# Bose-Einstein condensate in an elliptical waveguide 

Luca Salasnich<br>Dipartimento di Fisica e Astronomia "G. Galilei", Università di Padova, via F. Marzolo 8, I-35131, Padova, Italy<br>Istituto Nazionale di Fisica Nucleare, Sezione di Padova, Via Marzolo 8, 35131 Padova, Italy<br>Istituto Nazionale di Ottica del Consiglio Nazionale delle Ricerche, Unita di Sesto Fiorentino, Via Nello Carrara 2, 50019 Sesto Fiorentino, Italy

March 8, 2022


#### Abstract

We investigate the effects of spatial curvature for an atomic Bose-Einstein condensate confined in an elliptical waveguide. The system is well described by an effective 1D GrossPitaevskii equation with a quantum-curvature potential, which has the shape of a doublewell but crucially depends on the eccentricity of the ellipse. The ground state of the system displays a quantum phase transition from a two-peak configuration to a one-peak configuration at a critical attractive interaction strength. In correspondence of this phase transition the superfluid fraction strongly reduces and goes to zero for a sufficiently attractive BoseBose interaction.


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## 1 Introduction

How does a locally-varying spatial curvature influence the properties of low-dimensional quantum systems? This is a relevant question asked by scientists working in very different fields such as the
linear Schrödinger equation for a particle constrained on a curve manifold [1-3], but also quantum gravity [4] or quantum chaos [5]. It is well know [6-8] that the local curvature of a curve on the three-dimensional (3D) Euclidean space is characterized by the so-called geodesic curvature. This geodesic curvature $\kappa$ is an extrinsic quantity: it does not remain invariant if the curve is under the effect of a distance-preserving transformation [6-8]. Instead, the local curvature of a surface on the 3D Euclidean space is characterized by the so-called Riemann curvature tensor, which can be written in terms of the invariant Gaussian curvature and the not-invariant average curvature [6-8]. The quantum motion of a particle on a curved waveguide has been anayzed by several authors [9-14]. More recently, the highly nontrivial role of curvature for constrained quantum systems has been theoretically investigated with ultracold atomic gases confined in a quasi-1D [15-18] and quasi-2D configurations [19]. The main result of all these investigations is that the local curvature gives rise to a quantum-curvature potential.

In this paper we consider an atomic Bose-Einstein condensate (BEC) confined in a quasi-1D elliptical waveguide finding that the quantum-curvature potential has the shape of a double-well, if the eccentricity of the ellipse is different from zero. By numerically solving the 1D GrossPitaevskii equation of the BEC wavefunction under the effect of this quantum-curvature potential, we show that the ground state of the system is uniform along the waveguide only if the eccentricity $\epsilon$ of the ellipse is zero (circular waveguide with constant curvature). Instead, for $\epsilon \neq 0$ we find that the ground state is generically characterized by a two-peak configuration, where the peaks are located around the minima of the effective double-well potential. However, we discover that in the case of attractive interaction it exists a critical (negative) interaction strength below which the ground state exhibits a quantum phase transition from the two-peak configuration to a one-peak configuration. This is the analog of the spontaneous symmetry breaking, i.e. the modulational instability [20], of the uniform configuration predicted some years ago for an 1D attractive BEC in a circular waveguide $[21-23]$. Our results show that the critical interaction strength depends on the eccentricity $\epsilon$ of the ellipse in a non-trivial way. We also analyze the effect of a boost velocity on the BEC moving in the elliptical waveguide deriving the Leggett formula [24] for the superfluid fraction of a 1D bosonic system [25-27]. Our numerical investigation reveals that the superfluid fraction decreases dramatically in response to this quantum phase transition, eventually reaching zero for a sufficiently negative Bose-Bose interaction.

