

Development of a cold atomic muonium beam for next generation atomic physics and gravity experiments

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Abstract

A high-intensity, low-emittance atomic muonium ($M = \mu^+ + e^-$) beam is being developed, which would enable improving the precision of M spectroscopy measurements, and may allow a direct observation of the M gravitational interaction. Measuring the free fall of M atoms would be the first test of the weak equivalence principle using elementary antimatter (μ^+) and a purely leptonic system. Such an experiment relies on the high intensity, continuous muon beams available at the Paul Scherrer Institute (PSI, Switzerland), and a proposed novel M source. In this paper, the theoretical motivation and principles of this experiment are described.

31.1 Introduction

Muonium (M) is a two-body exotic atom consisting of a positive anti-muon (μ^+) and an electron (e^-). This purely leptonic system can be a unique precision probe to test bound-state QED without the influence of nuclear- and finite size effects. Laser spectroscopy of the M 1S-2S transition [1, 2], and microwave spectroscopy of the M ground state hyperfine structure [3] provided precision measurements of fundamental constants (muon mass, magnetic moment), while searches for muonium-antimuonium conversion put limits on the strength of charged lepton number violation [4]. Improvements in these measurements especially 1S-2S spectroscopy is strongly motivated by recent experiments measuring the anomalous muon $g-2$ [5]. A high intensity, cold atomic beam could significantly improve statistical limitations and systematic effects originating from the (residual) Doppler shift.

Another unique and so far unexplored facet of M is that its mass is dominated by the μ^+ , which is not only an elementary antiparticle, but also a second-generation lepton. Direct measurement of the gravitational interaction, thereby tests the weak equivalence principle of such particles, has not yet been attempted [6, 7]. Besides muonium, only antihydrogen ($\bar{H} = \bar{p} + e^+$) [8–10] and positronium ($\text{Ps} = e^- + e^+$) [11–13] have been proposed as laboratory candidates for antimatter gravity experiments, and M is the only viable candidate for testing gravity with purely leptonic, second generation matter.

31.1.1 The weak equivalence principle

The Standard Model (SM), as any local, Lorentz-invariant quantum field theory, incorporates CPT symmetry - the simultaneous transformations of charge conjugation (C) parity transformation (P) and time reversal (T) - as an exact symmetry [14]. An important consequence

of this is the equivalence of various measurable properties of matter and antimatter, such as the mass, the magnitude of the charge, and the strength of certain interactions. Comparative measurements between matter and antimatter put stringent limits on CPT violation by different experiments using mesons ($K_0 - \bar{K}_0$) [15] leptons ($e^+ - e^-$, $\mu^+ - \mu^-$) [16,17] and baryons ($p - \bar{p}$) [18–21].

With the lack of a unified theory of General Relativity (GR) and the SM, the considerations above however do not imply anything about the gravitational interaction of matter and antimatter. Our expectations originate from the assumed equivalency of the inertial and gravitational masses of particles, which is incorporated in GR as part of the equivalence principle [22, 23]. The exact formulation of this principle varies in the literature, and frequently cited as a collective of some these statements below:

1. Weak equivalence principle (WEP) or *universality of free-fall*: all particles (and antiparticles) fall with the same acceleration in a gravitational field.
2. Local position invariance (LPI): The outcome of any local non-gravitational experiment is independent of its location in space or time. Experimental consequences:
 - (a) the *universality of clocks* (WEP-c), meaning all systems regardless of their composition (e.g. matter or antimatter) experience the same local time.
 - (b) the lack of *variation of fundamental constants* (WEP-v) in time.
3. Local Lorentz invariance (LLI): The outcome of any local non-gravitational experiment in a free-falling laboratory is independent of its velocity.
4. Strong equivalence principle (SEP): states LLI and LPI combined and extended to the gravitational measurements as well (e.g. test bodies with significant contributions from their own gravitational field.)

