

The fermionic double smeared null energy condition

Duarte Fragoso^{1*} and Lihan Guo^{1,2†}

1 ITFA, Universiteit van Amsterdam,
Science Park 904, 1098 XH Amsterdam, the Netherlands
2 Dipartimento di Matematica e Fisica, Università Cattolica del Sacro Cuore,
Via della Garzetta 48, 25133 Brescia, Italy

* duarte.fragoso@student.uva.nl, † lihan.guo01@icatt.it

Abstract

Energy conditions are crucial for understanding why exotic phenomena such as traversable wormholes and closed timelike curves remain elusive. In this paper, we prove the Double Smeared Null Energy Condition (DSNEC) for the fermionic free theory in 4-dimensional flat Minkowski space-time, extending previous work on the same energy condition for the bosonic case [1] [2] by adapting Fewster and Mistry's method [3] to the energy-momentum tensor T_{++} . A notable difference from previous works lies in the presence of the $\gamma_0\gamma_+$ matrix in T_{++} , causing a loss of symmetry. This challenge is addressed by making use of its square-root matrix. We provide explicit analytic results for the massless case as well as numerical insights for the mass-dependence of the bound in the case of Gaussian smearing.

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1 Introduction

In general relativity, Einstein's equation itself doesn't impose any restrictions on the form of the energy-momentum tensor $T_{\mu\nu}$. This freedom allows the existence of solutions $G_{\mu\nu}$ that may lead to surprising phenomena, such as macroscopic traversable wormholes [4], closed timelike curves [5] or other causality violations. Energy conditions are essential for explaining why these phenomena have never been observed. The Null Energy Condition (NEC) is particularly important since it is essential in the proof of Penrose's singularity theorem [6] and the second law of black hole thermodynamics (or the Area Theorem) [7] [8] [9].

Previous work has been done for new types of energy conditions namely the Smeared Null Energy Condition (SNEC) [10] and the Double-Smeared Null Energy Condition (DSNEC) [1] which were used to deal with problems that arise when generalizing the NEC to a quantum setup [11–19]. Since this was only studied for the free bosonic theory, in this paper, we will focus on the SNEC and DSNEC for the fermionic theory. Our derivation will closely follow the reasoning of Fewster and Mistry [3] on the Quantum Weak Energy Inequalities for the Dirac field, who deduced a bound for the T_{00} component of the massive fermionic free theory in four-dimensional flat Minkowski space-time, and Wei-Wing *et al* [20], who generalized this result for Minkowski space-time of arbitrary dimensions.

We introduce operators $\mathcal{O}_{\mu i}$ that enable us to express the smeared energy-momentum tensor as the difference between a positive semi-definite operator and a c-number. The primary challenge in defining these operators arises from the presence of the $\gamma_0\gamma_+$ matrix in T_{++} , which reduces the symmetry of the problem. This obstacle is overcome by incorporating the square-root matrix of $\gamma_0\gamma_+$ in the definition of $\mathcal{O}_{\mu i}$.

In brief, the structure of the paper is as follows: In Section 2, we undertake the derivation outlined above and obtain an inequality for the once-smeared T_{++} . However, this inequality is completely trivial, i.e. the lower bound obtained is $-\infty$. We address this issue in Section 3.1 by applying the smearing in two directions, providing a new energy condition:

$$\langle T_{f_+f_-} \rangle \geq -\frac{2}{\pi^3} \int_0^\infty du \int_{\frac{m^2}{u}}^\infty dv \left(\frac{vu^3}{6} - \frac{m^2u^2}{2} + \frac{m^4u}{2v} - \frac{m^6}{6v^2} \right) |\hat{g}_+(u)|^2 |\hat{g}_-(v)|^2, \quad (1)$$

where f_\pm are smearing functions in space-time coordinates and \hat{g}_\pm denotes the Fourier transform of $\sqrt{f_\pm}$. Additionally, we present explicit results for the massless case in Section 3.2, where we employ a Gaussian distribution as the smearing function and derive a bound that depends rationally on the standard deviations,

$$\langle T_{f_+f_-} \rangle \geq -\frac{1}{12\pi^3\sigma_+^3\sigma_-}. \quad (2)$$

Finally, in Section 3.3, we provide numerical results concerning the mass-dependence of the bound. In particular, we observe that for large masses, the bound asymptotically tends to zero.

2 Derivation of the smearing null energy condition

In this section, we will derive a bound for the T_{++} component of the energy-momentum tensor when smeared over the x^+ -direction¹. The quantum field theory considered is the free fermion in Minkowski flat space-time. Note that, despite the bound derived being trivial, the idea can and will be used to deduce a non-trivial bound in Section 3.1.

¹The light-cone variables x^+ and x^- are defined in Appendix A, as well as the light-cone momentum coordinates k^+ and k^- .

52 First, let us write the symmetrized version of the energy-momentum tensor for the free
53 fermion (the Belinfante tensor):

$$T_{\mu\nu} = \frac{i}{4}(\bar{\psi}\gamma_\mu\partial_\nu\psi - \partial_\nu\bar{\psi}\gamma_\mu\psi + \bar{\psi}\gamma_\nu\partial_\mu\psi - \partial_\mu\bar{\psi}\gamma_\nu\psi). \quad (3)$$

54 In particular, we are interested in the light-cone component,

$$T_{++} = \frac{i}{2}(\psi^\dagger A\partial_+\psi - \partial_+\psi^\dagger A\psi), \quad (4)$$

55 where we define $A = \gamma_0\gamma_+$.

56 The decomposition of the fermionic quantum field into Fourier modes yields the following:

$$\psi(x) = \sum_{k,\alpha} b_\alpha(k)u^\alpha(k)e^{-ik\cdot x} + d_\alpha^\dagger(k)v^\alpha(k)e^{ik\cdot x}, \quad (5)$$

57 where here we are considering discrete Dirac quantization in a box of side L . At the end of
58 the derivation, we will take the continuous limit at $L \rightarrow +\infty$.

