

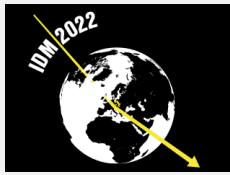
Sub-GeV Dark Matter Searches with EDELWEISS: New results and prospects

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October 4, 2022

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2 *14th International Conference on Identification of Dark Matter*
Vienna, Austria, 18-22 July 2022
doi:[10.21468/SciPostPhysProc.7](https://doi.org/10.21468/SciPostPhysProc.7)

3 Abstract

4 The Edelweiss collaboration performs light Dark Matter (DM) particles searches with
5 germanium bolometer collecting charge and phonon signals. Thanks to the Neganov-
6 Trofimov-Luke (NTL) effect, a RMS resolution of 4.46 electron-hole pairs was obtained
7 on a massive (200g) germanium detector instrumented with a NbSi Transition Edge Sen-
8 sor (TES) operated underground at the Laboratoire Souterrain de Modane (LSM). This
9 sensitivity made possible a search for WIMP using the Migdal effect down to 32 MeV/C²
10 and exclude cross-sections down to 10⁻²⁹ cm². It is the first measurement in cryogenic
11 germanium with such thermal sensor, proving the high relevance of this technology. Fur-
12 thermore, such TES have shown sensitivity to out-of-equilibrium phonons, paving the
13 way for EDELWEISS new experience CRYOSEL. This is an important step in the devel-
14 opment of Ge detectors with improved performance in the context of the EDELWEISS-
15 SubGeV program.

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1 Introduction

There were great progresses in the Dark Matter (DM) direct interaction with nuclei searches [1–3], for particle with masses ranging from $1 \text{ GeV}\cdot\text{c}^{-2}$ to $1 \text{ TeV}\cdot\text{c}^{-2}$ [4–6]. Nevertheless, because of the absence of signal in that region, the interest in extending the search to lower masses intensified [7–13]. Probing lower masses raise additional experimental constraints, such as the need of lower energy threshold (below 1 keV) as well as ionization or scintillation yield for nuclear recoil. To such constraints, collaborations have to add dealing with new types of low energy backgrounds [14–19].

In order to avoid the issues raised by the very low kinetic energy of the nuclear recoil, it is possible to use the Migdal effect [24–27]. This effect quantifies the probability that an atomic electron is released because of the DM particle scattering on a nucleus. The electrons are emitted with an energy typically much larger than the kinematic energy of the nuclear recoil [27], yielding signals much easier to detect.

In this context, the EDELWEISS collaboration used the Migdal effect to extend the mass range of DM particles search down to $32 \text{ MeV}\cdot\text{c}^{-2}$, using a 200 g cryogenic Ge detector equipped with a NbSi Transition Edge Sensor (TES) and operated underground at the Laboratoire Souterrain de Modane (LSM). This detector achieved a low energy threshold for electron recoil and appeared to be less sensitive to the Heat Only background (HO), an already known background made of events that are not associated with charges.

The development of the NbSi detector is part of the EDELWEISS-SubGeV program and a step toward the new generation of EDELWEISS detectors, Cryosel. Those detectors will be Ge bolometers able to sustain high biases and equipped with a new sensor able to tag down to a single charge and thus veto HO events [28].

2 New results : DM search with a detector equipped with TES

The search was performed at the LSM benefitting from the ultra-low background environment of the EDELWEISS-III cryostat [29] and the 4800 m.w.e. rock overburden. The experiment, its results and their interpretation are described in detail in [30]. The detector named NbSi209 is a 200 g Ge cylindrical crystal (48 mm in diameter and 20 mm in height) on top of which was lithographed a $\text{Nb}_x\text{Si}_{1-x}$ thin film TES [31] used as thermal sensor. The TES is operated around 44 mK, the temperature at which the transition between the normal and the superconducting states start to occur. The data from the ionization and heat channels were digitized at a frequency of 100 kHz, filtered, and continuously stored on disk with a digitization rate of 500 Hz. The bottom side is fully covered by an electrode biased at a voltage varying between $\pm 66 \text{ V}$. The bias is applied in order to trigger the so-called Neganov-Trofimov-Luke (NTL) effect [32, 33]. This effect states that the drift of the N electron-hole pairs across a voltage difference ΔV creates additional phonons of energy $E_{NTL} = Ne\Delta V$ (e is the elementary charge). This energy E_{NTL} adds to the recoil energy. In the case of events associated with charges, the mean gain of the NTL effect is $\langle g \rangle = 1 + e\Delta V / \epsilon_{\text{gamma}}$, where $\epsilon_{\text{gamma}} = 3.0 \text{ eV}$ is the average ionization energy in Ge for electron recoils [34]. The average resolution over the dataset is between 90 and 100 eV. For the NTL gain of $\langle g \rangle = 23$ corresponding to a bias of 66 V, this leads to a resolution of approximately 4 eV electron-equivalent (eV_{ee}).