The combination of the above weak statements (LLI with LPI, sometimes WEP included) is frequently referred to as Einstein’s equivalence principle. Most importantly, violation of one of these principles would not necessarily mean the violation of all, and depending on the underlying new physics, it would effect GR and the SM on different levels [23, 24]. Hence, testing the above equivalence principles independently in different experiments using different SM particles is essential [22, 23, 25].

For example, in Earth-based or satellite-borne laboratories, gravitational redshift experiments (WEP-c) and direct free-fall experiments (WEP) using different types of matter may be considered. WEP-c was tested to relatively high accuracy ($\Delta g/g < 10^{-6}$) using matter and antimatter clocks, H and \bar{H} [18, 24] as well as by measuring cyclotron frequencies of trapped p and \bar{p} [?, 19]. Such experiments arguably also constrain direct WEP-violation originating from certain SM extensions [24, 26]. However, direct gravitational free-fall experiments (tests of the WEP) have never been carried out using anything other than normal matter, more precisely macroscopic objects of different material composition, neutral atoms or neutrons.

31.2 Experiments for testing the WEP

The most rigorous tests of the WEP utilize Earth-based and satellite-borne experiments that either use the modern versions of the Eötvös torsion pendulum, or other sensitive accelerometers. These experiments compare gravitational accelerations of two macroscopic test masses (g_1, g_2) in terms of the Eötvös parameter

$$\eta(1, 2) = 2 \frac{|g_1 - g_2|}{|g_1 + g_2|}. \quad (31.1)$$

82 The highest precision comes from the satellite-borne MICROSCOPE experiment [27] for tita-
 83 nium and platinum, giving $\eta(\text{Ti, Pt}) = [1 \pm 9(\text{stat}) \pm 9(\text{syst})] \times 10^{-15}$, which is about an order
 84 or magnitude better than the best torsion pendulum results from the Eöt-Wash group [28]. On
 85 the largest mass scales, the Lunar Ranging Test is the most notable, constraining differences be-
 86 tween the Earth and Moon gravitational and inertial mass ratios to levels below $\sim 10^{-13}$ [29].

87 The WEP has been tested on the atomic scales as well. The latest atom interferometry
 88 results comparing two isotopes of rubidium in free-falling cold atom clouds confirmed a null
 89 measurement with $\eta(^{85}\text{Rb}, ^{87}\text{Rb}) = [1.6 \pm 1.8(\text{stat}) \pm 3.4(\text{syst})] \times 10^{-12}$ [30].

90 Gravitational acceleration has only been observed with one subatomic particle, the neu-
 91 tron. The most precise experiments were carried out using neutron refractometers [31], neu-
 92 tron spin-echo technique [32] and also the gravitational quantum states of ultracold neu-
 93 trons [33, 34]: they have reached an overall precision of $\sim 0.3\%$. New experiments plan to
 94 improve this by at least an order of magnitude [35].

95 In summary, WEP tests have limited the Eötvös parameter to $\eta < 1.3 \times 10^{-14}$ for different
 96 (macroscopic) elements. Future satellite-borne experiments may improve the precision by two
 97 orders of magnitude [23, 36].

98 31.2.1 Possibilities for new physics violating WEP in exotic atoms

99 Conservative extensions of the SM and GR that would differentiate matter and antimatter
 100 in a free fall experiment were discussed with the specific case of antihydrogen [24]. The
 101 possibilities discussed include extensions of the existing theories like Kostelecký's extension
 102 of the SM [37] containing Lorentz- and CPT violating terms, or minimal modifications of GR
 103 that would maintain core principles (like local Lorentz invariance, causality, description as
 104 a Riemannian manifold) but modify the dynamics described by the action by adding extra
 105 terms that modify the energy-momentum tensor. Several possibilities of 'fifth force' scenarios
 106 have also been discussed in the literature, most recently in e.g. Refs. [38, 39]. In Ref. [24] it
 107 is pointed out how such new vector bosons could have different couplings to the oppositely
 108 charged matter and antimatter, and how this would impact WEP measurements.