59 Now, with these expansions, we can expand the first term of T_{++} ,

$$\begin{aligned} \frac{i}{2}(\psi^\dagger A\partial_+\psi) &= \frac{1}{2} \sum_{k,\tilde{k},\alpha,\alpha'} \tilde{k}_+ b_\alpha^\dagger(k)b_{\alpha'}(\tilde{k})u_\alpha^\dagger(k)Au_{\alpha'}^\dagger(\tilde{k})e^{i(k-\tilde{k})\cdot x} \\ &\quad - \tilde{k}_+ b_\alpha^\dagger(k)d_{\alpha'}(\tilde{k})u_\alpha^\dagger(k)Av_{\alpha'}^\dagger(\tilde{k})e^{i(k+\tilde{k})\cdot x} \\ &\quad + \tilde{k}_+ d_\alpha(k)b_{\alpha'}(\tilde{k})v_\alpha^\dagger(k)Au_{\alpha'}^\dagger(\tilde{k})e^{-i(k+\tilde{k})\cdot x} \\ &\quad - \tilde{k}_+ d_\alpha(k)d_{\alpha'}^\dagger(\tilde{k})v_\alpha^\dagger(k)Av_{\alpha'}^\dagger(\tilde{k})e^{-i(k-\tilde{k})\cdot x}, \end{aligned} \quad (6)$$

60 and similarly for the second term. Normal ordering will switch $d_\alpha(k)$ with $d_{\alpha'}^\dagger(\tilde{k})$ providing an
61 additional minus sign:

$$\begin{aligned} :T_{++}: &:= \frac{1}{2} \sum_{k,\tilde{k},\alpha,\alpha'} (k_+ + \tilde{k}_+) [b_\alpha^\dagger(k)b_\alpha(\tilde{k})u_\alpha^\dagger(k)Au_{\alpha'}^\dagger(\tilde{k})e^{i(k-\tilde{k})\cdot x} \\ &\quad + d_{\alpha'}^\dagger(\tilde{k})d_\alpha(k)v_\alpha^\dagger(k)Av_{\alpha'}^\dagger(\tilde{k})e^{-i(k-\tilde{k})\cdot x}] \\ &\quad + (k_+ - \tilde{k}_+) [d_\alpha(k)b_\alpha(\tilde{k})v_\alpha^\dagger(k)Au_{\alpha'}^\dagger(\tilde{k})e^{-i(k+\tilde{k})\cdot x} \\ &\quad - b_\alpha^\dagger(k)d_{\alpha'}(\tilde{k})u_\alpha^\dagger(k)Av_{\alpha'}^\dagger(\tilde{k})e^{i(k+\tilde{k})\cdot x}]. \end{aligned} \quad (7)$$

62 We are interested in smear T_{++} in the x^+ -direction. So, let us put all the other inputs to zero:

$$\begin{aligned} :T_{++}:(x^+, 0) &= \frac{1}{2} \sum_{k,\tilde{k},\alpha,\alpha'} (k_+ + \tilde{k}_+) [b_\alpha^\dagger(k)b_\alpha(\tilde{k})u_\alpha^\dagger(k)Au_{\alpha'}^\dagger(\tilde{k})e^{i(k_+-\tilde{k}_+)\cdot x^+} \\ &\quad + d_{\alpha'}^\dagger(\tilde{k})d_\alpha(k)v_\alpha^\dagger(k)Av_{\alpha'}^\dagger(\tilde{k})e^{-i(k_+-\tilde{k}_+)\cdot x^+}] \\ &\quad + (k_+ - \tilde{k}_+) [d_\alpha(k)b_\alpha(\tilde{k})v_\alpha^\dagger(k)Au_{\alpha'}^\dagger(\tilde{k})e^{-i(k_++\tilde{k}_+)\cdot x^+} \\ &\quad - b_\alpha^\dagger(k)d_{\alpha'}(\tilde{k})u_\alpha^\dagger(k)Av_{\alpha'}^\dagger(\tilde{k})e^{i(k_++\tilde{k}_+)\cdot x^+}]. \end{aligned} \quad (8)$$

63 For general configurations, the expression above is point-wise unbounded from below, so we
64 have to introduce a smearing. Define, for a smearing function f for which we assume to have
65 the positivity condition $f = g^2$ for some other real function g , the smeared energy-momentum
66 tensor component:

$$T_f = \int_{-\infty}^{+\infty} dx^+ :T_{++}:(x^+, 0)f(x^+). \quad (9)$$

67 By the definition of the fourier transform, $\hat{f}(k) = \int_{-\infty}^{+\infty} dx f(x) e^{-ikx}$, we get the following
68 expression:

$$T_f = \frac{1}{2} \sum_{k, \tilde{k}, \alpha, \alpha'} (k_+ + \tilde{k}_+) [b_\alpha^\dagger(k) b_\alpha(\tilde{k}) u_\alpha^\dagger(k) A u_{\alpha'}(\tilde{k}) \hat{f}(\tilde{k}_+ - k_+) + d_{\alpha'}^\dagger(\tilde{k}) d_\alpha(k) v_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k}) \hat{f}(k_+ - \tilde{k}_+)] + (k_+ - \tilde{k}_+) [d_\alpha(k) b_\alpha(\tilde{k}) v_\alpha^\dagger(k) A u_{\alpha'}(\tilde{k}) \hat{f}(k_+ - \tilde{k}_+) - b_\alpha^\dagger(k) d_{\alpha'}(\tilde{k}) u_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k}) \hat{f}(-k_+ - \tilde{k}_+)]. \quad (10)$$

69 Note that the matrix A has eigenvalues $\lambda_1 = \lambda_2 = 0$, $\lambda_3 = \lambda_4 = 2$, so it is positive semi-definite.
70 Then it is possible to find the matrix B such that $B^\dagger B = B^\dagger B = A$, i.e. B is the square root matrix
71 of A . It is easy to obtain B explicitly but we will only use its existence.

