lew: POWHEG:vetoCount = 10000
698 699 700 701 702	 The following settings are simply left at their recommended values (the same as for main31.cmnd); see the onlin manual section on POWHEG for details: POWHEG:pTemt = 0 POWHEG:emitted = 0 POWHEG:pTdef = 1
703 704 705 706	 For completeness, (we note that we have anyway turned both MPI and QED showers off in this study): POWHEG:MPIveto = 0 POWHEG:QEDveto = 2
707 708 709	The event files generated by POWHEG should be provided in exactly the same way as for PYTHIA+POWHEG. If the POWHEG events were generated in several separate batches, for instance, the resulting files can be read as usual, using PYTHIA's "subruns" functionality:
710 711 712	<pre>! Powheg Subruns. Beams:frameType = 4 Main:numberOfSubruns = 3</pre>
713 714 715 716	Main:subrun = 0 Beams:LHEF = POWHEG-BOX-V2/VBF_H/run/pwgevents-0001.lhe
717 718 719 720	Main:subrun = 1 Main:LHEFskipInit = on Beams:LHEF = POWHEG-BOX-V2/VBF_H/run/pwgevents-0002.lhe
721 722 723	Main:subrun = 2 Main:LHEFskipInit = on Beams:LHEF = POWHEG-BOX-V2/VBF_H/run/pwgevents-0003.lhe

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1

724 B VINCIA CKKW-L Setup

Since PYTHIA version 8.304, the release is shipped with VINCIA's own implementation of
 the CKKW-L merging technique, suitably modified for sector showers.

In the spirit of the last section, let us again assume you have a main program running CKKW-L merging with PYTHIA's default ("simple") shower. (We note that this is a hypothetical setup for the purpose of this study, as the default merging implementation in PYTHIA 8.3 does not handle VBF processes. An algorithmic fix is planned for PYTHIA version 8.307 or later.) The following changes are needed to alter it to run VINCIA's CKKW-L merging instead, with changes again highlighted in red.

```
Turn VINCIA and its sector showers on<sup>6</sup>:
PartonShowers:model = 2 # Use Vincia's shower model.
Vincia:sectorShowers = on # Turn sector showers on.
Disable VINCIA components that are not (yet) handled by the merging:
Vincia:ewMode = 0 # Switch off QED/EW showers.
```

⁶We note that as of now, sector showers are on per default in VINCIA and this flag is listed here only for completeness.

738	VINCIA:InterleaveresDec= off # No interleaved resonance decays.
739	Vincia:helicityShower = off # Use helicity-averaged antennae.
740	These three limitations are intended to be temporary and may be lifted in future
741	updates; users are encouraged to check for changes mentioning VINCIA's merging
742	implementation in the Update History section of PYTHIA's HTML manual in releases
743	from 8.307 onwards.
744	Enable the merging machinery and set the merging scale definition (in this study,
745	all event samples were regulated by a $k_{\rm T}$ cut, so $k_{\rm T}$ -merging is turned on):
746	Merging:doMerging = on # Turn merging machinery on.
747	Merging:doKTMerging = on # Set kT as merging scale.
748	Set the merging scale to the desired value in GeV (note that the cuts on the event
749	samples should be more inclusive than the ones in the merging!):
750	Merging:TMS = 20 # Value of the merging scale in GeV.
751 •	Replace the Process string by one obeying VINCIA's syntax, i.e. encased in curly
752	brackets and with whitespaces between particles, and switch the dedicated VBF
753	treatment on:
754	Merging:process = { p p > h0 j j } # Define the hard process.
755	Vincia:mergeVBF = on # Enable merging in VBF systems.
756	Set the number of additional jets with respect to the Born process (e.g. for the VBF
757	process considered here, the number of additional jets is 4, while the total number
758	of jets is 6):
759	Merging:nJetMax = 4 # Merge samples with up to 4 additional jets.

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