109 The resulting theoretical possibilities are narrow, especially in light of existing WEP mea-
 110 surements on ordinary matter that arguably constrain effects of antimatter gravity via the core
 111 principles above and the potential and kinetic energies incorporated in the rest mass [26], and
 112 WEP-c measurements that already set constraints on GR extensions [24]. The overall conclu-
 113 sion from theory is that while possible violations of WEP in antihydrogen free-fall experiments
 114 may be envisaged, present viable models that do not break the principles of the GR or SM
 115 suggest that they are small, and almost certainly already constrained with WEP-c experiments
 116 at the $\Delta g/g < 10^{-6}$ level [24]. This consideration also applies to the proposed positronium
 117 experiments [11–13] that would probe the antimatter counterpart of the electron.

118 The same considerations however do not necessarily apply to muonium, which contains
 119 an elementary antiparticle from the second generation (μ^+). Direct gravitational tests have
 120 never been carried out before neither with μ^+ nor μ^- . Hence, we may not need to envision
 121 long-range vector bosons (fifth forces) that differentiate matter and antimatter to explain an
 122 unexpected result, but could explore other new physics that couples differently to muons than
 123 electrons. In the light of recent precision experiments that show intriguing discrepancies in
 124 the charged lepton sector like the muon $g-2$ anomaly [5] or the B anomalies [40], such exotic
 125 BSM physics may not be so far fetched.

126 As to WEP-c tests, next generation experiments of the 1S-2S transition frequency of M
 127 have the capability of reaching ~ 0.1 ppm fractional precision, and of being sensitive to the
 128 effects of gravitational redshift change while the laboratory travels in the solar system (annual
 129 modulations of the gravitational potential in perihelion-aphelion) [41]. The interpretation of

130 the muon g-2 result as a clock measurement [5, 41] may also bring some intriguing hints in
 131 the same direction.

132 We also note that there has been an ambiguity in interpreting what experiments with com-
 133 posite objects like neutrons or neutral atoms already tell us about the connection of gravity to
 134 the SM particles and interactions [26, 41]. About 99 % of the rest mass of protons and neutrons
 135 comes from the strong interaction that confines the constituent quarks. Nuclear binding- and
 136 kinetic energies further shift the mass up to $\sim 9 \text{ MeV}/c^2$ per nucleon, while electrostatic in-
 137 teractions with another few eV/c^2 . In this sense, direct gravity experiments have so far tested
 138 mainly binding energies from the strong interaction.

139 However, the mass of the muonium is dominated by the elementary muon mass, which is
 140 a fundamental parameter in the SM. Hence measuring muonium gravity may provide cleaner
 141 access to understanding the connection of gravity to elementary particles in the absence of an
 142 overwhelming strong interaction.

143 31.3 Prospects for a gravity experiment with a novel M beam

144 A direct gravity experiment using muonium is inherently challenging due to the short lifetime
 145 ($\tau \sim 2.2 \mu\text{s}$) of the μ^+ and the fact that M atoms must be created in matter, while experiments
 146 must be carried out *in vacuo*. These imply that we need to envision experiments using propa-
 147 gating atomic beams. A straightforward method is to use atom interferometry, which is known
 148 to be a sensitive method to observe inertial forces [30]. However, this requires ultracold atomic
 149 clouds, or well-collimated atomic beams with small transverse momentum.

150 Present vacuum muonium sources are room temperature, porous materials that allow com-
 151 bination of the muon with an electron from the bulk, and a following quick diffusion inside
 152 the nanoscopic pores (See Figure 31.1 A). Laser ablated silica aerogel is one of the best room
 153 temperature converters; the microscopic holes created by the laser enhance the emission of
 154 the M atoms into vacuum. Such sources provide $\sim 3\%$ muon-to-vacuum M conversion using
 155 surface μ^+ beams of 28 MeV/c momentum [42]. However, such converters produce a M beam
 156 with broad (thermal) energy and angular ($\sim \cos\theta$) distributions.