72 Define the following family of operators for $i \in \{1, 2, 3, 4\}$ and $\mu \in \mathbb{R}$:

$$\mathcal{O}_{\mu i} = \sum_{k, \alpha} \hat{g}(-k_+ + \mu) \overline{b_\alpha(k) (B u_\alpha(k))_i} + \hat{g}(k_+ + \mu) d_\alpha^\dagger(k) (B v_\alpha(k))_i, \quad (11)$$

$$\mathcal{O}_{\mu i}^\dagger = \sum_{k, \alpha} \hat{g}(-k_+ + \mu) b_\alpha^\dagger(k) (u_\alpha^\dagger(k) B^\dagger)_i + \hat{g}(k_+ + \mu) d_\alpha^\dagger(k) (v_\alpha^\dagger(k) B^\dagger)_i. \quad (12)$$

73 So that \mathcal{O}_μ for a fixed $\mu \in \mathbb{R}$ is a four-dimensional vector of operators, and \mathcal{O}_μ^\dagger a co-vector of
74 the same type. Using the anti-commutation relations of the fields, one finds

$$\begin{aligned} \mathcal{O}_\mu^\dagger \mathcal{O}_\mu = S_\mu^v \mathbb{1} + \sum_{k, \tilde{k}, \alpha, \alpha'} & \hat{g}(-k_+ + \mu) \overline{\hat{g}(-\tilde{k}_+ + \mu) b_\alpha^\dagger(k) b_{\alpha'}(\tilde{k}) u_\alpha^\dagger(k) A u_{\alpha'}(\tilde{k})} \\ & - \hat{g}(k_+ + \mu) \overline{\hat{g}(\tilde{k}_+ + \mu) d_{\alpha'}^\dagger(k) d_\alpha(\tilde{k}) v_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k})} \\ & + \hat{g}(k_+ + \mu) \overline{\hat{g}(-\tilde{k}_+ + \mu) d_\alpha(k) b_{\alpha'}(\tilde{k}) v_\alpha^\dagger(k) A u_{\alpha'}(\tilde{k})} \\ & + \hat{g}(-k_+ + \mu) \overline{\hat{g}(\tilde{k}_+ + \mu) b_\alpha^\dagger(k) d_{\alpha'}^\dagger(\tilde{k}) u_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k})}, \end{aligned}$$

75 where we have defined

$$S_\mu^v \equiv \sum_{k, \alpha} \hat{g}(k_+ + \mu) \hat{g}(\tilde{k}_+ + \mu) \delta_{\alpha, \alpha'} \delta_{k, \tilde{k}} v_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k}) = \sum_{k, \alpha} \hat{g}(k_+ + \mu) \hat{g}(k_+ + \mu) v_\alpha^\dagger(k) A v_\alpha(k). \quad (13)$$

76 Proven in the literature [3], the following lemma allows us to recover T_f .

Lemma 1 *Let $f = g^2$ with g a real, smooth, compactly-supported² function. Then the following identity holds:*

$$(k_+ + \tilde{k}_+) \hat{f}(k_+ - \tilde{k}_+) = \frac{1}{\pi} \int_{-\infty}^{\infty} d\mu \mu \hat{g}(k_+ - \mu) \overline{\hat{g}(\tilde{k}_+ - \mu)}.$$

77 Using this lemma,

$$T_f = \frac{1}{2\pi} \int_{-\infty}^{\infty} d\mu \mu (\mathcal{O}_\mu^\dagger \mathcal{O}_\mu - S_\mu^v \mathbb{1}). \quad (14)$$

78 One can then compute the anti-commutator of the operator \mathcal{O} ,

$$\begin{aligned} \{\mathcal{O}_{\mu i}^\dagger, \mathcal{O}_{\mu i}\} &= \sum_{k, \tilde{k}, \alpha, \alpha'} (\hat{g}(-k_+ + \mu) \overline{\hat{g}(-\tilde{k}_+ + \mu) \delta_{\alpha, \alpha'} \delta_{k, \tilde{k}} u_\alpha^\dagger(k) A u_{\alpha'}(\tilde{k})} \\ &+ \hat{g}(k_+ + \mu) \overline{\hat{g}(\tilde{k}_+ + \mu) \delta_{\alpha, \alpha'} \delta_{k, \tilde{k}} v_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k})}) \mathbb{1} \\ &= (S_{-\mu}^u + S_\mu^v) \mathbb{1}. \end{aligned} \quad (15)$$

²Note that the assumption of compact support is stronger than necessary. For example, the lemma still holds for g a Gaussian distribution, since the rapid decay of the function secures convergence of the integral.

79 Note that since g is real valued, $|\hat{g}|$ is even.

80 Using the anti-commutation relation obtained above, we can split the integral,

$$\begin{aligned} T_f &= \frac{1}{2\pi} \int_{-\infty}^{\infty} d\mu \mu (\mathcal{O}_{\mu i}^\dagger \mathcal{O}_{\mu i} - S_\mu^v \mathbb{1}) \\ &= \frac{1}{2\pi} \int_0^{\infty} d\mu \mu (\mathcal{O}_{\mu i}^\dagger \mathcal{O}_{\mu i} - S_\mu^v \mathbb{1}) + \frac{1}{2\pi} \int_{-\infty}^0 d\mu \mu (S_{-\mu}^u \mathbb{1} - \mathcal{O}_{\mu i} \mathcal{O}_{\mu i}^\dagger). \end{aligned} \quad (16)$$

81 Notice that if $\mu \geq 0$, then $\mu \langle \mathcal{O}_{\mu i}^\dagger \mathcal{O}_{\mu i} \rangle_\psi \geq 0$, and similarly if $\mu \leq 0$, then $-\mu \langle \mathcal{O}_{\mu i} \mathcal{O}_{\mu i}^\dagger \rangle_\psi \geq 0$, for
82 any state $|\psi\rangle$. Hence,

$$\begin{aligned} \langle T_f \rangle_\psi &\geq -\frac{1}{2\pi} \int_0^{\infty} d\mu \mu S_\mu^v + \frac{1}{2\pi} \int_{-\infty}^0 d\mu \mu S_{-\mu}^u \\ &= -\frac{1}{2\pi} \int_0^{\infty} d\mu \mu (S_\mu^v + S_\mu^u). \end{aligned} \quad (17)$$

83 The computation preformed in appendix B shows that

$$\sum_\alpha u_\alpha^\dagger(k) A u_\alpha(k) = \frac{2}{V} \left(1 - \frac{k^1}{\omega_k}\right), \quad (18)$$

$$\sum_\alpha v_\alpha^\dagger(k) A v_\alpha(k) = \frac{2}{V} \left(1 - \frac{k^1}{\omega_k}\right). \quad (19)$$

84 Finally, plugging it in the expression for S^u and S^v and taking the continuous limit $\frac{1}{V} \sum_{\vec{k}} \rightarrow \int \frac{d^3\vec{k}}{(2\pi)^3}$,
85 we come to the conclusion that

$$\langle T_f \rangle_\psi \geq -\frac{2}{\pi} \int_0^{\infty} d\mu \mu \int \frac{d^3k}{(2\pi)^3} |\hat{g}(k_+ + \mu)|^2 \left(1 - \frac{k^1}{\omega_k}\right). \quad (20)$$

86 We will denote this bound as \mathcal{B}_1 , where the subscript 1 represents the number of smearing
87 directions, and the dependence on the smearing function is implicit.