## 2 Quantum-curvature potential

We consider a Bose-Einstein condensate (BEC) made of $N$ identical bosonic atoms of mass $m$. The atoms are constrained to move along a curve $\mathcal{C}$ by the presence of a strong harmonic potential of frequency $\omega_{\perp}$ in the local transverse plane with respect to $\mathcal{C}$. The characteristic length of the transverse confinement is $l_{\perp}=\sqrt{\hbar /\left(m \omega_{\perp}\right)}$ where $\hbar$ is the reduced Planck constant. We introduce a local system $(s, u, v)$ of coordinates, where $s$ is the curvilinar abscissa (arclength) along $\mathcal{C}$ while $u$ and $v$ are two coordinates of the transverse plane [15-18]. In this way the Lagrangian density of our problem is given by

$$
\begin{equation*}
\mathscr{L}=\frac{i \hbar}{2}\left(\Psi^{*} \partial_{t} \Psi-\Psi \partial_{t} \Psi^{*}\right)-\frac{\hbar^{2}}{2 m}|\nabla \Psi|^{2}-\frac{m \omega_{\perp}^{2}}{2}\left(u^{2}+v^{2}\right)|\Psi|^{2}-\frac{1}{2} g|\Psi|^{4} \tag{1}
\end{equation*}
$$

where $\Psi(s, u, v, t)$ is the BEC wavefunction normalized to one and $g=4 \pi \hbar^{2} a_{s}(N-1) / m$ is the 3D strength of the contact inter-atomic potential with $a_{s}$ the s-wave scattering length. Clearly, the Laplacian operator $\nabla^{2}$ must be written in terms of the local system $(s, u, v)$ of coordinates [15-18]. Assuming the factorization

$$
\begin{equation*}
\Psi(s, u, v, t)=\psi(s, t) \frac{e^{-\frac{\left(u^{2}+v^{2}\right)}{2 \sigma(s, t)^{2}}}}{\pi^{1 / 2} \sigma(s, t)} \tag{2}
\end{equation*}
$$

and inserting this ansatz into the Lagrangian density, after integration over $u$ and $v$ one gets [18, 28, 29],

$$
\begin{align*}
\overline{\mathscr{L}} & =\frac{i \hbar}{2}\left(\psi^{*} \partial_{t} \psi-\psi \partial_{t} \psi^{*}\right)-\frac{\hbar^{2}}{2 m}\left|\partial_{s} \psi\right|^{2}+\frac{\hbar^{2} \kappa^{2}(s)}{8 m}|\psi|^{2} \\
& -\left(\frac{\hbar^{2}}{2 m} \frac{1}{\sigma^{2}}+\frac{m \omega_{\perp}^{2}}{2} \sigma^{2}\right)|\psi|^{2}-\frac{1}{2} \frac{g}{2 \pi \sigma^{2}}|\psi|^{4} \tag{3}
\end{align*}
$$

where $\kappa(s)$ is the local geodesic curvature of $\mathcal{C}$, and the conditions $\sigma \kappa \ll 1$ and $\sigma \ll \xi$ must hold, with $\xi=\hbar / \sqrt{2 g|\Psi|^{2}}$ the 3D healing length [18]. The Euler-Lagrange equations of the 1D action functional with respect to the 1 D wavefunction $\psi(s, t)$ and the transverse width $\sigma(s, t)$ are

$$
\begin{equation*}
i \hbar \partial_{t} \psi=\left[-\frac{\hbar^{2}}{2 m} \partial_{s}^{2}-\frac{\hbar^{2} \kappa^{2}(s)}{8 m}+\frac{\hbar^{2}}{2 m} \frac{1}{\sigma^{2}}+\frac{m \omega_{\perp}^{2}}{2} \sigma^{2}+\frac{2 \hbar^{2} a_{s}(N-1)}{m \sigma^{2}}|\psi|^{2}\right] \psi \tag{4}
\end{equation*}
$$

and

$$
\begin{equation*}
\sigma^{2}=l_{\perp}^{2} \sqrt{1+2 a_{s}(N-1)|\psi|^{2}} \tag{5}
\end{equation*}
$$

Eq. (4), equipped with Eq. (5), is the time-dependent 1D nonpolynomial Schrödinger equation (NPSE) [28,29] for the wavefunction $\psi(s, t)$ of the BEC moving along the curve $\mathcal{C}$ (see also [18]). As previously discussed, the geodesic curvature $\kappa(s)$ gives rise to an effective potential

$$
\begin{equation*}
U_{Q}(s)=-\frac{\hbar^{2} \kappa(s)^{2}}{8 m} \tag{6}
\end{equation*}
$$

This curvature potential $U_{Q}(s)$ is quantum because it involves the square of the reduced Planck constant $\hbar$. At fixed atomic mass $m$, only if the square of the curvature $\kappa(s)$ is sufficently large the effects of this quantum-curvature potential become relevant.