157 Mesoporous materials have been shown to convert μ^+ to vacuum M with efficiencies of
 158 40% at room temperature when using a highly moderated, keV energy μ^+ beam; this has an
 159 intensity four orders-of-magnitude lower than a surface muon beam. These low-energy muons
 160 penetrate only a few μm into the surface, but are emitted with wide energy- and angular
 161 distributions [43]. Improving the source quality by cooling these samples results in lower
 162 emission rates, with no observable emission below $\sim 50 \text{ K}$ due to the decreased diffusion
 163 constant, and the sticking of M to the pore walls that occurs unavoidably with any conventional
 164 M converter [43, 44].

165 31.3.1 Vacuum muonium from superfluid helium

166 Superfluid helium (SFHe) may overcome the above mentioned difficulties due to its inert na-
 167 ture that rejects impurities from its bulk even at the lowest temperatures. This can be qualita-
 168 tively explained by the unusually small mean distance ($\sim 0.3 \text{ nm}$) of the condensed He atoms:
 169 when implanting large impurity atoms or negative ions, nearby He atoms will be repelled by
 170 the Pauli core repulsion [45], resulting in a spherical cavity (bubble) around the impurity. This
 171 exercises an inward pressure that results in a positive chemical potential of M, that results in
 172 the ejection of the impurity from the bulk when they reach the surface.

173 The principle of the proposed M source relying on this mechanism [6, 46] is summarized in
 174 Figure 31.1 (b). The μ^+ are stopped in the bulk of SFHe, where they capture an electron from
 175 the ionization trails. The M atom formed in the bubble state (M^*) diffuses to the surface where

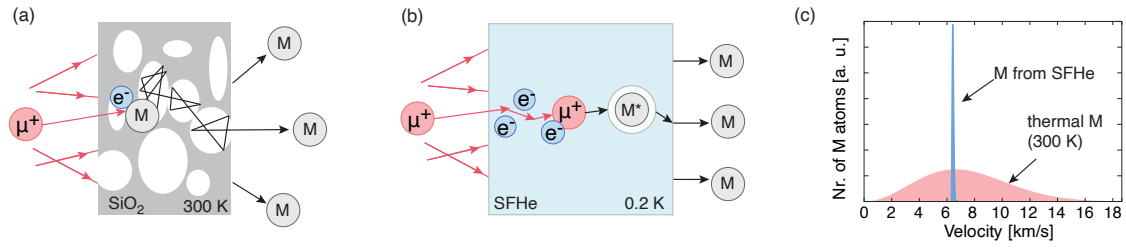


Figure 31.1: (a) Principle of a conventional μ^+ -to-vacuum-M converter based on porous materials. (b) Principle of a SFHe-based converter. (c) Comparison of the expected M velocity distribution from SFHe (blue) and a mesoporous (red) converters.

176 it will be emitted perpendicularly, with kinetic energy defined by the chemical potential, only
 177 slightly broadened by thermal energies (Figure 31.1 (c)).

178 The chemical potentials for ^4He , ^3He , H, D and T in SFHe have been calculated [47, 48],
 179 and these predictions have been experimentally verified for ^4He , ^3He and D [49]. Modelling
 180 M atoms as a light hydrogen isotope gives an approximate chemical potential of $E/k_B \approx 270$ K
 181 [50], implying that the M atom will leave the SFHe surface with a well defined longitudinal
 182 velocity of $v_M \sim 6300$ m/s. The velocity spread and the transverse velocities are given in first
 183 approximation by the thermal motion of the M^* bubble in the liquid. Predicting this is difficult
 184 without a microscopic theory of the quantum liquid.