88 Unfortunately, the integral obtained in equation (20) diverges. By definition, $k_+ = \frac{1}{2}(\omega_k + k_1)$
89 $= \frac{1}{2}(\omega_k - k^1) = \frac{1}{2}(\sqrt{k_1^2 + k_2^2 + k_3^2 + m^2} - k^1)$, so we can change the integral variable accord-
90 ingly. Using that the measure transforms as $dk_+ = \frac{1}{2}\left(\frac{k^1}{\omega_k} - 1\right)dk^1$, we have that the bound is
91 proportional to

$$\mathcal{B}_1 \propto - \int_0^{\infty} d\mu \int dk_+ \mu |\hat{g}(k_+ + \mu)|^2 \int dk^2 dk^3. \quad (21)$$

92 We can then note that the integrals in k_2 and k_3 are decoupled and they will contribute with
93 the volume of the space in those directions. Since the integral in μ and k_+ does not vanish
94 for non-trivial smearing functions, the expression above diverges, which means the bound is
95 completely trivial.

96 This outcome is clearly unsatisfactory since our aim was to derive a non-trivial lower
97 bound. Such a bound would allow us to explore the extent to which the Null Energy Con-
98 dition (NEC) is violated within the framework of free fermionic quantum field theory.

99 However, this divergence is not unexpected, drawing an analogy with the divergence of
100 the bound of the free bosonic theory [10], when the UV cut-off approaches zero. The main
101 issue is that, in order to obtain a convergent integral, it is necessary to fully smear it in the
102 time direction. Note that t is linearly dependent on x^+ and x^- due to $x^+ + x^- = t$. Hence, we
103 expect that, if we smear T_{++} in both light-cone directions, we will obtain a convergent lower
104 bound.

105 3 Double smeared null energy condition

106 3.1 Derivation of the non-trivial bound

107 In this section, we will prove a convergent lower bound, \mathcal{B}_2 , by smearing T_{++} in both the
 108 x^+ and x^- -direction. In general, the smearing function can be of the form $f(x^+, x^-)$. For
 109 practical purposes, we restrict our argument to the case where f is separable, i.e. the function
 110 factors multiplicatively $f(x^+, x^-) = f_+(x^+)f_-(x^-)$. In other words, we are now interested in
 111 obtaining an upper bound for

$$T_{f_+f_-} = \int dx^+ \int dx^- : T_{++} : (x^+, x^-, 0) f_+(x^+) f_-(x^-). \quad (22)$$

112 Keeping in mind that $e^{ik \cdot x} = e^{ik_+ x^+} e^{ik_- x^-} e^{-ik_\perp \cdot x^\perp}$ and using the definition of Fourier trans-
 113 form, we can carry out the same procedure as before to write

$$\begin{aligned} T_{f_+f_-} = & \frac{1}{2} \sum_{k, \tilde{k}, \alpha, \alpha'} (k_+ + \tilde{k}_+) [b_\alpha^\dagger(k) b_\alpha(\tilde{k}) u_\alpha^\dagger(k) A u_{\alpha'}(\tilde{k}) \hat{f}_+(\tilde{k}_+ - k_+) \hat{f}_-(\tilde{k}_- - k_-) + d_{\alpha'}^\dagger(\tilde{k}) d_\alpha(k) \\ & v_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k}) \hat{f}_+(k_+ - \tilde{k}_+) \hat{f}_-(k_- - \tilde{k}_-)] + (k_- - \tilde{k}_-) [d_\alpha(k) b_\alpha(\tilde{k}) v_\alpha^\dagger(k) A u_{\alpha'}(\tilde{k}) \\ & \hat{f}_+(k_+ - \tilde{k}_+) \hat{f}_-(k_- - \tilde{k}_-) - b_\alpha^\dagger(k) d_{\alpha'}(\tilde{k}) u_\alpha^\dagger(k) A v_{\alpha'}(\tilde{k}) \hat{f}_+(-k_+ - \tilde{k}_+) \hat{f}_-(-k_- - \tilde{k}_-)]. \end{aligned} \quad (23)$$

114 Denoting $g_\pm = \sqrt{f_\pm}$, we define the new operators:

$$\mathcal{O}_{\mu i} = \sum_{k, \alpha} \overline{\hat{G}(-k + \mu) b_\alpha(k) (B u_\alpha(k))_i} + \overline{\hat{G}(k + \mu) d_\alpha^\dagger(k) (B v_\alpha(k))_i} \quad (24)$$

$$\mathcal{O}_{\mu i}^\dagger = \sum_{k, \alpha} \hat{G}(-k + \mu) b_\alpha^\dagger(k) (u_\alpha^\dagger(k) B^\dagger)_i + \hat{G}(k + \mu) d_\alpha^\dagger(k) (v_\alpha^\dagger(k) B^\dagger)_i, \quad (25)$$

115 where $\hat{G}(k + \mu) = \hat{g}_+(k_+ + \mu_+) \hat{g}_-(k_- + \mu_-)$ and μ_\pm are two dummy variables that shall be
 116 integrated out at the end. We will denote $\mu = (\mu_+, \mu_-)$.