Under the assumption that $\sigma \simeq l_{\perp}$, which corresponds to a very strong transverse confinement, the 1D NPSE becomes the familiar 1D Gross-Pitaevskii (GPE) equation

$$
\begin{equation*}
i \hbar \partial_{t} \psi=\left[-\frac{\hbar^{2}}{2 m} \partial_{s}^{2}-\frac{\hbar^{2} \kappa(s)^{2}}{8 m}+\hbar \omega_{\perp}+\frac{2 \hbar^{2} a_{s}(N-1)}{m l_{\perp}^{2}}|\psi|^{2}\right] \psi \tag{7}
\end{equation*}
$$

It is very important to stress that, from Eq. (5], the condition $\sigma \simeq l_{\perp}$ implies $2 a_{s}(N-1)|\psi|^{2} \ll 1$. In the rest of the paper we will work within this 1 D regime. n the new version of the manuscript I shall discuss the role Clearly, Eq. (7) is reliable in the weak-coupling and strong-transverseconfinement regime, where both beyond-mean-field and transverse-size effects are very small.

## 3 Properties of the elliptical waveguide

We now choose an ellipse for the curve $\mathcal{C}$. By using cartesian coordinates its defining equation reads

$$
\begin{equation*}
\frac{x^{2}}{a^{2}}+\frac{y^{2}}{b^{2}}=1 \tag{8}
\end{equation*}
$$

where $a$ and $b$ are the lengths of the two semi-axes of the ellipse. Here we assume that $a \geq b$, such that $a$ is the length of the major semi-axis. The eccentricity of the ellipse is defined as [8]

$$
\begin{equation*}
\epsilon=\sqrt{1-\frac{b^{2}}{a^{2}}} \tag{9}
\end{equation*}
$$

Clearly, $0 \leq \epsilon<1$ and for $\epsilon=0$ we obtain a circle of radius $R=a=b$. Introducing the angle $\phi \in[0,2 \pi]$ we can write

$$
\begin{array}{r}
x=a \cos (\phi) \\
y=b \sin (\phi) \tag{11}
\end{array}
$$



Figure 1: Quantum-curvature potential $U_{Q}(s)$, Eq. [6], induced by the geodesic curvature $\kappa(s)$ of an ellipse, as a function of the arclength $s$, where $a$ is the length of the major semi-axis, $L=$ $a E(2 \pi, \epsilon)$ is the perimeter of the ellipse, and $E_{a}=\hbar^{2} /\left(m a^{2}\right)$ a characteristic energy. The curves are obtained for different values of the eccentricity $\epsilon$.
and the arclength $s$ along the ellipse can be expressed with the formula [8]

$$
\begin{equation*}
s=a E(\phi, \epsilon) \tag{12}
\end{equation*}
$$

where

$$
\begin{equation*}
E(\phi, \epsilon)=\int_{0}^{\phi} \sqrt{1-\epsilon^{2} \sin ^{2}\left(\phi^{\prime}\right)} d \phi^{\prime} \tag{13}
\end{equation*}
$$

is the incomplete elliptic integral of the second kind. It follows that the perimeter $L$ of the ellipse reads

$$
\begin{equation*}
L=a E(2 \pi, \epsilon), \tag{14}
\end{equation*}
$$

such that for $\epsilon=0$ we have $L=2 \pi a$ because $E(2 \pi, 0)=2 \pi$. Instead, for $\epsilon \rightarrow 1$ we have $L \rightarrow 4 a$ because $E(2 \pi, 1)=4$. We conclude that $4 a<L \leq 2 \pi a$. The geodesic curvature $\kappa$ of the ellipse can be written as [8]

$$
\begin{equation*}
\kappa=\frac{1}{a} \frac{\sqrt{1-\epsilon^{2}}}{\left(\sin ^{2}(\phi)+\sqrt{1-\epsilon^{2}} \cos ^{2}(\phi)\right)^{3 / 2}} \tag{15}
\end{equation*}
$$