185 Based on [47], the M^* acquires an effective mass of $m_M^* \approx 2.5 m_{\text{He}}$ due to hydrody-
 186 namic back-flow effects in SFHe, similar to all hydrogen isotopes [50]. In a simplified model,
 187 the M^* loses energy in a 200 mK isotopically-pure superfluid ^4He solely by creating rotons
 188 and phonons (no scattering on ^3He), until its kinetic energy falls below the roton gap [51]
 189 ($\Delta_{\text{rot}}/k_B = 8.6$ K), resulting in thermal velocities distributed below $v_t \approx 110$ m/s. Thermally
 190 available phonons are sparse at this temperature, hence scattering on phonons is unlikely on
 191 the relevant μs timescales [52]. The small effective mass of the M^* suggests we can neglect
 192 other hydrodynamic effects like vortex nucleation as well [53], and assume that M^* moves
 193 afterwards ballistically in the SFHe medium, with average velocities of $\bar{v}_t \approx v_t/2$. This allows
 194 a large fraction of the atoms to escape from $\sim 100 \mu\text{m}$ thick SFHe layers, a thickness that can
 195 efficiently stop μ^+ beams of 10-12 MeV/c momentum.

196 In summary, with the assumptions above and neglecting further surface effects, we expect
 197 efficient muon-to-vacuum-M ($\sim 10\text{--}30\%$) conversion with a mean atomic velocity of $v_M \approx 6.3$
 198 mm/ μs in the longitudinal direction (originating from the chemical potential), and a spread
 199 given approximately as $v_t \approx 0.11$ mm/ μs from the thermal velocities above. This yields to
 200 a momentum bite of $< 0.01\%$, and $\alpha \approx v_t/v_M \approx 17$ mrad angular distribution. Moreover,
 201 the cold temperature of the SFHe (~ 200 mK) leads to a small saturated vapor density
 202 (equivalent to UHV conditions at room temperature) which is needed to reduce the collision
 203 of the vacuum M with the He gas that would degrade the quality of the M beam.

204 We have constructed a 200 mK cryogenic target cooled by a dilution refrigerator for the
 205 first proof-of-principle experiments to test the above theoretical assumptions, and presently
 206 carrying out the first measurements at PSI [54].

207 31.3.2 Free fall experiment using M-atom interferometry

208 If the M atoms are initially at rest in the vertical direction and obey the weak equivalence
 209 principle, they fall a mere $\Delta x = \frac{1}{2}gt^2 = 600$ pm in a time of $t = 5\tau$. The measurement of
 210 this tiny gravitational fall needs precise knowledge of the initial momentum of the atoms, and
 211 requires strict momentum selection. Two periodic gratings (G1 and G2) with horizontal slits
 212 of pitch d and spaced by a distance L could be used to achieve this momentum selection as

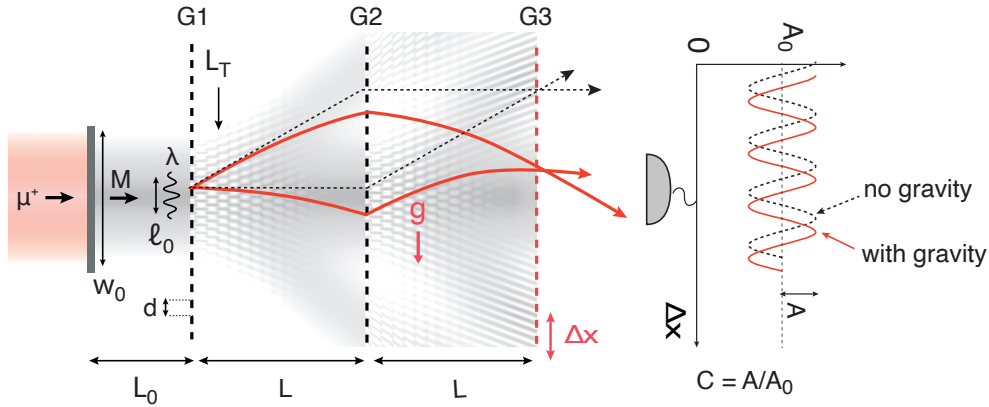


Figure 31.2: A three-grating interferometer used to measure the gravitational interaction of M atoms. The quantum diffraction pattern caused by the gratings G1 and G2 with a fully coherent beam is given in grey. Classical trajectories (red and dashed lines) are shown to illustrate the effect of gravity on the measured interference pattern appearing at G3. The vertical shift of the interference pattern caused by the gravitational acceleration g is detected by measuring the transmitted M rate while scanning G3 in vertical direction. See details in text.