117 Since the Fourier transform of the product is the convolution, $f = g^2$ implies that $\hat{f} = \hat{g} * \hat{g}$
 118 $= \int d\mu \hat{g}(\mu) \hat{g}(k - \mu)$, so we have

$$\hat{f}_-(k_- - \tilde{k}_-) = \int d\mu \hat{g}_-(\mu) \overline{\hat{g}_-(\mu - (k_- - \tilde{k}_-))} \quad (26)$$

$$= \int d\mu_- \hat{g}_-(k_- - \mu_-) \overline{\hat{g}_-(\tilde{k}_- - \mu_-)}, \quad (27)$$

119 where we changed the variable $\mu = k_- - \mu_-$ and used that since g_- is real, $\overline{\hat{g}_-(x)} = \hat{g}_-(-x)$.
 120 Applying lemma 1 to f_+ and g_+ , one obtains

$$(k_+ + \tilde{k}_+) \hat{f}_+(k_+ - \tilde{k}_+) = \frac{1}{\pi} \int_{-\infty}^{\infty} d\mu_+ \mu_+ \hat{g}_+(k_+ - \mu_+) \overline{\hat{g}_+(\tilde{k}_+ - \mu_+)}. \quad (28)$$

121 Then for the double smearing case, applying lemma 1 again, we have

$$(k_+ + \tilde{k}_+) \hat{f}_+(k_+ - \tilde{k}_+) \hat{f}_-(k_- - \tilde{k}_-) = \int d\mu_+ d\mu_- \mu_+ \hat{g}_+(k_+ - \mu_+) \overline{\hat{g}_+(\tilde{k}_+ - \mu_+) \hat{g}_-(k_- - \mu_-) \overline{\hat{g}_-(\tilde{k}_- - \mu_-)}}. \quad (29)$$

122 In an analogous way as before, we can prove that

$$T_{f_+, f_-} = \int_0^{+\infty} d\mu_+ \int_0^{+\infty} d\mu_- \mu_+ (\mathcal{O}_\mu^\dagger \mathcal{O}_\mu - S_\mu^v \mathbb{1}), \quad (30)$$

123 where now the definition of S_μ^v is different from the once-smear case:

$$S_\mu^v = \sum_{k, \alpha} |\hat{g}_+(k_+ + \mu_+)|^2 |\hat{g}_-(k_- + \mu_-)|^2 v_\alpha^\dagger(k) A v_\alpha(k) \quad (31)$$

$$= \frac{2}{V} \sum_k |\hat{g}_+(k_+ + \mu_+)|^2 |\hat{g}_-(k_- + \mu_-)|^2 \left(1 - \frac{k^1}{\omega_k}\right) \quad (32)$$

124 and the anti-commutator is what we expect, with the new definitions of S_μ^v and S_μ^u :

$$\{\mathcal{O}_{\mu i}^\dagger, \mathcal{O}_{\mu i}\} = (S_{-\mu}^u + S_\mu^v) \mathbb{1}. \quad (33)$$

125 In the end, we get the following bound,

$$\langle T_{f_+, f_-} \rangle \geq -\frac{2}{\pi} \int_0^{+\infty} d\mu_+ \int_0^{+\infty} d\mu_- \int \frac{d^3 \vec{k}}{(2\pi)^3} \mu_+ |\hat{g}_+(k_+ + \mu_+)|^2 |\hat{g}_-(k_- + \mu_-)|^2 \left(1 - \frac{k^1}{\omega_k}\right). \quad (34)$$

126 We know that

$$k_+ = \frac{1}{2}(\omega_k + k_1), \quad (35)$$

$$k_- = \frac{1}{2}(\omega_k - k_1), \quad (36)$$

$$\omega_k^2 = k_1^2 + k_2^2 + k_3^2 + m^2. \quad (37)$$

127 So setting $k_\perp := \sqrt{k_2^2 + k_3^2}$ we obtain,

$$4k_+ k_- = k_2^2 + k_3^2 + m^2 = k_\perp^2 + m^2, \quad (38)$$

128 and it's straightforward to find that

$$d(k_+ k_-) = \frac{1}{2} k_\perp dk_\perp. \quad (39)$$

129 Since $dk_1 \wedge dk_2 \wedge dk_3 = dk_1 \wedge dk_\perp \wedge k_\perp d\theta$, we can rewrite part of our integral measure in
130 terms of dk^+ , dk^- and $d\theta$, i.e. $dk_1 \wedge dk_2 \wedge dk_3 = 2(k_+ + k_-) dk_- \wedge dk_+ \wedge d\theta$.

131 With all the considerations discussed above, one can change the variable of the integral in
132 the double-smear bound (34),

$$\begin{aligned} \langle T_{f_+, f_-} \rangle &\geq -\frac{2}{\pi} \int_0^{+\infty} d\mu_+ \int_0^{+\infty} d\mu_- \int_{\mathcal{D}} 2(k_- + k_+) dk_- dk_+ 2\pi \frac{1}{(2\pi)^3} \mu_+ |\hat{g}_+(k_+ + \mu_+)|^2 |\hat{g}_-(k_- + \mu_-)|^2 \left(\frac{2k_+}{k_- + k_+}\right) \\ &= -\frac{2}{\pi^3} \int_0^{+\infty} d\mu_+ \int_0^{+\infty} d\mu_- \int_{\mathcal{D}} dk_+ dk_- \mu_+ k_+ |\hat{g}_+(k_+ + \mu_+)|^2 |\hat{g}_-(k_- + \mu_-)|^2 \end{aligned} \quad (40)$$

133 where the integration domain is $\mathcal{D} = \{k_\pm \geq 0 | k_+ k_- \geq m^2\}$.

134 Equation (40) can be further simplified by changing variables. Setting $u = k_+ + \mu_+$ and
135 $v = k_- + \mu_-$,

$$\langle T_{f_+, f_-} \rangle \geq -\frac{2}{\pi^3} \int_0^\infty du \int_{\frac{m^2}{u}}^\infty dv \int_{\frac{m^2}{v}}^u dk_+ \int_{\frac{m^2}{k_+}}^v dk_- (u - k_+) k_+ |\hat{g}_+(u)|^2 |\hat{g}_-(v)|^2. \quad (41)$$

136 Performing the k^- and k^+ integrals, we can present our main result.

$$\langle T_{f_+f_-} \rangle \geq -\frac{2}{\pi^3} \int_0^\infty du \int_{\frac{m^2}{u}}^\infty dv \left(\frac{vu^3}{6} - \frac{m^2u^2}{2} + \frac{m^4u}{2v} - \frac{m^6}{6v^2} \right) |\hat{g}_+(u)|^2 |\hat{g}_-(v)|^2 \quad (42)$$

137 It's worth mentioning that the form of our bound looks simpler than the result for the bosonic
138 case in [1] and [2]. In the next subsections, we will explore this bound in more specific
139 circumstances, where we can find simpler analytic expressions or numerical results.