At fixed $a$ and $\epsilon$, the maximum of the curvature is obtained for $\phi=0$ and $\phi=\pi$, i.e.

$$
\begin{equation*}
\kappa_{\max }=\frac{1}{a} \frac{1}{\left(1-\epsilon^{2}\right)^{1 / 4}} \tag{16}
\end{equation*}
$$

while the minimum of the curvature is obtained for $\phi=\pi / 2$ and $\phi=3 \pi / 2$, i.e.

$$
\begin{equation*}
\kappa_{\min }=\frac{1}{a} \sqrt{1-\epsilon^{2}} \tag{17}
\end{equation*}
$$

Thus, for $\epsilon=0$ we have $\kappa_{\max }=\kappa_{\min }=1 / a$ while for $\epsilon \rightarrow 1$ we have $\kappa_{\max } \rightarrow+\infty$ and $\kappa_{\text {min }} \rightarrow 0$. We conclude that $1 / a \leq k_{\max }<+\infty$ and $0<k_{\min } \leq 1 / a$. In general, the


Figure 2: Probablility density $\rho(s)$ of the non-interacting BEC ground state in an ellipse as a function of the arclength $s$, where $a$ is the length of the major semi-axis and $L=a E(2 \pi, \epsilon)$ is the perimeter of the ellipse. The curves are obtained for different values of the eccentricity $\epsilon$. Here the s-wave scattering length $a_{s}$ is set to zero or, equivalently, the number $N$ of particles is set to one.
formula which gives the curvature $\kappa$ as a function of the arclength $s$ is called Cesaro equation. Unfortunately, in the case of the ellipse there is no Cesaro equation. In other words, an analytical formula of $\kappa$ as a function of $s$ is not available. However, from Eqs. (12) and (15), fixing the length $a$ and the eccentricity $\epsilon$ of the ellipse, we can easily plot $\kappa$ vs $s$ using $\phi$ as dummy variable. More explicitly: we calculate separately $\kappa$ vs $\phi$ and $s$ vs $\phi$, and then we plot $\kappa$ vs $s$. The curvature $\kappa(s)$ has a the periodic structure of $\kappa(s)$. By increasing $\epsilon$, the perimeter $L$ of the ellipse slightly decreases while $\kappa_{\max }$ and $\kappa_{\min }$ pull away. In Fig. 1 we plot the quantum-curvature potential $U_{Q}(s)$, Eq. [6), induced by the curvature $\kappa(s)$ of the ellipse for different values of the eccentricity $\epsilon$. The figure clearly shows that, for $\epsilon \neq 0, U_{Q}(s)$ is symmetric double-well potential where the depth of the wells becomes larger by increasing the eccentricity. The minima (maxima) of the quantum-curvature potential $U_{Q}(s)$ are in correspondence to the maxima (minima) of the curvature $\kappa(s)$.

## 4 BEC ground state in elliptical waveguide

The time-independent 1D GPE is obtained from Eqs. (7) setting

$$
\begin{equation*}
\psi(s, t)=\Phi(s) e^{-i\left(\mu+\hbar \omega_{\perp}\right) t / \hbar} \tag{18}
\end{equation*}
$$

In this way we have

$$
\begin{equation*}
\mu \Phi=\left[-\frac{\hbar^{2}}{2 m} \partial_{s}^{2}-\frac{\hbar^{2} \kappa(s)^{2}}{8 m}+\frac{2 a_{s}(N-1) \hbar^{2}}{m l_{\perp}^{2}}|\Phi|^{2}\right] \Phi \tag{19}
\end{equation*}
$$

that is the 1D GPE equation for the stationary wavefunction $\Phi(s)$, such that $\rho(s)=|\Phi(s)|^{2}$ is the probability density of finding the BEC at the position $s$. In Fig. 2 we report the probability


