

1 **Abstract**

2

3 The high-energy cosmic rays entering the Earth's atmosphere throw light upon many
4 **different** aspects of Astroparticle **Physics** and Particle Physics. **The This work presented**
5 **in this symposium** outlines investigative learning about these **high-energy** primaries
6 based on a 7 mini array DEASA in Agra, India. DEASA (Dayalbagh Educational Air
7 Shower Array) consists of eight plastic scintillators, each with an area of 1 square meter.
8 This array covers an area of 260 square meters and is the first array in the northern part of
9 our country. 10 **The A real-life applications of the cosmic ray particles where the effect of**
11 **cosmic rays in space has been studied and the best material has been found to protect the**
12 **astronaut from the galactic cosmic rays. Poly materials were found to be the best material**
13 **due to the large amount of hydrogen (H) and low atomic number (Z).** It is observed that
14 **the** equivalent dose is minimum (107 sieverts) for Polystyrene compared to the other
15 materials. 15 Finally the high energy muons have been studied to image nuclear caskets
16 called muon tomography. In this study, a dry cask container has been simulated which
17 contains the 17 UO₂ rods and the muon scattering has been observed.

17 **MINOR COMMENT:** *The paragraph between lines 9-12 must be written clearly. It is*
18 *confusing. Also, use the active voice in places and sentences describing your own results*
19 *and findings. For example, here is a suggestion: We found poly materials are the best*
20 *material to protect astronauts from galactic cosmic rays due to a large amount of*
21 *hydrogen (H) and low atomic number (Z). Try to join the previous sentence to this one*
22 *with coherence.*

23 **It is We** observed that **the** equivalent dose is minimum (107 sieverts) for Polystyrene
24 compared to the other materials. **Finally**, the **high-energy** muons have been studied to
25 image nuclear caskets called muon tomography. **In this study, a dry cask container been**
26 **simulated which contains the UO₂ rods and the muon scattering has been observed.**

27

28 **MINOR COMMENTS:** *The marked paragraph between lines 25-26 must be written clearly.*
29 *This is an abstract. It must be concise. The paragraph starting on line 23, must be*
30 *continuous to line 22*

31

32 **MAJOR COMMENT:** *The abstract seems to be a copy-paste of the same*
33 *sentence or similar words in/from the main text. Please avoid the latter.*
34 *Change the abstract somehow. The main idea is there but is not a copy-paste*
35 *literally (e.g. compare with marked text in lines 102-104). Authors abstract*
36 *must write the abstract in a pretty concise and short way. It must have a brief*
37 *context, aim/goal/scope, results, and conclusions. No more. This work is a*
38 *peer review article, so authors must write it as it is and with more care.*

39 1 Introduction

40

41 The cosmic rays **mainly** come from radioactive decay inside the stars, **explosive**
42 **supernovae's**, **supernovae**, **the** Sun and pulsars **etc** **(do not use wordy-or etc.**
43 **Use another word like mainly, mostly or similars.)**. The higher energy ones
44 seem to be coming from **supermassive** black holes at the heart of galaxies. On
45 reaching the Earth's atmosphere, they produce showers of particles **which that**
46 pass through us almost 500 times **in** a minute. The cosmic flux is **an important a**
47 **crucial** tool for calibrating particle detectors, **and** this study is **also** being done for
48 DEASA detectors **also**. The muons entering our detectors do not have a constant
49 flux but slightly more in summer and lower in winter. This is connected to pions
50 which have decayed into muons in the shower. In **summer**, the air warms and
51 expands, leading to more gaps between air molecules allowing pions to **reach**
52 further **so as to** decay into muons. In **winter**, the air is cold and dense resulting **in**
53 higher collisions of pions leading to fewer decays into muons. The cosmic rays
54 before entering the atmosphere are **mostly primarily energetic** galactic **energetic**
55 particles coming from inside the galaxy **and more energetic** extragalactic **with**
56 **energetic particles** from the active galactic nuclei, quasars or **gamma-ray** bursts.
57 These energetic particles affect the human body in many different ways as a
58 **published study [1] says that the twins physiology, memory abilities and genes for**
59 **one of the twin's on Earth and other on ISS for 340 days. The study confirms that space-**
60 **time manipulates the genes, and affects the human immune system.** The exposed person
61 suffers from **less in** mental reasoning and memory **loss, and** studies are going on for
62 **long-term** ailments. One of the stickiest problems for NASA is how to shield **astronauts**
63 from energetic cosmic rays and solar flares. The air shower developed by an energetic
64 particle entering the atmosphere grows with depth into hadronic and electromagnetic
65 particles at **the** sea level. These muons can look into the interiors of impenetrable
66 structures in parallel to the x-ray imaging of our body. The difference **being is** that X-rays
67 have to be produced in the laboratory, **and** muons are always available. This feature
68 defines them as a good tool for **impenetrable** imaging **the impenetrable** structures like
69 pyramids **and volcano** to nuclear reactor containers. Muons **travelling** through a structure
70 will be stopped along the path or scattered depending on the thickness and density of the
71 material. The plastic scintillator lights up when a charged particle passes through so we
72 design the simulation with 46 a nuclear casket **surrounded surrounding** by two plastic
73 scintillators. **Finally, this** muon imaging technique has been used to image the interiors of
74 the nuclear reactors at **the** Fukushima Daiichi plant[2] **also**.