140 3.2 Massless, Gaussian-smearred bound

141 Let us first investigate the massless case, where we can compute some analytic results for
142 specific smearing functions. Take the Gaussian function $|\hat{g}_+(u)|^2 = \sigma_+ e^{-(\sigma_+ u)^2}$ (similarly
143 $|\hat{g}_-(v)|^2 = \sigma_- e^{-(\sigma_- v)^2}$)³ as a particular example of smearing function. By changing variables
144 ($\tilde{u} = \sigma_+ u$ and $\tilde{v} = \sigma_- v$), from equation (42) we obtain the expression for the bound,

$$\langle T_{f_+f_-} \rangle \geq -\frac{1}{3\pi^3} \int_0^\infty d\tilde{u} \int_0^\infty d\tilde{v} \frac{1}{\sigma_+^3 \sigma_-} \tilde{u}^3 \tilde{v} e^{-\tilde{u}^2} e^{-\tilde{v}^2} \quad (43)$$

$$= -\frac{1}{12\pi^3 \sigma_+^3 \sigma_-}, \quad (44)$$

145 which turns out to be a satisfactory finite negative number. So, we obtained a non-trivial
146 lower bound for the doubled-smearred T_{++} for the simple case where the smearing is Gaus-
147 sian. Moreover, σ_+ has a larger effect on the bound comparatively to σ_- . This asymmetry of
148 the dependence on the deviations is expected since the energy-momentum tensor component
149 considered has, by definition, a preferred space-time direction.

150 Since large σ_\pm correspond to a wide smearing in space-time, we expect the bound to
151 approach zero. This is indeed in agreement with the well-studied null energy condition (NEC)
152 [21]. On the other hand, in the $\sigma_\pm \rightarrow 0$ case, i.e. there is no smearing in space-time, we obtain
153 a trivial bound. This is expected since the expected value of the energy-momentum evaluated
154 at a particular space-time point is generally unbounded.

155 3.3 Mass dependence of the Gaussian-smearred bound

156 One can also wonder about how this bound, which will now denote by \mathcal{B}_2 , depends on the
157 mass. Let us choose the two smearing functions to be Gaussians with standard deviation $\sigma = 1$,
158 i.e. $|\hat{g}_+(x)|^2 = |\hat{g}_-(x)|^2 = e^{-x^2}$. By dimensional analysis, we have $[\sigma] = -1$. Now the bound
159 takes the following form:

$$\mathcal{B}_2 = -\frac{2}{\pi^3} \int_0^\infty du \int_{\frac{m^2}{u}}^\infty dv \left(\frac{vu^3}{6} - \frac{m^2u^2}{2} + \frac{m^4u}{2v} - \frac{m^6}{6v^2} \right) e^{-(u^2+v^2)}. \quad (45)$$

160 We can numerically integrate the expression above to obtain the following relation between
161 \mathcal{B}_2 and the mass, which is shown in Figure 1. Since we work with natural units, u and v are
162 in the same unit as m , so $[\mathcal{B}_2] = 4$.

163 Note that in the highly massive region, the lower bound approaches zero. We can un-
164 derstand this result in the following qualitative way. Roughly speaking, quantum effects are
165 relevant when the de Broglie wavelength of the particle, $(\frac{h}{mv})$, is much greater than the char-
166 acteristic size of the system, d . In our case, we simply take this d to be the smearing length.

³We are defining our Gaussian function slightly different from the usual form $\frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{1}{2}(\frac{x}{\sigma})^2}$ here. This way, σ_\pm are the smearing lengths in space-time since the Fourier transform of a Gaussian function with .

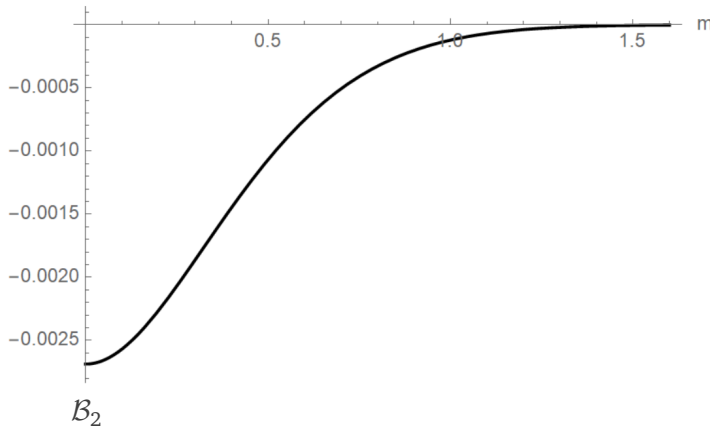


Figure 1: The bound \mathcal{B}_2 as a function of the mass m .

167 For small m , quantum behavior becomes prominent, but as m increases, classical behavior
 168 dominates. Given that the classical case satisfies the Null Energy Condition (NEC), the bound
 169 is anticipated to approach zero as m becomes large, which is verified numerically in the figure
 170 above.

171 4 Conclusion

172 In this work, we investigated the Double Smeared Null Energy Condition for the fermionic
 173 free theory in 4-dimensional flat Minkowski space-time. We first obtained an inequality for the
 174 once-smeared T_{++} . We addressed its triviality later by applying the smearing in two directions,
 175 providing a new energy condition. We offered explicit analytic results for the massless case
 176 and numerical insights for the mass-dependence of the later bound in the case of Gaussian
 177 smearing.

178 Regarding the outlook of this research, as mentioned in [1], understanding the behavior
 179 of DSNEC in interacting field theories is generally still an open question. Since we are using
 180 different methods than those used in [1] to derive the DSNEC for fermions, our approach may
 181 be promising for interacting field theories. This is because our techniques are not as specifically
 182 tailored to treat free field theories.

183 Moreover, our discussion is limited to Minkowski spacetime. Extending the DSNEC to
 184 curved spaces is a highly significant direction, given that it is crucial for its application in
 185 semiclassical gravity. It would then be interesting to explore a generalized version of our
 186 results in different curved spacetimes.

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191 A Conventions

192 In this paper, we work in 4 dimensional Minkowski spacetime with the “mostly minus” signa-
193 ture in natural units ($c = \hbar = 1$).