Figure 3: Probablility density $\rho(s)$ of the attractive BEC ground state in an ellipse as a function of the arclength $s$, where $a$ is the length of the major semi-axis and $L=a E(2 \pi, \epsilon)$ is the perimeter of the ellipse. The curves are obtained with eccentricity $\epsilon=0.9$ for different values of the adimensional interaction strength $\gamma=2 a a_{s}(N-1) / l_{\perp}^{2}$, where $a_{s}$ is the 3D s-wave scattering length and $l_{\perp}$ is the characteristic length of the transverse harmonic confinement.
density $\rho(s)$ of the ground-state of the non-interacting ( $a_{s}=0$ or, equivalently, $N=1$ ) BEC confined along the ellipse as a function of the arclength $s$. Our results are obtained by solving Eq. (7) with a Crank-Nicolson predictor-corrector method and imaginary time. In the figure, the curves correspond to different values of the eccentricity $\epsilon$. Clearly, for $\epsilon=0$ the ground state is uniform along the ellipse. However, for $\epsilon \neq 0$ the ground state is no more uniform due to a non-constant curvature $\kappa(s)$ which implies a non-constant effective potential $U_{Q}(s)=$ $-\hbar^{2} \kappa(s)^{2} /(8 m)$. By increasing the eccentricity $\epsilon$ the localization of $\rho(s)$ around the minima of $U_{Q}(s)$, where the curvature is larger, becomes more evident.

## 5 Quantum phase transition

It is interesting to investigate the effect of the inter-atomic interaction on the ground state properties of the system. In adimensional units the interaction strength reads $\gamma=g /\left(2 \pi l_{\perp}^{2} a E_{a}\right)=2 a a_{s}(N-$ 1) $/ l_{\perp}^{2}$ with $E_{a}=\hbar^{2} /\left(m a^{2}\right)$. In Fig. 3 we consider the case of an attractive BEC and plot our numerical results obtained for different values of a negative $\gamma$ with fixed eccentricity $\epsilon=0.9$. Quite remarkably, for $\gamma<-1.5$ the ground state has a spontaneous symmetry breaking: one of the two local mimima contains more bosons. Indeed for $\gamma=-2$ this single-well localization is very clear. It is important to observe that this kind of quantum phase transition happens for any $\epsilon$. In the case of a circle $(\epsilon=0)$ a similar spontaneous symmetry breaking was predicted about 20 years ago [21-23]. Actually, this quantum phase transition, or spontaneous symmetry breaking, is nothing else than the modulational instability of the ground-state configuration, induced by the appearance of an imaginary component in the energies of the elementary excitations of the ground state [20]. However, for $\epsilon=0$ there is a quantum phase transition from a uniform configuration to a
single-peak configuration, while for the $\epsilon \neq 0$ there is a quantum phase transition from a two-peak configuration to a single-peak configuration. This quantum phase transition was observed some years ago with an attractive BEC of ${ }^{39} \mathrm{~K}$ atoms, where the double-well potential was created by intersecting two pairs of laser beams [31]. Our double-well system is slightly different because the particles tunnel from one well to the other well following two different curved paths; moreover, our elipsoidal configuration offers also the possibility of having persistent currents. For a sufficiently attractive BEC the transverse width $\sigma$ of the BEC becomes smaller than $l_{\perp}$ and there is the collapse of the single-peak configuration. For $\epsilon=0$ the 1D NPSE predicts the collapse of this single-peak configuration at $\gamma_{c}=(4 / 3)\left(a / l_{\perp}\right)$ [28,30]. Thus, our numerical results of Fig. 3, obtained from the 1D GPE, are fully reliable under the condition $a / l_{\perp} \gg 1$. This is again the condition of a tight transverse confinement.

