

# DEASA STUDIES AND APPLICATIONS TO SPACE PHYSICS AND MUON TOMOGRAPHY

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## 3 Abstract

4 The high energy cosmic rays entering the Earth's atmosphere throw light upon many dif-  
5 ferent aspects of Astroparticle Physics and Particle Physics. The work presented in this  
6 symposium outlines investigative learning about these high energy primaries based on a  
7 mini array DEASA in Agra, India. DEASA (Dayalbagh Educational Air Shower Array) con-  
8 sists of eight plastic scintillators each with an area of 1 square meter. This array covers  
9 an area of 260 square meters and is the first array in the northern part of our country.  
10 The real-life applications of the cosmic ray particles where the effect of cosmic rays in  
11 space has been studied and the best material has been found to protect the astronaut  
12 from the galactic cosmic rays. Poly materials were found to be the best material due to  
13 large amount of hydrogen (H) and low atomic number (Z). It is observed that equiva-  
14 lent dose is minimum (107 sieverts) for Polystyrene compared to the other materials.  
15 Finally the high energy muons have been studied to image nuclear caskets called muon  
16 tomography. In this study, a dry cask container has been simulated which contains the  
17 UO2 rods and the muon scattering has been observed.

## 18 1 Introduction

19 The cosmic rays come from radioactive decay inside the stars, explosive supernovae's, Sun  
20 and pulsars etc. The higher energy ones seem to be coming from super massive black holes at  
21 the heart of galaxies. On reaching the Earth's atmosphere, they produce showers of particles  
22 which pass through us almost 500 times in a minute. The cosmic flux is an important tool  
23 for calibrating particle detectors and this study is being done for DEASA detectors also. The  
24 muons entering our detectors do not have a constant flux but slightly more in summer and  
25 lower in winter. This is connected to pions which have decayed into muons in the shower.  
26 In summer the air warms and expands, leading to more gap between air molecules allowing  
27 pions to reach further so as to decay into muons. In winter the air is cold and dense resulting  
28 into higher collisions of pions leading to fewer decays into muons.

29 The cosmic rays before entering the atmosphere are mostly galactic energetic particles coming  
30 from inside the galaxy, extragalactic with energetic particles from the active galactic nuclei,  
31 quasars or gamma ray bursts. These energetic particles affect the human body in many dif-  
32 ferent ways as published study [1] says that the twins physiology, memory abilities and genes

33 for one of the twin on Earth and other on ISS for 340 days. The study confirms that space time  
34 manipulates the genes, affects the human immune system. The exposed person suffers from  
35 loss in mental reasoning and memory and studies are going on for long term ailments. One  
36 of the stickiest problems for NASA is how to shield astronaut from energetic cosmic rays and  
37 solar flares.

38 The air shower developed by an energetic particle entering the atmosphere grows with depth  
39 into hadronic and electromagnetic particles at the sea level. These muons can look into the  
40 interiors of impenetrable structures in parallel to the x-ray imaging of our body. The difference  
41 being that X-rays have to be produced in the laboratory and muons are always available. This  
42 feature defines them as a good tool for imaging the impenetrable structures like pyramids,  
43 volcano's to nuclear reactor containers .Muons traveling through a structure will be stopped  
44 along the path or scattered depending on the thickness and density of the material. The plastic  
45 scintillator lights up when a charged particle passes through so we design the simulation with  
46 a nuclear casket surrounding by two plastic scintillators. This muon imaging technique has  
47 been used to image the interiors of the nuclear reactors at Fukushima Daiichi plant[2] also.

## 48 2 DEASA

49 The cosmic ray flux decreases rapidly with energy as  $E^{-2.7}$  around  $10^{14}$  eV, hence impossible to  
50 have direct measurements. At these energies the secondaries produced at sea level increases  
51 with primary particle energy. The change in transverse momentum and scattering of the sec-  
52 ondary particles with the atmospheric particles leads to their lateral spread on ground. This  
53 process of almost parallel arrival of the secondary particles reaching ground is called extensive  
54 air shower in which the spread is between  $10^4 m^2$  to  $10 km^2$ .