194 The position light-cone coordinates are given by $x^\pm = t \pm x^1$. The corresponding metric
195 tensor for the coordinates (x^+, x^-, x^2, x^3) is,

$$g_{\mu\nu} = \begin{pmatrix} 0 & 1/2 & 0 & 0 \\ 1/2 & 0 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}. \quad (\text{A.1})$$

196 In momentum space we will denote $k_\pm = \frac{1}{2}(\omega_k \pm k_1)$ such that $k \cdot x = k_+ x^+ + k_- x^- + x_i \cdot k^i$.
197 In our convention, the 4-by-4 gamma matrices are defined as

$$\gamma^0 = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}, \quad (\text{A.2})$$

198

$$\gamma^i = \begin{pmatrix} 0 & \sigma_i \\ -\sigma_i & 0 \end{pmatrix}. \quad (\text{A.3})$$

199 Since we are working with the mostly-minus metric we obtain

$$\gamma_0 = \gamma^0, \quad (\text{A.4})$$

200

$$\gamma_i = -\gamma^i. \quad (\text{A.5})$$

201 In particular,

$$\gamma_+ = \gamma_0 + \gamma_1 = \gamma_0 - \gamma^1 = \begin{pmatrix} 1 & 0 & 0 & -1 \\ 0 & 1 & -1 & 0 \\ 0 & 1 & -1 & 0 \\ 1 & 0 & 0 & -1 \end{pmatrix}, \quad (\text{A.6})$$

202 and we define

$$A = \gamma_0 \gamma_+ = \begin{pmatrix} 1 & 0 & 0 & -1 \\ 0 & 1 & -1 & 0 \\ 0 & -1 & 1 & 0 \\ -1 & 0 & 0 & 1 \end{pmatrix} = \mathbb{1} - \mathbb{L}, \quad (\text{A.7})$$

203 where $\mathbb{1}$ is the identity matrix, and \mathbb{L} is the exchange matrix.

204 B Explicit computations for $\sum_\alpha u_\alpha^\dagger(k) A u_\alpha(k)$

205 In Appendix B, the explicit computations for $\sum_\alpha u_\alpha^\dagger(k) A u_\alpha(k)$ will be made.

206 Note that $\sum_\alpha u_\alpha^\dagger(k) A u_\alpha(k) = \frac{2}{V} - \sum_\alpha u_\alpha^\dagger(k) \mathbb{L} u_\alpha(k)$, using the decomposition $A = \mathbb{1} - \mathbb{L}$ and
207 the normalization of $u_\alpha(k)$. Then, we can write

$$u_1(k) = \begin{bmatrix} a \\ Cb \end{bmatrix}, \quad (\text{B.1})$$

208 where we define

$$a = \sqrt{\frac{\omega_k + m}{2\omega_k V}} \begin{bmatrix} 1 \\ 0 \end{bmatrix}, \quad (\text{B.2})$$

$$b = \frac{1}{\sqrt{2\omega_k(\omega_k + m)V}} \begin{bmatrix} 1 \\ 0 \end{bmatrix}, \quad (\text{B.3})$$

$$C = \vec{\sigma} \cdot \vec{k}. \quad (\text{B.4})$$

209 Using this notation we obtain that

$$u_1^\dagger(k) \mathbb{L} u_1(k) = a^\dagger \sigma_1 C b + b^\dagger C^\dagger \sigma_1 a. \quad (\text{B.5})$$

210 The matrix σ_1 appears in the non-zero blocks of \mathbb{L} .

211 Now,

$$a^\dagger \sigma_1 C b = \sqrt{\frac{\omega_k + m}{2\omega_k V}} \begin{bmatrix} 1 & 0 \end{bmatrix} k^1 \frac{1}{\sqrt{2\omega_k(\omega_k + m)V}} \mathbb{1}_{2 \times 2} \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad (\text{B.6})$$

$$+ \sqrt{\frac{\omega_k + m}{2\omega_k V}} \begin{bmatrix} 1 & 0 \end{bmatrix} k^2 \frac{1}{\sqrt{2\omega_k(\omega_k + m)V}} \begin{bmatrix} i & 0 \\ 0 & -i \end{bmatrix} \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad (\text{B.7})$$

$$+ \sqrt{\frac{\omega_k + m}{2\omega_k V}} \begin{bmatrix} 1 & 0 \end{bmatrix} k^3 \frac{1}{\sqrt{2\omega_k(\omega_k + m)V}} \begin{bmatrix} 0 & -1 \\ 1 & 0 \end{bmatrix} \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad (\text{B.8})$$

$$= \frac{1}{V} \frac{1}{2\omega_k} (k^1 + ik^2). \quad (\text{B.9})$$

212 For the second term, we have something similar:

$$a^\dagger \sigma_1 C b = \sqrt{\frac{\omega_k + m}{2\omega_k V}} \begin{bmatrix} 1 & 0 \end{bmatrix} k^1 \frac{1}{\sqrt{2\omega_k(\omega_k + m)V}} \mathbb{1}_{2 \times 2} \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad (\text{B.10})$$

$$+ \sqrt{\frac{\omega_k + m}{2\omega_k V}} \begin{bmatrix} 1 & 0 \end{bmatrix} k^2 \frac{1}{\sqrt{2\omega_k(\omega_k + m)V}} \begin{bmatrix} -i & 0 \\ 0 & i \end{bmatrix} \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad (\text{B.11})$$

$$+ \sqrt{\frac{\omega_k + m}{2\omega_k V}} \begin{bmatrix} 1 & 0 \end{bmatrix} k^3 \frac{1}{\sqrt{2\omega_k(\omega_k + m)V}} \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix} \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad (\text{B.12})$$

$$= \frac{1}{V} \frac{1}{2\omega_k} (k^1 - ik^2). \quad (\text{B.13})$$

213 With this we conclude that

$$u_1^\dagger(k) \mathbb{L} u_1(k) = \frac{1}{V} \frac{1}{2\omega_k} k^1. \quad (\text{B.14})$$

214 The analogous computation for $u_2(k)$ yields the same result. Summing both, we obtain:

$$u_\alpha(k)^\dagger(k) \mathbb{L} u_\alpha(k) = \frac{1}{V} \frac{1}{\omega_k} k^1. \quad (\text{B.15})$$

215 In the same way, we obtain the exact same result for $v_\alpha(k)^\dagger(k) \mathbb{L} v_\alpha(k)$.

216 References

- 217 [1] J. R. Fliss and B. Freivogel, *Semi-local bounds on null energy in qft*, SciPost Physics **12**(3),
218 084 (2022), doi:[10.21468/SciPostPhys.12.3.084](https://doi.org/10.21468/SciPostPhys.12.3.084).