## 6 Superfluid fraction

Let us consider the effect of a boost velocity $v_{B}$ on the BEC moving along the ellipse. In this case Eq. 19 is modified as follows

$$
\begin{equation*}
\mu \Phi=\left[\frac{1}{2 m}\left(-i \hbar \partial_{s}-m v_{\mathrm{B}}\right)^{2}-\frac{\hbar^{2} \kappa^{2}(s)}{8 m}+\frac{2 a_{s}(N-1) \hbar^{2}}{m l_{\perp}^{2}}|\Phi|^{2}\right] \Phi \tag{20}
\end{equation*}
$$

Now we set

$$
\begin{equation*}
\Phi(s)=\frac{n(s)^{1 / 2}}{\sqrt{N}} e^{i \theta(s)} \tag{21}
\end{equation*}
$$

where $n(s)=N \rho(s)$ is the local number density of the BEC. Morover, we introduce the local velocity field

$$
\begin{equation*}
v(s)=\frac{\hbar}{m} \partial_{s} \theta(s) \tag{22}
\end{equation*}
$$

Inserting these formulas into Eq. (20) we obtain 1D stationary equations of zero-temperature superfluid hydrodynamics

$$
\begin{equation*}
\mu=-\frac{\hbar^{2}}{2 m \sqrt{n}} \partial_{s}^{2} \sqrt{n}+\frac{m}{2}\left(v-v_{\mathrm{B}}\right)^{2}-\frac{\hbar^{2} \kappa(s)^{2}}{8 m}+\frac{2 a_{s}(1-1 / N) \hbar^{2} n}{m l_{\perp}^{2}} \tag{23}
\end{equation*}
$$

and also

$$
\begin{equation*}
\partial_{s}\left[n\left(v-v_{B}\right)\right]=0 \tag{24}
\end{equation*}
$$

Eq. (24) implies that $n\left(v-v_{B}\right)=J$, where $J$ is a constant current density. This result is very interesting because it says that if $n(s)$ has spatial variations then also $v(x)$ must have spatial variations.

Inspired by Ref. [27], we now introduce the average value of the velocity $v(x)$ in a region $[a, b]$ of the ellipse as

$$
\begin{equation*}
\bar{v}=\frac{1}{(b-a)} \int_{a}^{b} v(s) d s \tag{25}
\end{equation*}
$$

Then, from the previous equations we obtain

$$
\begin{equation*}
\bar{v}=\frac{1}{(b-a)} \int_{a}^{b}\left(\frac{J}{n(s)}+v_{B}\right) d s=\frac{J}{\bar{n}_{s}}+v_{B} \tag{26}
\end{equation*}
$$

where

$$
\begin{equation*}
\bar{n}_{s}=\frac{1}{\frac{1}{(b-a)} \int_{a}^{b} \frac{1}{n(s)} d s} \tag{27}
\end{equation*}
$$

The number density $\bar{n}_{s}$ can be interpreted as the superfluid number density of the stationary state in the spatial region $[a, b]$. Indeed, Eq. (27] is the 1D version of the formula obtained by Leggett [24] for a supersolid with spatial periodicity $(b-a)$, and recently discussed by others [25-27]. If the stationary state $\Psi(s)$ moves with the average velocity $\bar{v}$, its current density reads $J=\bar{n}_{s}\left(\bar{v}-v_{B}\right)$, where $\bar{v}$ is the average velocity in the region $[a, b]$ and $\bar{n}_{s}$ the corresponding superfluid number density. We can also introduce

$$
\begin{equation*}
\bar{n}=\frac{1}{(b-a)} \int_{a}^{b} n(s) d s \tag{28}
\end{equation*}
$$

that is the average number density in the region $[a, b]$. Consequently, the superfluid fraction of the BEC in the region $[0, L]$ reads

$$
\begin{equation*}
f_{s}=\frac{\bar{n}_{s}}{\bar{n}}=\frac{1}{\frac{N}{L^{2}} \int_{0}^{L} \frac{1}{n(s)} d s} \tag{29}
\end{equation*}
$$

where $N=\int_{0}^{L} n(s) d x=L \bar{n}$. This formula can be also obtained as the response of the linear momentum of the BEC to the boost velocity $\bar{v}_{B}$, that is the non-classical translational inertia of the system [25].