The detector was activated with a strong AmBe neutron source. This produces short-lived ^{71}Ge isotopes which decay by electron capture in the K, L, and M shells, with de-excitation x-ray lines at 10.37, 1.30, and 0.16 keV, respectively. The isotopes being produced uniformly throughout the detector volume, the absorption of the low energy x-ray lines allows to probe the DM response in the entire detector volume, providing a precise characterization of low

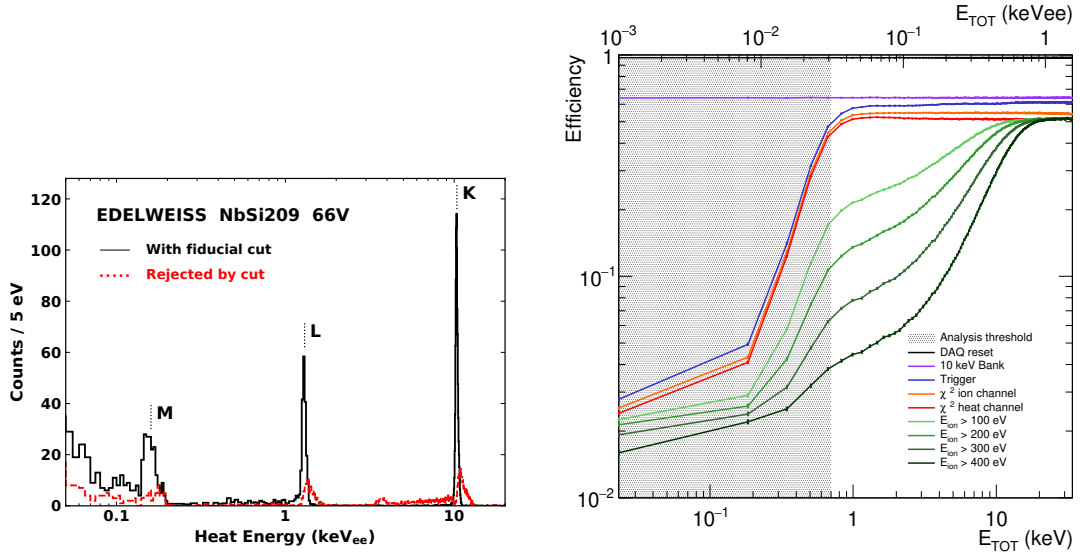


Figure 1: Left: energy spectrum recorded at a bias of 66 V following the ^{71}Ge emitted x-ray lines induced by the neutron activation of the Ge detector, resulting in the strong lines characteristic of ^{71}Ge x-ray emission. Right : efficiency as a function of the input phonon heat energy of simulated events inserted in the data stream in units of keV (lower axis), or keV-electron-equivalent (keV_{ee}, upper axis). Each color line corresponds to a step in the trigger and data analysis process. Those figures have been taken from [30] under RNP/22/OCT/058564 license.

71 energy calibration and data selection efficiency. Those lines are visible in Fig. 1, left panel,
 72 which shows the energy spectrum for the phonon signal from calibration data recorded at
 73 a bias of 66 V. A sample of 10667 K-shell events with energies between 2 and 12.6 keV_{ee},
 74 recorded at 66 V was used to assess the efficiency of data processing, triggering and selection
 75 as function of the energy. In order to probe the energy dependency, those events were scaled
 76 to the desired energy and injected at random time in the data stream. Those events went
 77 through the same analysis pipeline as the original data. The Figure 1, right panel shows the
 78 efficiency as a function of the energy of the injected pulse at various steps of the triggering,
 79 processing and selection of the data analysis. This procedure of event simulation accounts for
 80 actual noise conditions in data, as well as reconstruction and selection biases. The efficiency
 81 plateau is around $\sim 50\%$ for the bulk of the selection, in top of which a stringent criterion
 82 on the ionization energy is added in order to reject most of the HO background, the effect on
 83 efficiency is shown by the green curves.