- 219 [2] J. R. Fliss, B. Freivogel and E.-A. Kontou, *The double smeared null energy condition*,
220 SciPost Physics **14**(2), 024 (2023), doi:[10.21468/SciPostPhys.14.2.024](https://doi.org/10.21468/SciPostPhys.14.2.024).
- 221 [3] C. Fewster and B. Mistry, *Quantum weak energy inequalities for the dirac field in flat*
222 *spacetime*, Physical Review D **68**, 105010 (2003), doi:[10.1103/PhysRevD.68.105010](https://doi.org/10.1103/PhysRevD.68.105010).
- 223 [4] L. Ford and T. A. Roman, *Quantum field theory constrains traversable wormhole geometries*,
224 Physical Review D **53**(10), 5496 (1996), doi:[10.1103/PhysRevD.53.5496](https://doi.org/10.1103/PhysRevD.53.5496).
- 225 [5] F. Lobo and P. Crawford, *Time, closed timelike curves and causality*, The nature of time:
226 Geometry, physics and perception pp. 289–296 (2003), doi:[10.1007/978-94-010-0155-7_30](https://doi.org/10.1007/978-94-010-0155-7_30).
227
- 228 [6] R. Penrose, *Gravitational collapse and space-time singularities*, Physical Review Letters
229 **14**(3), 57 (1965), doi:[10.1103/PhysRevLett.14.57](https://doi.org/10.1103/PhysRevLett.14.57).
- 230 [7] S. W. Hawking, *Black holes in general relativity*, Communications in Mathematical Physics
231 **25**, 152 (1972), doi:[10.1007/BF01877517](https://doi.org/10.1007/BF01877517).
- 232 [8] J. M. Bardeen, B. Carter and S. W. Hawking, *The four laws of black hole mechanics*,
233 Communications in mathematical physics **31**, 161 (1973), doi:[10.1007/BF01645742](https://doi.org/10.1007/BF01645742).
- 234 [9] R. M. Wald, *Quantum field theory in curved spacetime and black hole thermodynamics*,
235 University of Chicago press, ISBN 978-0-226-87027-4 (1994).
- 236 [10] B. Freivogel and D. Krommydas, *The smeared null energy condition*, Journal of High
237 Energy Physics **2018**(12), 1 (2018), doi:[10.1007/JHEP12\(2018\)067](https://doi.org/10.1007/JHEP12(2018)067).
- 238 [11] N. Graham and K. D. Olum, *Achronal averaged null energy condition*, Physical Review D
239 **76**(6), 064001 (2007), doi:[10.1103/PhysRevD.76.064001](https://doi.org/10.1103/PhysRevD.76.064001).
- 240 [12] A. C. Wall, *Proving the achronal averaged null energy condition from the generalized second*
241 *law*, Physical Review D **81**(2), 024038 (2010), doi:[10.1103/PhysRevD.81.024038](https://doi.org/10.1103/PhysRevD.81.024038).
- 242 [13] D. Urban and K. D. Olum, *Averaged null energy condition violation in a conformally flat*
243 *spacetime*, Physical Review D **81**(2), 024039 (2010), doi:[10.1103/PhysRevD.81.024039](https://doi.org/10.1103/PhysRevD.81.024039).
- 244 [14] E.-A. Kontou and K. D. Olum, *Proof of the averaged null energy condition in a classical*
245 *curved spacetime using a null-projected quantum inequality*, Physical Review D **92**(12),
246 124009 (2015), doi:[10.1103/PhysRevD.92.124009](https://doi.org/10.1103/PhysRevD.92.124009).
- 247 [15] T. Faulkner, R. G. Leigh, O. Parrikar and H. Wang, *Modular hamiltonians for deformed half-*
248 *spaces and the averaged null energy condition*, Journal of High Energy Physics **2016**(9),
249 1 (2016), doi:[10.1007/JHEP09\(2016\)038](https://doi.org/10.1007/JHEP09(2016)038).
- 250 [16] R. Bousso, Z. Fisher, S. Leichenauer and A. C. Wall, *Quantum focusing conjecture*, Physical
251 Review D **93**(6), 064044 (2016), doi:[10.1103/PhysRevD.93.064044](https://doi.org/10.1103/PhysRevD.93.064044).
- 252 [17] R. Bousso, Z. Fisher, J. Koeller, S. Leichenauer and A. C. Wall, *Proof of*
253 *the quantum null energy condition*, Physical Review D **93**(2), 024017 (2016),
254 doi:[10.1103/PhysRevD.93.024017](https://doi.org/10.1103/PhysRevD.93.024017).
- 255 [18] S. Balakrishnan, T. Faulkner, Z. U. Khandker and H. Wang, *A general proof of the*
256 *quantum null energy condition*, Journal of High Energy Physics **2019**(9), 1 (2019),
257 doi:[10.1007/JHEP09\(2019\)020](https://doi.org/10.1007/JHEP09(2019)020).

- 258 [19] F. Ceyhan and T. Faulkner, *Recovering the qnec from the anec*, Communications in Math-
259 ematical Physics **377**, 999 (2020), doi:[10.1007/s00220-020-03751-y](https://doi.org/10.1007/s00220-020-03751-y).
- 260 [20] S. Wei-Xing, Y. Hong-Wei, L. Fei, W. Pu-Xun and R. Zhong-Zhou, *Restrictions on nega-*
261 *tive energy density for the dirac field in flat spacetime*, Chinese Physics **15**, 934 (2006),
262 doi:[10.1088/1009-1963/15/5/011](https://doi.org/10.1088/1009-1963/15/5/011).
- 263 [21] E. Curiel, *A primer on energy conditions*, Towards a theory of spacetime theories pp.
264 43–104 (2017), doi:[10.1007/978-1-4939-3210-8_3](https://doi.org/10.1007/978-1-4939-3210-8_3).