213 shown in Figure 31.2.

214 The classical and quantum regime of this device is characterized by de Broglie wavelength
 215 of the atoms, $\lambda = h/p$, and grating pitch d in terms of the Talbot length, $L_T = d^2/\lambda$, which is
 216 approximately 18 microns for thermal M atoms with $\lambda_M \approx 0.56$ nm. If the grating distances
 217 are much smaller than the Talbot length ($L \ll L_T$, the diffraction of the atoms can be neglected
 218 during propagation in the device, and this classical device is called a Moiré deflectometer. With
 219 the choice of much smaller grating pitch or larger distances $L \gg L_T$ diffraction and in general
 220 the wave nature of the atoms become significant, and we work on an interferometer.

221 With both classical and quantum cases, trajectory selection at G1 and G2 will result in an
 222 intensity pattern with the same periodicity d at a distance L after G2. Gravitational accelera-
 223 tion and deflection of the atoms causes a phase shift $\delta\phi$ of this pattern in the vertical direction
 224 as $\delta\phi = 2\pi g T^2/d$, where $T = L/v_M$ is the M time of flight between each pair of gratings.

225 Direct observation of this sub-micron patterns and sub-nanometer shifts needed for mea-
 226 suring M gravity would be extremely hard. It is possible however to carry out an indirect
 227 measurement using a third grating (G3) of the same pitch d , placed at distance L from G2.
 228 By counting the total rate of M atoms transmitted through G3 as a function of the G3 vertical
 229 position Δx the phase shift can be measured.

230 The contrast of the intensity pattern C is defined by the ratio of the amplitude and the
 231 average yield $C = A/A_0$ as shown in Figure 31.2. When the three gratings work as an inter-
 232 ferometer, this contrast strongly depends on the transverse coherence length of the beam, ℓ_0 ,
 233 that determines how many slits of G1 are illuminated with a coherent wavefront. This coher-
 234 ence length in relation to the beam width w_0 and the interferometer parameters (the grating
 235 periodicity d and distances L) together with the de Broglie wavelength (λ) of the atoms is
 236 sufficient to estimate to describe the interferometer performance in the first approximation.
 237 In analogy to statistical optics (Van Cittert-Zernike theorem [55]), we can relate the trans-
 238 verse coherence length of the M beam to the transverse momentum distribution of the atoms:
 239 $\ell_0 = \frac{1}{2} \frac{\lambda}{\alpha} \approx 16$ nm, where α is the above mentioned angular spread of the M source. This
 240 initial transverse coherence is naturally increasing as the atoms experience diffraction on the

241 first grating. In simplified terms, diffraction results in a new coherent wavefront, that expands
 242 along the angle of diffraction. Regardless whether the 3-grating device works in the classical
 243 regime or as an interferometer, the sensitivity in measuring the gravitational acceleration g is
 244 given by [56]

$$\Delta g = \frac{1}{2\pi T^2} \frac{d}{C\sqrt{N}}, \quad (31.2)$$

245 where N is the number of M atoms transmitted through G3 and measured by the detector
 246 given by

$$N = N_0 \varepsilon_0 e^{-(t_0+2T)/\tau} (T_G)^3 \varepsilon_{\text{det}}, \quad (31.3)$$

247 with N_0 being the number of M atoms produced at the M source, and ε_0 the M transport
 248 efficiency from the source to G1. The M decay is accounted for by the third term $e^{-(t_0+2T)/\tau}$,
 249 where t_0 is the time of flight from the source to G1. The number of detected M atoms is
 250 further reduced by the M detection efficiency ε_{det} , and by the limited transmission T_G of a
 251 single grating. The short lifetime of the muon necessitates a gain in sensitivity by using a
 252 small grating pitch d . Maximal sensitivity, as a tradeoff between phase shift $\delta\phi$ and statistics
 253 N , is obtained for $T \approx 6 - 8 \mu\text{s}$ corresponding to an interferometer length of 40-50 mm.