84 The data have been interpreted in terms of limits on the spin-independent interaction of
 85 DM particles with target atoms through the so-called Migdal effect using calculations from
 86 [27, 35]. In addition, because of the underground location of the detector, the action of the
 87 stopping power of the rock overburden on the DM particle flux has been taken into account
 88 [36–39]. The search was performed with data recorded between March 2019 and June 2020,
 89 only considering samples with a baseline heat energy resolution lower than 140 eV RMS. The
 90 resulting average resolution is 102 ± 12 eV RMS, corresponding to 4.46 ± 0.54 eV_{ee} RMS
 91 once the NTL gain $\langle g \rangle$ is considered. Half of the dataset was blinded to perform the search,
 92 the other half was used to set the analysis criteria. The number of events in the chosen energy
 93 range (the regions of interest ROI's), is used to set the 90% C.L. Poisson upper limit on DM
 94 particle interacting through Migdal effect with Ge nuclei. The signal energy distributions for
 95 a WIMP of 50 MeV, corrected for efficiency and smeared to detector resolution for upper and
 96 lower excluded cross-sections are shown in Fig 2 left panel in green and red, respectively, as

97 well as their corresponding ROI's. The non-blinded sample was used to derive a data model
 98 in order to set the ROI's prior to unblind the rest of the data. This model corresponds to the
 blue curve in Fig 2 left panel.

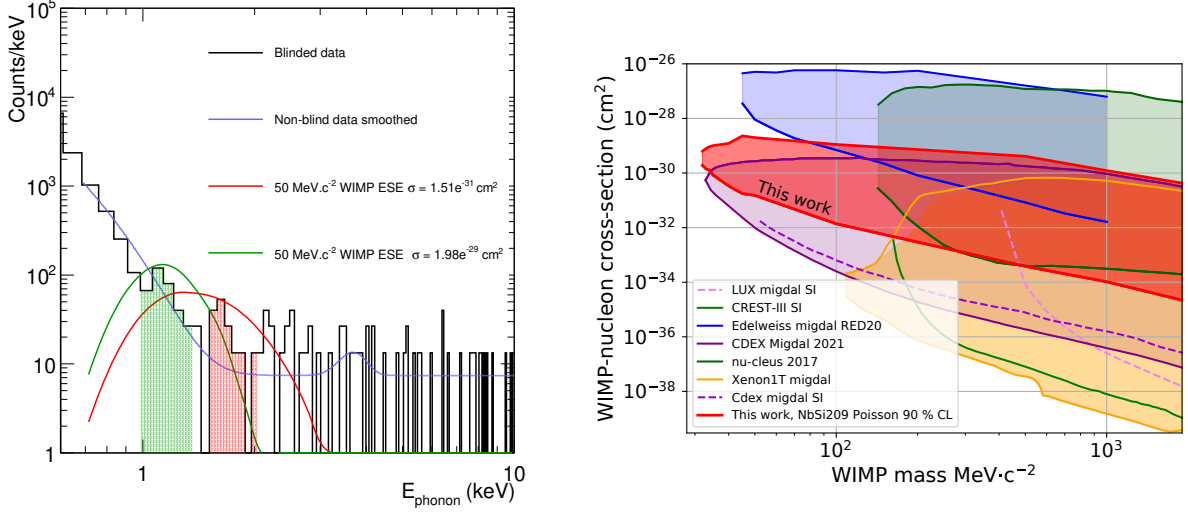


Figure 2: Left: energy spectrum after all criteria are applied for the blinded dataset in counts per keV (black histogram). Reference sample data smoothed by analytical function (plain blue). The red and green curves show the excluded Migdal spectra smeared to detector resolution, corrected for the Earth-shielding effect and efficiency corrected for WIMPs of $50 \text{ MeV}\cdot\text{c}^{-2}$ for lower and upper excluded cross-section. Right : 90% C.L. upper limit on the cross-section for Spin-Independent interaction between DM and Ge nuclei through the Migdal effect. The red contour is obtained by taking into account the slowing of the DM particle flux through the material above the detector. These results are compared to other experiments [19, 20, 23, 40–42]. Those figures have been taken from [30] under RNP/22/OCT/058564 license.