55 This phenomenon gave insight into:

- 56 1. Particle Physics from air shower spread.
- 57 2. The direction of secondaries arriving on ground tells about high energy particles.
- 58 3. The cosmic ray energy spectrum.
- 59 4. Mass of primary cosmic rays.

60 DEASA is a mini array of eight plastic scintillators each with an area of 1 square meter, has  
61 been set up as shown in Figure 1. This array covers an area of 260 square meters and is the  
62 first array in the northern part of our country. The pulses from the eight detectors are being  
63 manually studied. The pulse amplitudes, Time over Threshold, rise time, fall time and Full width  
64 half maximum are being observed to study correlations between them. The calibration of the  
65 12 dynode Photomultiplier tubes attached to each of the eight detectors has been completed  
66 and the flux measured is around 13500 counts per minute. Daily monitoring of the detectors  
67 is maintained in log book.

## 68 3 SPACE STUDIES

69 The real-life applications of the cosmic ray particles is the effect of cosmic rays in space to  
70 find out the best material to protect the astronaut from the solar energetic particles(SEP) and  
71 galactic cosmic rays(GCR) [3]. The water phantom was irradiated with primary proton follow-  
72 ing a galactic cosmic ray energy spectrum with different shielding materials. The secondary  
73 particles are created with interactions between protons and the shield material in Geant4.



Figure 1: The mini array

74 Poly materials were found to be the best material due to large amount of hydrogen (H) and  
75 low atomic number (Z). High H leads to fragmentation of the heavy GCR particles into small  
76 fragments and low Z produces maximum number of secondaries. Poly materials are 16 per-  
77 cent more effective than aluminum in reducing the dose equivalent with only 1.5148 g/cm<sup>2</sup>  
78 of material. Polymeric materials are expected to play an important role in protecting the as-  
79 tronauts on future missions. It is observed that equivalent dose is minimum (107 sievert) for  
80 Polystyrene as compared to the other materials.

## 81 4 MUON TOMOGRAPHY

82 The second study defines the application of muons to identify nuclear wastage using plastic  
83 scintillation detectors [4] in muon tomography. In this study, a dry cask container has been  
84 simulated which contains the UO<sub>2</sub> rods (varying in number) and the muon scattering has been  
85 observed [4]. This shows that when the dry cask is filled with the rods, muons scattered to  
86 the maximum angle and if the dry cask is empty, the muon will pass through it straight with-  
87 out getting scattered. The scattering of energetic muons of range 3 GeV – 10 GeV from these  
88 containers with dimensions from Narora Nuclear plant, Uttar Pradesh(U.P). The parameters  
89 measured are energy loss, radiation length and scattering angle for different number of rods  
90 gives us patterns which describe the inside of the containers without opening them. The ra-  
91 diation length is the average distance required for an electron to lose 1/e of its energy and  
92 measured in cm. The multiple scattering of the muons is mostly due to Coulomb scattering of  
93 muons in the target with charge Z, calculated analytically.

94 Muon with energy 3 GeV loses 3.64 MeV/c energy in concrete and the scattering angle  
95 was found to be 4.14 mrad whereas the radiation length was 10.91 cm. These calculations  
96 have been done for different energy muons scattered from Iron, Lead and Uranium targets in  
97 this paper.

## 98 5 Conclusion

99 These studies prove that high energy quantum fields consciously assist us in applications  
100 beyond the accelerating sources from which they arrive and reach far beyond the human -  
101 machine interface. Neutrinos although being nearly massless give solutions to Dark matter,  
102 Dark energy in cosmology, muons being tiny particles can scan structures like nuclear plants,  
103 submarines etc. Hadron fields have applications in medical physics such as hadron therapy

104 and carbon ion therapy for the cancer patients.  
105 The importance of cosmic ray studies at DEASA is that students can understand quantum  
106 sensors, the electronics for fast pulses and count rates analysis over the different seasons.  
107 Another important aspect of these studies is the Monte Carlo simulations in Geant4 and COR-  
108 SIKA. These codes give a wide angle view to the user and applications in space physics, nuclear  
109 science and many other areas.

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