254 A calculation of the interferometer parameters to extract the contrast C , uses an approxi-
 255 mation of the M source with a Gaussian Schell-model beam [57], and adapted mutual intensity
 256 functions that are widely used to describe the propagation of partially coherent light [55]. Us-
 257 ing realistic parameters on the initial beam size and quality expected from the superfluid source
 258 above, the fringe contrast of $C \approx 0.3$ at the exact position of G3 can be achieved. The contrast
 259 in this three-grating setup is less sensitive to the beam quality, but the sensitivity of the high
 260 contrast region along the propagation axis is, and shrinks to few μm . Such a measurement thus
 261 requires precise G3 positioning with μm -accuracy in the optical axis, and below-nm-accuracy
 262 in the vertical direction.

263 From (31.2) we see that determining the sign of g (more precisely to reach $\Delta g/g = 1$)
 264 in about one day, requires the detection of 3.2 M/s, assuming a contrast $C = 0.3$. Following
 265 (31.3), and taking pessimistic estimates from Monte Carlo simulations and initial detector
 266 and grating studies studies by using $T_G = 0.3$, $\varepsilon_0 = 0.75$ and $\varepsilon_{\text{det}} = 0.3$, at the source we
 267 need $N_0 \approx 1.4 \times 10^4$ M/s. As a comparison the πE5 beam line at PSI can presently deliver
 268 $3.6 \times 10^6 \mu^+/s$ at a momentum of 10 MeV/c within a transverse area of about 400 mm². At this
 269 muon momentum we can expect a muon-to-vacuum-M conversion efficiency of about 0.1-0.3
 270 based on the above discussion. This will result in M rates of up to $\sim 1.1 \times 10^6$ M/s. These
 271 high rates may allow a further collimation of the M beam to a 5×1 mm area, which would put
 272 less strain on grating production and alignment and would cut the number of useful M atoms
 273 conservatively by a factor $5 \text{ mm}^2/400 \text{ mm}^2 = 0.013$. Using these parameters where there is
 274 room for contingency, we expect to produce the necessary rate of $\sim 5 \times 10^4$ M/s in an small
 275 area of $\sim 5 \times 1 \text{ mm}^2$, and reach the goal sensitivity of $\Delta g = \frac{9.8 \text{ m/s}^2}{\sqrt{\# \text{ days}}}$ with present μ^+ sources.
 276 An increase by two orders of magnitude in μ^+ rates expected by the proposed HIMB (High
 277 Intensity Muon Beam [58]) project at PSI will further improve the sensitivity of to g .

278 31.4 Summary and outlook

279 With the development of a novel, cold atomic M beam with high yields of $10^4 - 10^5$ M/s and an-
 280 gular divergence of $\alpha \sim 10 - 20$ mrad, direct measurement of the gravitational acceleration of
 281 M seems feasible on a $\Delta g/g = 10^{-2}$ level of precision. While this precision is not comparable
 282 to present tests of the equivalence principle using normal matter ($\Delta g/g < 10^{-15}$), this exper-
 283 iment would be the first direct free fall demonstration using second generation (anti)matter.
 284 Moreover, the purely leptonic content of the atom would make it possible to study gravity for
 285 the first time in the absence of large binding energies from the strong interaction.

286 We are presently carrying out feasibility studies, and developing the first prototype of the
287 cryogenic atomic source and the accompanying detector system needed for this experiment at
288 PSI. We are also investigating further theoretical aspects using realistic M beams, and work-
289 ing on production methods for the 100-nm-pitch M interferometer and stabilization methods
290 needed for this precision.

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