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100 The resulting limits are shown by the red contour (delimited by the thick red line) in
 101 Fig 2, right panel. Those DM constraints go down to $32 \text{ MeV}\cdot\text{c}^{-2}$. Below that mass, the large
 102 cross-section that would yield an observable signal is prevented by the stopping effect of the
 103 overburden and shielding. This contour is compared with results from other experiments [19,
 104 20,23,40–42] and with the previous EDELWEISS Migdal search performed above ground. This
 105 shows orders of magnitude of improvement with this previous search and that the underground
 106 operation of the detector did not jeopardize the potential of a Migdal search.

107 Nevertheless, despite the efforts made to reject the HO contamination, this analysis was
 108 background limited. Although its HO nature is not questionable at high energy because of the
 109 ionization signal associated with the heat signal, the ionization resolution of $\sim 200 \text{ eV}$ (RMS)
 110 prevent such conclusion at lower energy. Fig 3 shows energy distribution of data recorded at
 111 15 V and 66 V in red and black, respectively. The lack of quantitative shape differences of
 112 those spectrums is a hint of their HO nature. The fit of a power law (αE^β) yields identical
 113 slopes within uncertainties $\beta \sim 3.40$ for both spectra. The compatible fitted powers indicate
 114 that no NTL amplification occurs, and therefore, that no charges are involved. The flatness
 115 of the ratio shown on the bottom pad enhanced this observation. Furthermore, assuming the
 116 that both the HO background and the possible electronic background follow the same power
 117 law, a limit on the fraction of events associated with charges x can be set, the resulting upper
 118 limit is $x < 0.0004$ at 90%C.L. This confirms that the nature of this background is mainly HO
 119 in the energy range of the search between 0.8 to 2.8 keV.

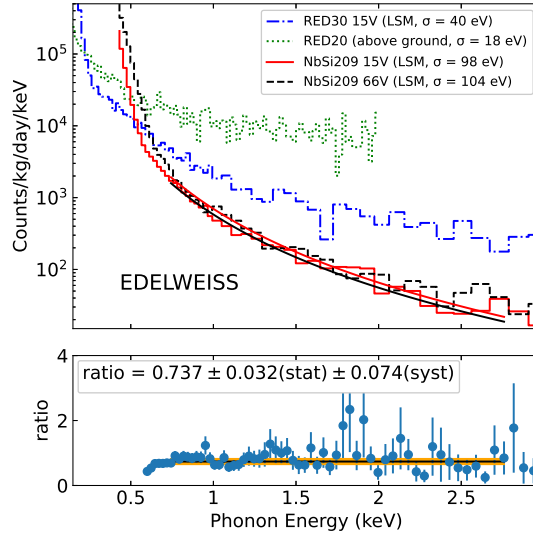


Figure 3: Heat energy spectra of events recorded with RED20 operated above ground at 0 V (green) [22], events with no ionization ($E_{ion} < 0$) for the RED30 detector operated at 15 V (blue), NbSi209 operated at 15 V (red) and 66 V (black). The fitted power law on NbSi spectra when operated at 66 V and 15 V in black and red respectively. The lower figure shows the ratio of NbSi209 distributions recorded at 15 and 66 V (blue) and the associated fit of a constant (black line) and its uncertainty band (orange). This figure has been taken from [30] under RNP/22/OCT/058564 license.

120 3 Prospect : Cryosel

121 The next step in the EDELWEISS Sub-GeV program is to improve both the phonon resolution
 122 of the detectors and the maximum bias at which they can be operated. Developments are
 123 ongoing on the study and reduction of the HO background, as it is one of the limiting factor
 124 in future searches in EDELWEISS. For example, a new detector with a novel phonon sensor,
 125 consisting of a $10 \mu\text{m}$ thin NbSi line called Super Conducting Single Electron Device (SSED)
 126 evaporated on top of a 40 g Ge crystal is under study. Such sensor on a detector operated at
 127 high biases (~ 200 V) has a sensitivity to the athermal phonons triggered by the NTL effect of a
 128 single charge drifting in the strong electric field. This leads to the possibility of evaluating the
 129 purely thermal nature of HO events and thus vetoing this background. The goal of CRYOSEL
 130 is to successfully couple this SSED with a high performance bolometer with 20 eV (RMS) of
 131 baseline resolution. The high-resolution thermal measurement is obtained by using a NTD-Ge
 132 thermistor with 20 eV phonon resolution and applying a bias of 200 V on the detector. This
 133 design is shown in the left panel of Fig 4 along with the structure of the expected electric field.
 134 The right panel shows the expected sensitivity of such design to WIMP nuclei interaction, this
 135 would be competitive with the performances of much more massive detectors and cover areas
 136 of the parameters space that have yet to be excluded. It would also cover all the milestone
 137 models [43] for DM interacting with electron. This program is in it R&D stage and prototypes
 138 are currently being produced and tested.