In Fig. 4 we plot our numerical results of the superfluid fraction $f_{s}$ as a function of the adimensional interaction strength $\gamma$ for different values of the eccentricity $\epsilon$ of the ellipse. For positive values of $\gamma$ the superfluid fraction $f_{s}$ is close to 1 also with $\epsilon=0.9$. However, for $0 \leq$ $\gamma<1$ and a very large eccentricity $(\epsilon=0.99)$ we find $f_{s} \simeq 0.95$. This result is quite reasonable because the wavefunction is strongly localized in the two well of the quantum-curvature potential. For negative values of $\gamma$ the most interesting effect appears in Fig. 4, around $\gamma \simeq-1.6$ the superfluid fraction $f_{s}$ quckly decreases and it goes to zero for very large negative values of $\gamma$. This is exacly the quantum phase transition from a two-peak configuration to a one-peak configuration. As previously stressed, the one-peak configuration becomes modulationally unstable when at least one of the energies of its elementary excitations acquires an imaginary component. A similar modulational instability [20] happens in the formation of a train of bright solitons from a singlepeak Bose-Einstein condensate, induced by a sudden change in the sign of the scattering length from positive to negative [32]. Our Fig. 4reveals that the critical strength $\gamma_{c}$ crucially depends on the eccentricity $\epsilon$, such as the behavior of $f_{s}$ as a function of $\gamma$ for $\gamma<\gamma_{c}$. In particular, we find that $\gamma_{c}$ slightly reduces by increasing $\epsilon$, but this effect is quite weak.

An important remark is that Eq. (29) has been derived here without any assumption about the sign of $\gamma$. Moreover, the absence of superfluidity for $\gamma=0$ is true only in the thermodynamic limit. In a ring there is a finite energy gap between the ground state and the first excited state also for $\gamma=0$. The Bose system we are considering has a finite size because it is confined in a finite elliptical ring. Finally, the 3D version of Eq. (27) was proposed historically by Leggett to characterize the superfluid density of a supersolid [24], while here Eq. (27) is used to determine the superfluid density of a Bose-Einstein condensate which is not supersolid but it is instead spatially modulated due to the crucial interplay between elliptical confinement and attractive interaction.

## 7 Conclusions

The main goal of this paper was to understand the role of a locally-varying curvature for a BoseEinstein condensate confined in an elliptical waveguide. The proposed setup, and the double-well quantum-curvature potential that we have found, can be experimentally achieved by using ultracold atoms, which are a paradigmatic physical platform due to the high experimental tunability of inter-atomic interactions and trapping potentials. For instance, one can trap $N=10^{4}$ ultracold Rb atoms by using a rapidly moving laser beam which creates a time-averaged elliptic-shaped toroidal


Figure 4: Superfluid fraction $f_{s}$ of the BEC ground state in an ellipse as a function of the adimensional interaction strength $\gamma=2 a a_{s}(N-1) / l_{\perp}^{2}$, where $a$ is the length of the major semi-axis, $a_{s}$ is the 3 D s-wave scattering length and $l_{\perp}$ is the characteristic length of the transverse harmonic confinement. The curves are obtained for three values of the eccentricity $\epsilon$ of the ellipse.
optical dipole potential [33]. The length $a$ of the major semi-axis of the ellpse can be $a \simeq 100$ microns and the transverse length $l_{\perp}=5$ microns. The scattering length $a_{s}$ could be then tuned by using an external magnetic field, which induces a Fano-Feshbach resonance [34]. Despite the fact that we focused on space curvature rather than space-time curvature, we believe that our results can be of interest not only to atomic and condensed matter physics researchers, but also to a large community working on general relativity and relativistic quantum field theory.

Acknowledgements. The author thanks Francesco Ancilotto, Koichiro Furutani, Francesco Minardi, and Andrea Tononi for useful discussions.

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