75 **MINOR COMMENT: The marked paragraph between lines 51-53 must be written clearly.**
76 **It is confuse, first authors are describing work [1] and then with a THE STUDY is not clear**
77 **if still is [1] or the current work. Remark and clarify what [1] is saying, and what the current**
78 **work is obtained.**

79 **2 DEASA**

80

81 The cosmic ray flux decreases rapidly with energy as $E^{-2.7}$ around 1014 49 eV. **Hence**, is
82 impossible to 50 have direct measurements. ~~At these energies, the~~ **The** secondaries
83 produced at sea level **increase** with primary particle energy **at these energies**. The change
84 in transverse momentum and scattering of the secondary particles with the atmospheric
85 particles leads to their lateral spread on **the** ground. This process of almost parallel arrival
86 of the secondary particles reaching **the** ground is called extensive air shower in which the
87 spread is between 104m² to 10km². This phenomenon gave insight into: 1. Particle Physics
88 from air shower spread. 2. The direction of secondaries arriving on **the** ground tells about
89 **high-energy** particles. 3. The cosmic ray energy spectrum. 4. Mass of primary cosmic rays.
90 DEASA is a mini array of eight plastic **scintillators**, each with an area of 1 square meter,
91 **which** has been set up as shown in Figure 1. This array covers an area of 260 square
92 meters and is the first array in the northern part of our country. The pulses from the eight
93 detectors are **being** manually studied. The pulse amplitudes, **time over the threshold**, rise
94 time, fall time, **and** full width half maximum are being observed to study correlations between
95 them. The calibration of the 12 dynode **photomultiplier** tubes attached to each of the eight
96 detectors has been completed and the flux measured is around 13500 counts per minute.
97 Daily monitoring of the detectors is maintained in **the** log book.

98

99 **3 SPACE STUDIES**

100

101 **The A real-life application of the cosmic ray particles is the effect of cosmic rays in space**
102 **to find out the best material to protect the astronaut from the solar energetic particles(SEP)**
103 **and galactic cosmic rays(GCR) [3] (COMMENT: CHECK LINES 32-38).** The water phantom
104 was irradiated with primary proton following a galactic cosmic ray energy spectrum with
105 different shielding materials. The secondary particles are created with interactions between
106 protons and the shield material in Geant4.

107

108 We found poly materials ~~were found to be~~ are the best material due to a large amount of
109 hydrogen (H) and low atomic number (Z). **(COMMENT: CHECK LINES 32-38)**. High H leads
110 to fragmentation of the heavy GCR particles into small fragments, **and** low Z produces a
111 maximum number of secondaries. Poly materials are 16 percent more effective than
112 **aluminum** in reducing the dose equivalent with only 1.5148 g/cm² **of material**. Polymeric
113 materials are expected to play an important role in protecting **the** astronauts on future
114 missions. It is observed that **the** equivalent dose is minimum (107 sievert) for Polystyrene
115 as compared to the other materials.

116

117 **4 MUON TOMOGRAPHY**

118

119 The second study defines the application of muons to identify nuclear wastage using plastic
120 scintillation detectors [4] in muon tomography. In this study, a dry cask container has been
121 simulated, **which** contains the UO₂ rods (varying in number), **and the muon** scattering has
122 been observed [4]. This shows that when the dry cask is filled with the rods, muons **are**
123 scattered to the maximum angle, **and** if the dry cask is empty, the muon will pass through it
124 straight **without** getting scattered. The scattering of energetic muons of range 3 GeV – 10
125 GeV from these containers with dimensions from Narora Nuclear plant, Uttar Pradesh(U.P.).
126 The parameters measured are energy loss, radiation length, **and** scattering angle for a
127 different number of rods gives us patterns **which that** describe the inside of the containers
128 without opening them. The radiation length is the average distance required for an electron
129 to lose 1/e of its energy and **is** measured in cm. The multiple scattering of the muons is
130 **mostly due primarily due to the** Coulomb scattering of muons in the target with charge Z,
131 calculated analytically. Muon with energy 3 GeV ~~loses~~ **loses** 3.64 MeV/c energy in
132 concrete, **and we found** the scattering angle ~~was found~~ to be 4.14 mrad **(COMMENT**
133 **define or mention what is a mrad to put in context)**, **whereas** the radiation length was
134 10.91 cm. These calculations have been done for different energy muons scattered from
135 Iron, Lead and Uranium targets in this paper.

136

137 **5 Conclusion**

138

139 These studies prove that **high-energy** quantum fields consciously assist us in applications
140 beyond the accelerating sources ~~from which~~ they arrive and reach far beyond the **human-**
141 **machine** interface. **Neutrinos**, although ~~being~~ nearly **massless**, give solutions to Dark
142 matter, Dark energy in cosmology, muons being tiny particles **that** can scan structures like
143 nuclear plants, submarines, **etc (COMMENT: Avoid to use etc. Use other words and be**
144 **more specific)**. Hadron fields have applications in medical physics, **such** as hadron therapy
145 and carbon ion therapy for ~~the~~ cancer patients. The importance of cosmic ray studies at
146 DEASA is that students can understand quantum sensors, the electronics for fast pulses,
147 **and the analysis of count rates** ~~count rates analysis~~ over the different seasons. Another
148 **important critical** aspect of these studies is the Monte Carlo simulations in Geant4 and
149 CORSIKA. These codes give a **wide-angle** view to the user and applications in space
150 physics, nuclear science, **and** many other areas.

151

152

153

154

155 **Acknowledgments**

156

157 The author acknowledges the financial support from the Director, Dayalbagh Educational
158 Institute for setting up the DEASA experiment.

159

160 **REFERENCES**

161 **The four references have different formats and using types. Use the accepted style**
162 **for the journal and uniformize**