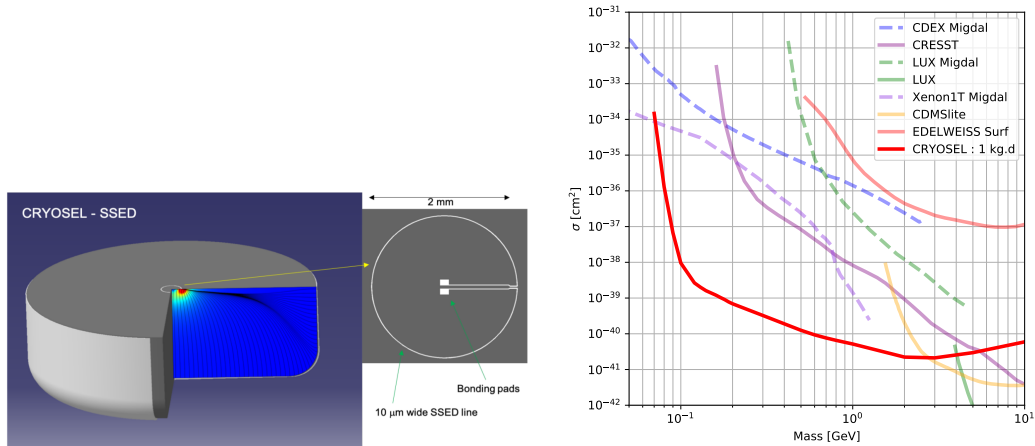


Figure 4: Left: Cryosel 30 g Ge cryogenic detector with covering bottom Al electrode and the SSED detector layout including the simulated electric field-lines. The intensity of the electric field rises significantly under the SSED line (red region) Right: thick red line: projected sensitivity for WIMP-nucleon spin-independent interaction (90%CL) with a 1 kg.d exposure of the CRYOSEL detector (~ 1 month with a 30 g detector).

139 4 Conclusion

140 The EDELWEISS collaboration has searched for DM particle interaction exploiting the Migdal
 141 effect with WIMP masses between $32 \text{ MeV}\cdot\text{c}^{-2}$ and $2 \text{ GeV}\cdot\text{c}^{-2}$ using a 200 g Ge detector op-
 142 erated underground at the LSM. The search constrains a new region of the parameters space
 143 for cross-sections close to 10^{-29} cm^2 . In the context of its EDELWEISS-SubGeV program, the
 144 collaboration is also investigating new methods to significantly reduce HO backgrounds by
 145 improving its ionization resolution with the use of new cold preamplifiers [44], and by devel-
 146 oping NbSi-instrumented devices able to tag the out-of-equilibrium NTL phonons associated
 147 to a single electron.

148 **Funding information** The EDELWEISS project is supported in part by the French Agence
 149 Nationale pour la Recherche (ANR) and the LabEx Lyon Institute of Origins (ANR-10-LABX-
 150 0066) of the Université de Lyon within the program “Investissements d’Avenir” (ANR-11-IDEX-
 151 00007), by the P2IO LabEx (ANR-10-LABX-0038) in the framework “Investissements d’Avenir”
 152 (ANR-11-IDEX-0003-01) managed by the ANR (France), and the Russian Foundation for Ba-
 153 sic Research (grant No. 18-02-00159). This project has received funding from the European
 154 Union’s Horizon 2020 research and innovation program under the Marie Skłodowska-Curie
 155 Grant Agreement No. 838537. B.J. Kavanagh thanks the Spanish Agencia Estatal de Investi-
 156 gación (AEI, MICIU) for the support to the Unidad de Excelencia María de Maeztu Instituto
 157 de Física de Cantabria, ref. MDM-2017-0765. We thank J.P. Lopez (IP2I) and the Physics
 158 Department of Université Lyon 1 for their contribution to the radioactive sources.

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